

# **WG1 - Neutrino Oscillations**

## **Summary and Answers to Questions**



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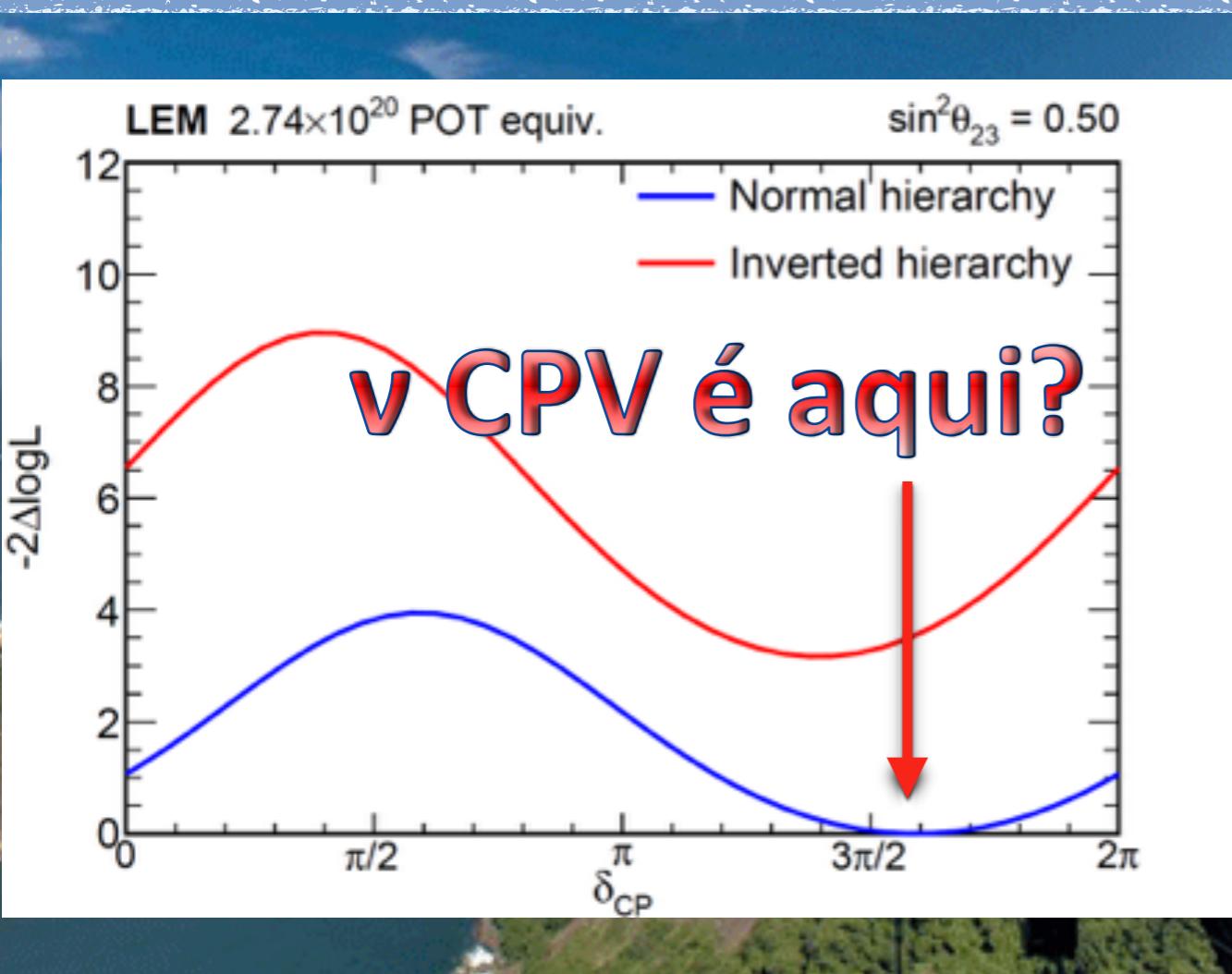
**NuFact 2015**

**CBPF, Rio de Janeiro, Brasil**

**August 15, 2015**

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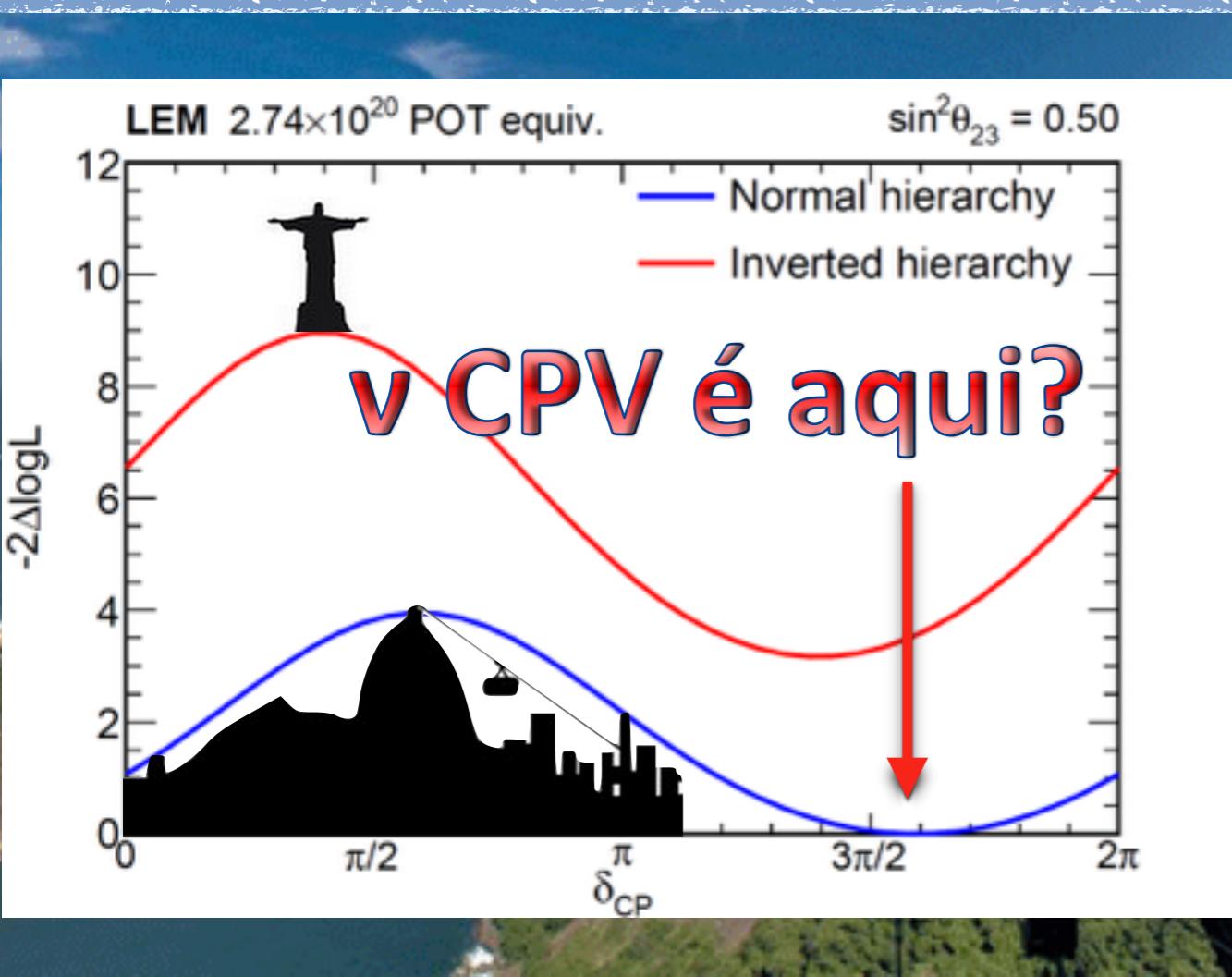
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# Q1: v Mass and Flavor Models

**Question from NuFact '14:** What are the new developments and predictions from flavor models on neutrino oscillation parameters? What precision do we need to achieve to probe them? Which parameters (or combinations of them) are more powerful to test them?

# v Mass and Flavor Models

## ► Mu-Chun Chen - Neutrino Mass and Flavor Models

- Simple "Tribimaximal" (TBM) mixing and “golden ratio” models are ruled out by non-zero  $\theta_{13}$ 
  - Try to construct models from TBM + higher order corrections using continuous or discrete family symmetries
  - Leads to sum rules connecting mixing angles and CP phase

### • Tri-bimaximal Mixing Pattern

$$U_{TBM} = \begin{pmatrix} \sqrt{2/3} & \sqrt{1/3} & 0 \\ -\sqrt{1/6} & \sqrt{1/3} & -\sqrt{1/2} \\ -\sqrt{1/6} & \sqrt{1/3} & \sqrt{1/2} \end{pmatrix}$$

### “Golden Ratio”

$$U_{GR} \equiv \begin{pmatrix} \cos \theta_{12} & \sin \theta_{12} & 0 \\ -\frac{\sin \theta_{12}}{\sqrt{2}} & \frac{\cos \theta_{12}}{\sqrt{2}} & \frac{1}{\sqrt{2}} \\ -\frac{\sin \theta_{12}}{\sqrt{2}} & \frac{\cos \theta_{12}}{\sqrt{2}} & -\frac{1}{\sqrt{2}} \end{pmatrix}$$

A<sub>5</sub> P. Ballett, S. Pascoli, J. Turner (2015)

$G_e$	$\theta_{12}$	$\theta_{23}$	$\sin \alpha_{ji}$	$\delta$
$\mathbb{Z}_3$	$35.27^\circ + 10.13^\circ r^2$	$45^\circ$	0	$90^\circ$
				$270^\circ$
$\mathbb{Z}_5$	$31.72^\circ + 8.85^\circ r^2$	$45^\circ \pm 25.04^\circ r$	0	$0^\circ$
		$45^\circ$	0	$180^\circ$
$\mathbb{Z}_2 \times \mathbb{Z}_2$	$36.00^\circ - 34.78^\circ r^2$	$31.72^\circ + 55.76^\circ r$	0	$90^\circ$
		$58.28^\circ - 55.76^\circ r$	0	$270^\circ$
				$0^\circ$
				$180^\circ$
				$0^\circ$
				$180^\circ$

Double Tetrahedral Group T': double covering of A<sub>4</sub>

$$\tan^2 \theta_\odot \simeq \tan^2 \theta_{\odot, TBM} + \frac{1}{2} \theta_c \cos \delta$$

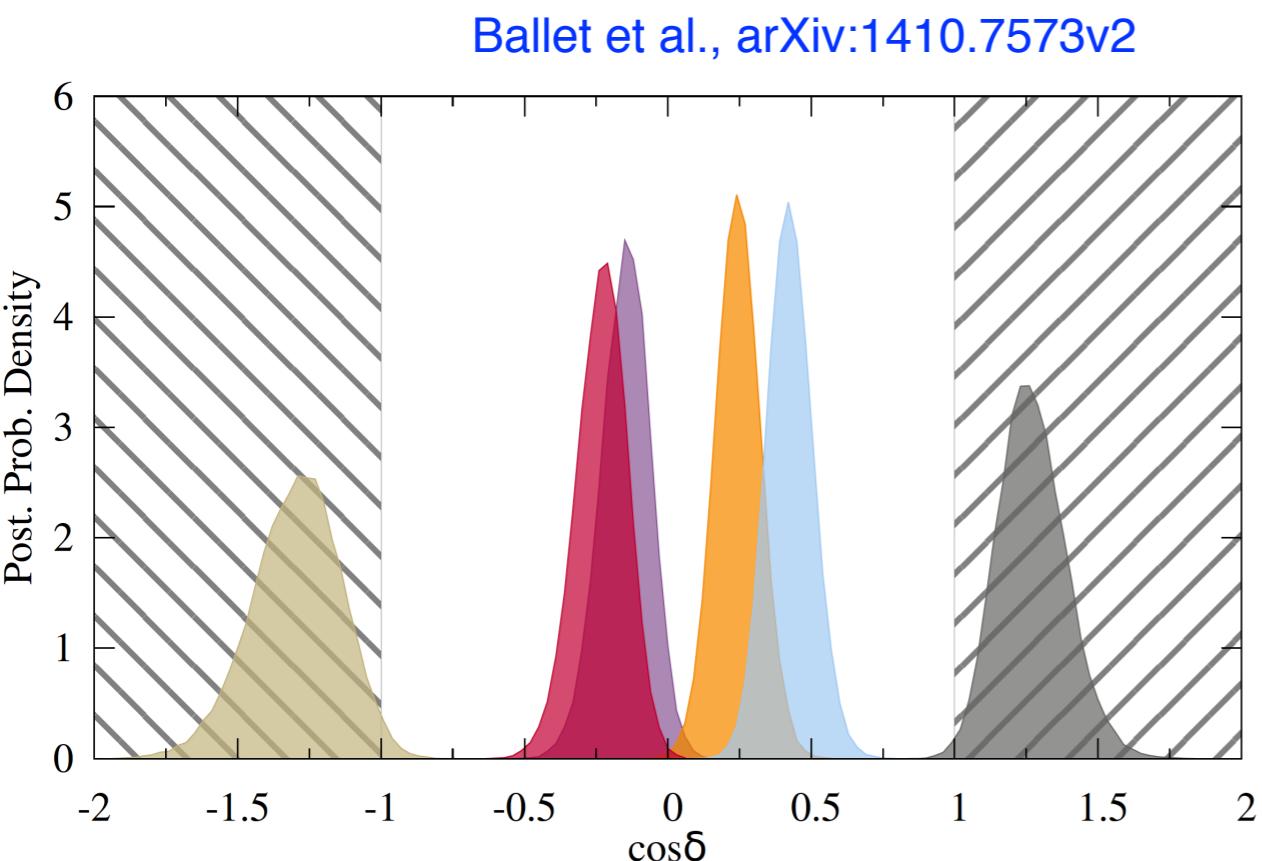
M.-C.C, K.T. Mahanthappa (2007, 2009)

TABLE I. Numerical predictions for the correlations found in this paper. The dimensionless parameter  $r \equiv \sqrt{2} \sin \theta_{13}$  is constrained by global data to lie in the interval  $0.19 \lesssim r \lesssim 0.22$  at  $3\sigma$ . The predictions for  $\theta_{12}$  and  $\theta_{23}$  shown here neglect terms of order  $\mathcal{O}(r^4)$  and  $\mathcal{O}(r^2)$ , respectively.

# $\nu$ Mass and Flavor Models

## ► Alan Bross - Panel Discussion

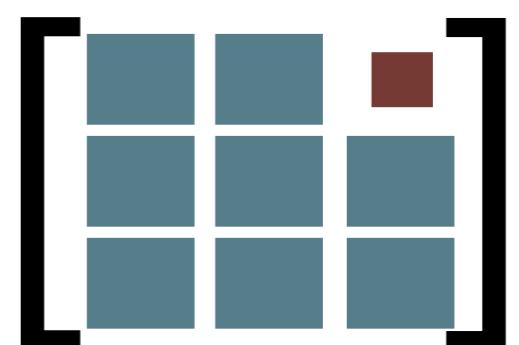
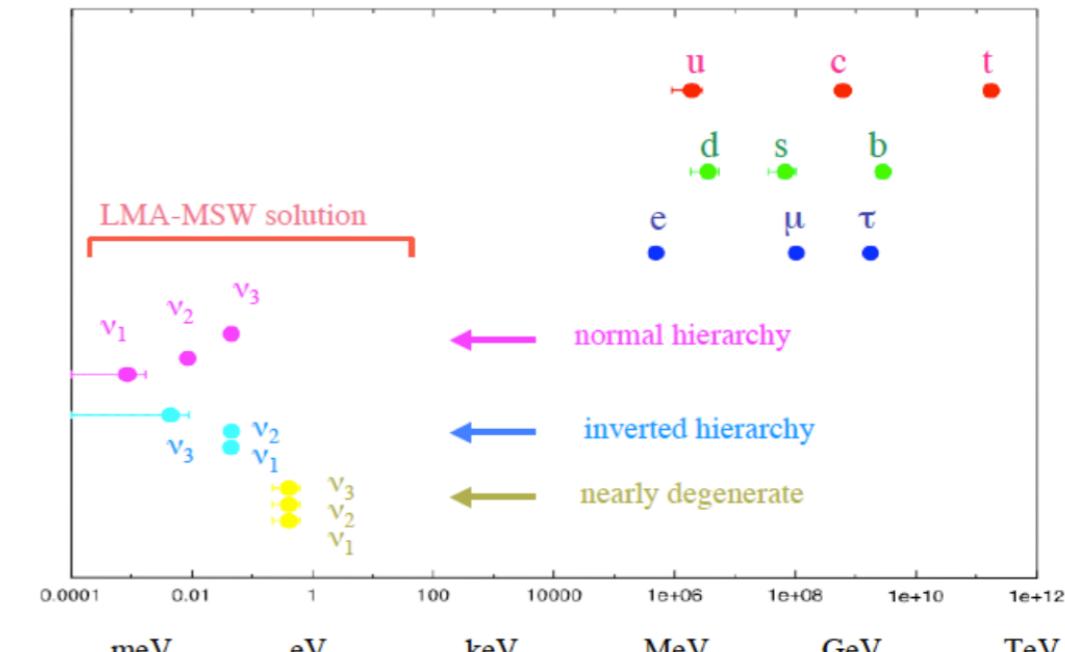
$$m_\nu \ll m_e, u, d$$



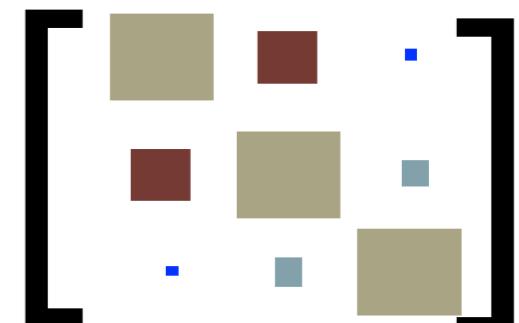
- Allowing  $3\sigma$  ranges on mixing angle measurements

$$\theta_{12} = \theta_{12}^\nu + \theta_{13} \cos \delta .$$

- More precise measurements of mixing angles and precision measurement of  $\delta$  needed!



leptonic mixing



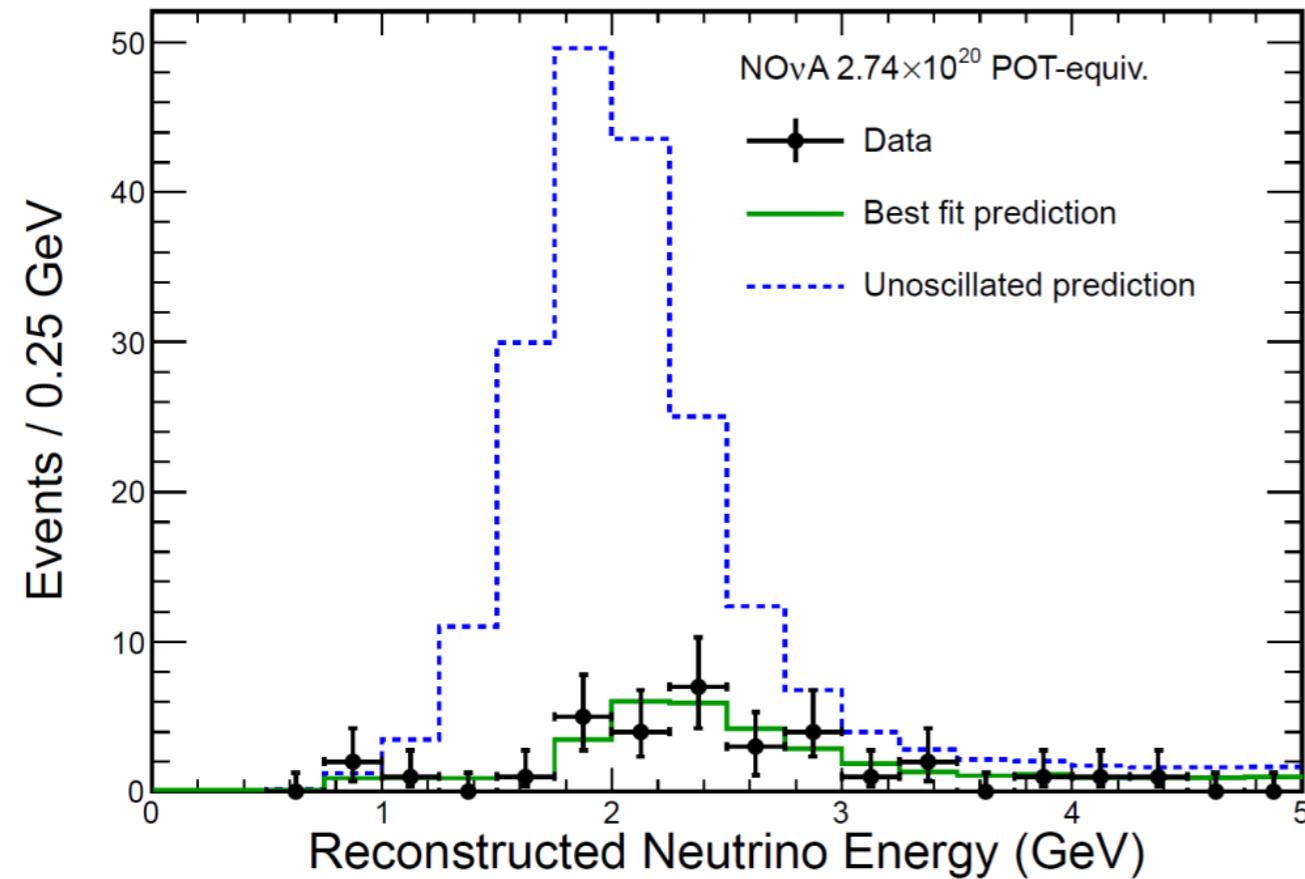
quark mixing

# New and Current Results

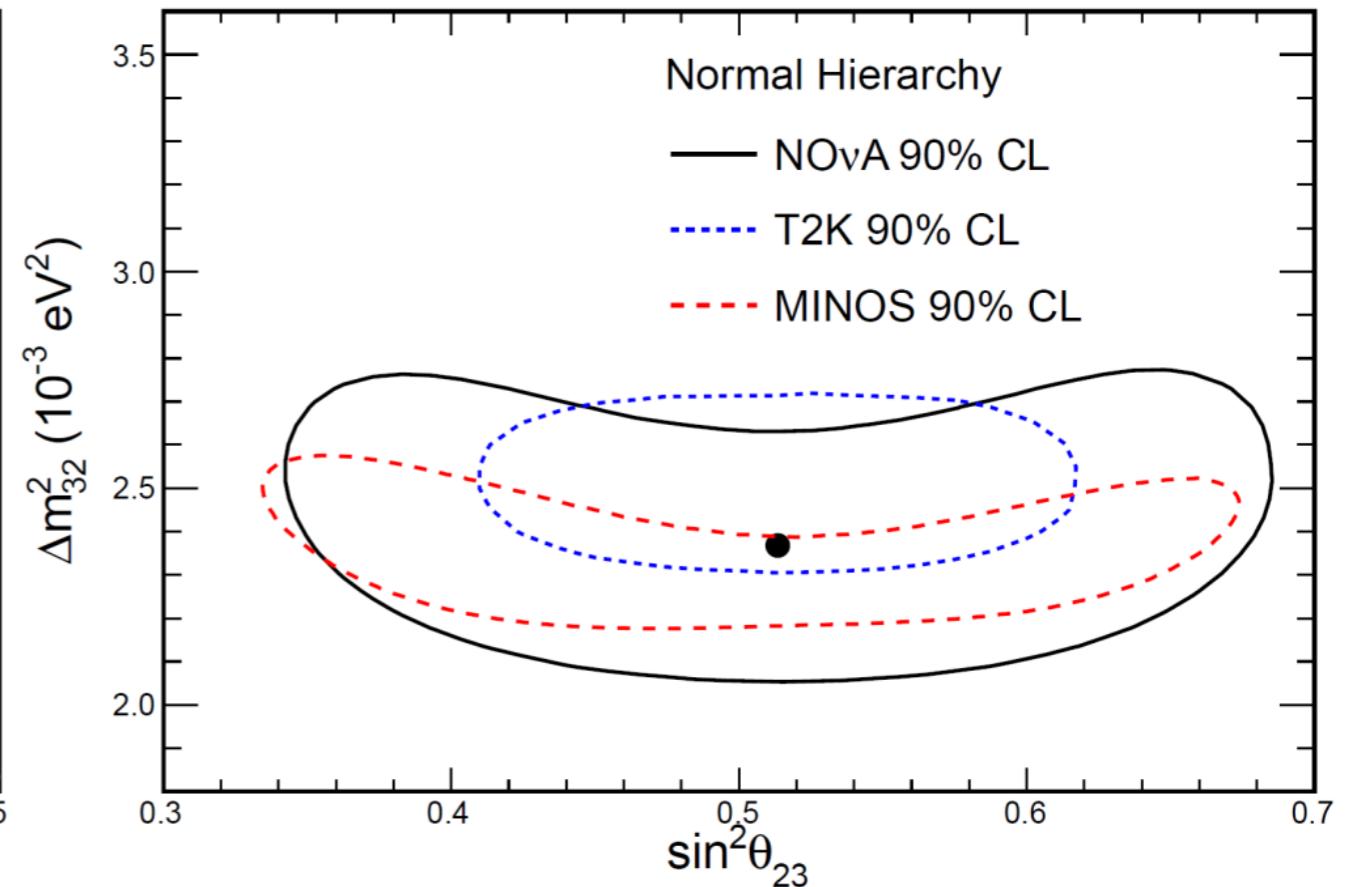
## ► NOvA - Mayly Sanchez

### Numu Disappearance

#### NOvA Preliminary



#### NOvA Preliminary



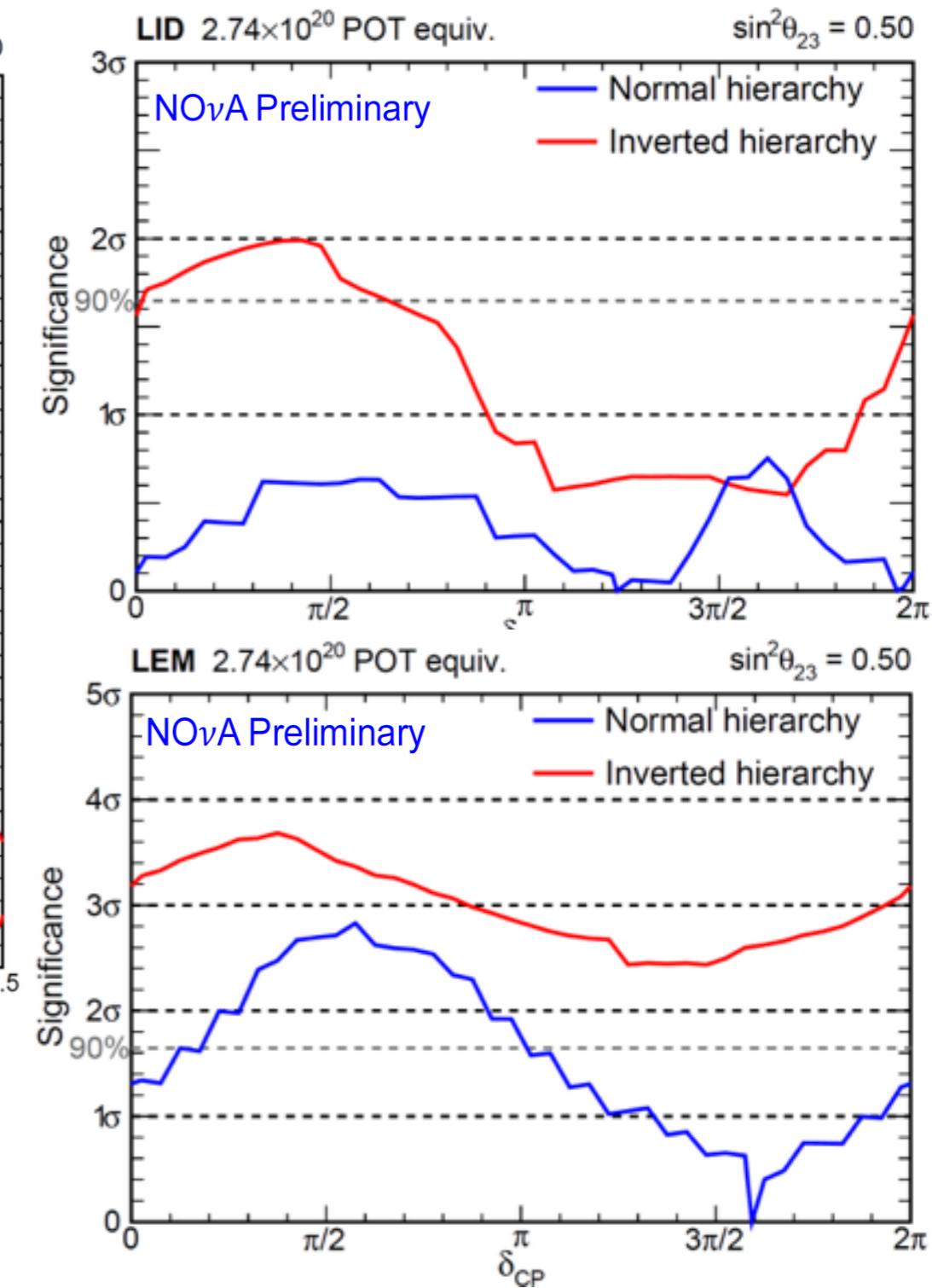
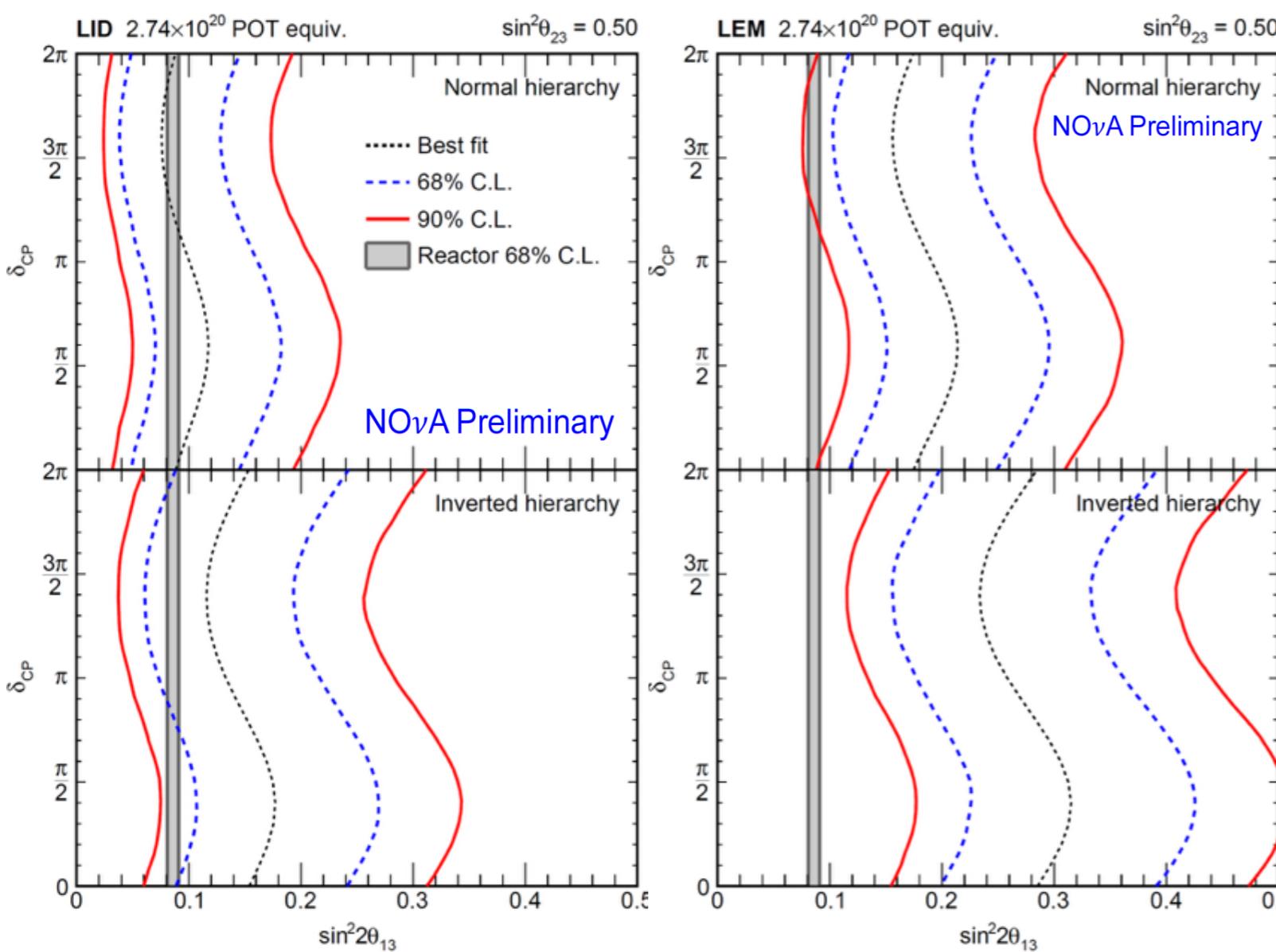
- 201  $\nu_\mu$  CC candidates predicted, 33 observed
  - Starting to be competitive with existing measurements with just 8% of nominal exposure

$$\Delta m_{32}^2 = +2.37^{+0.16}_{-0.15} \text{ [normal ordering]}$$
$$\Delta m_{32}^2 = -2.40^{+0.14}_{-0.17} \text{ [inverted ordering]}$$
$$\sin^2 \theta_{23} = 0.51 \pm 0.10$$

# New and Current Results

## ► NOvA - Mayly Sanchez

### Nue Appearance

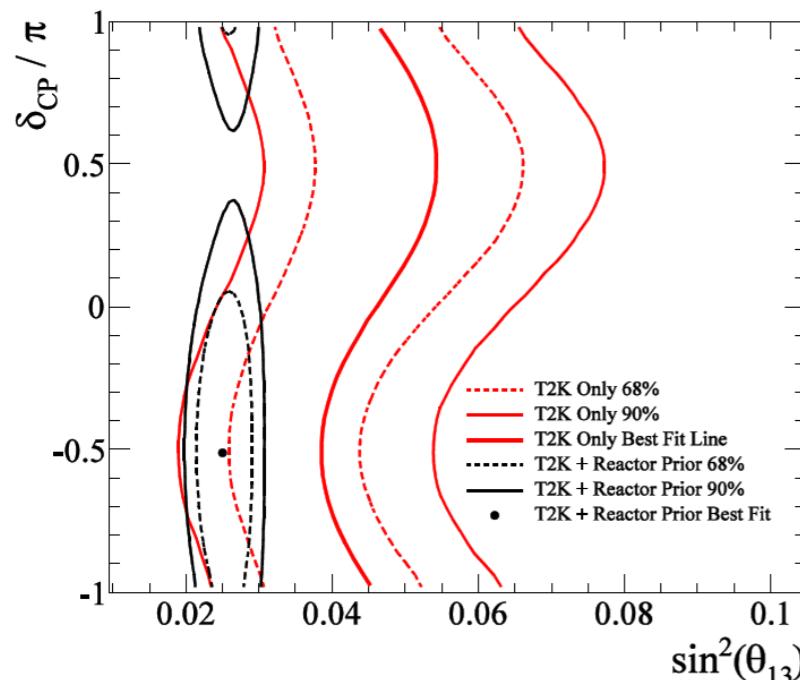


► 6 (LID) and 11 (LEM)  $\nu_e$  CC candidates found

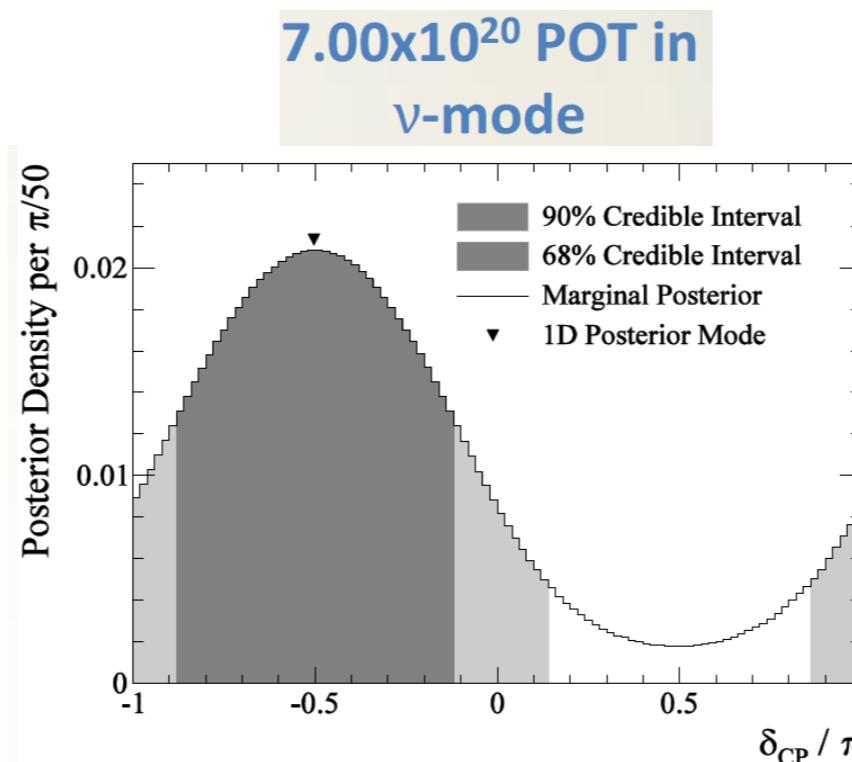
- Consistent with reactor measurements of  $\theta_{13}$
- Both selectors prefer Normal Hierarchy
- Both selectors prefer  $\delta_{CP}$  near  $3\pi/2$

# New and Current Results

## ► T2K - Kirsty Duffy



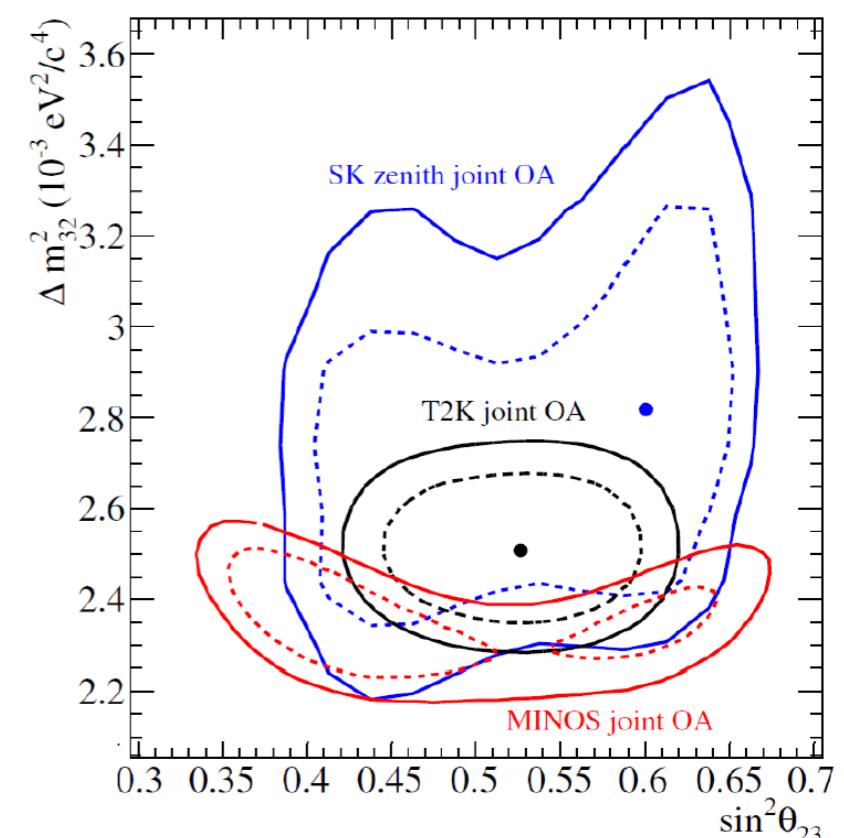
First measurement of  $\nu_e$  appearance ( $7.3\sigma$ ).  
Independent measurement of  $\theta_{13}$  (analyses performed with and without reactor constraint on  $\theta_{13}$ ,  $\sin^2 2\bar{\theta}_{13} = 0.095 \pm 0.01$ )



90% constraint on  $\delta_{CP}$

- Preference for normal hierarchy
- Preference for  $\delta_{CP} = -\pi/2$

## Current Results



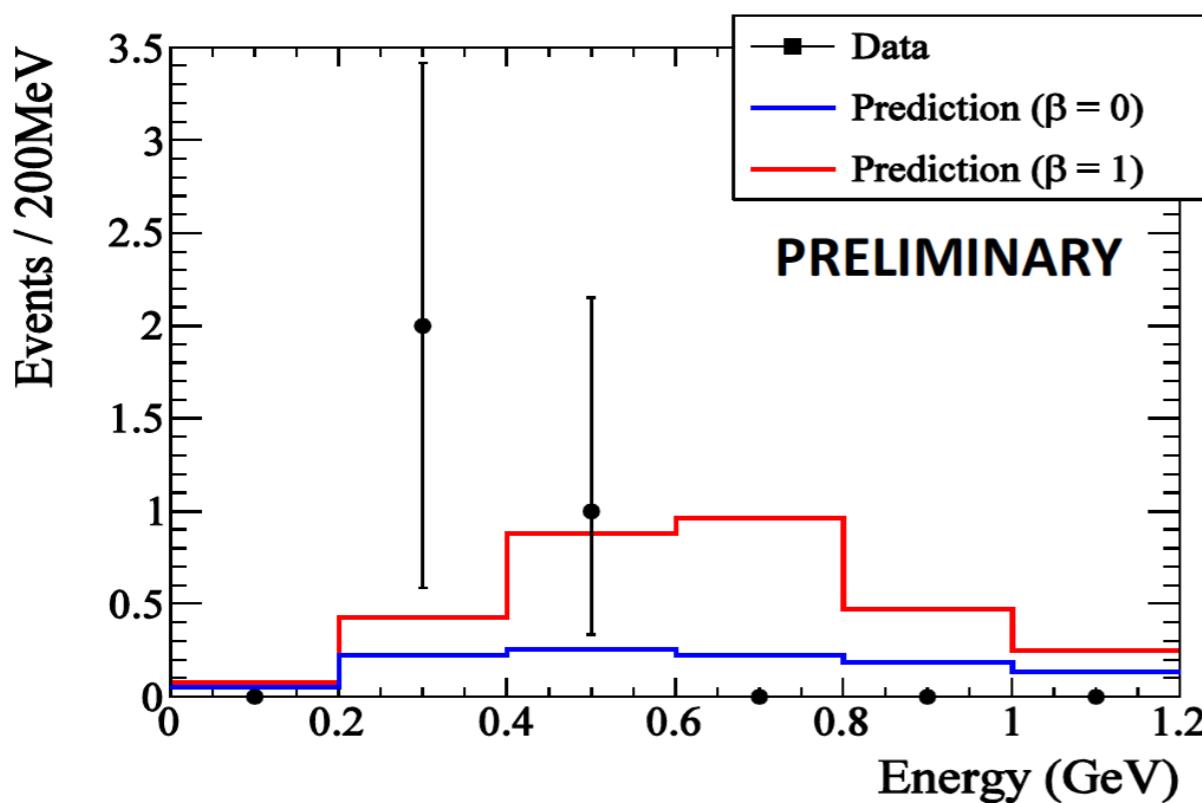
World-leading measurement of  $\theta_{23}$ .  
Significant measurement of  $\Delta m^2_{32}$ .

Abe, K. et al, Physical Review D 91.7 (2015): 072010

# New and Current Results

## ► T2K - Kirsty Duffy

### Anti-Nue Appearance



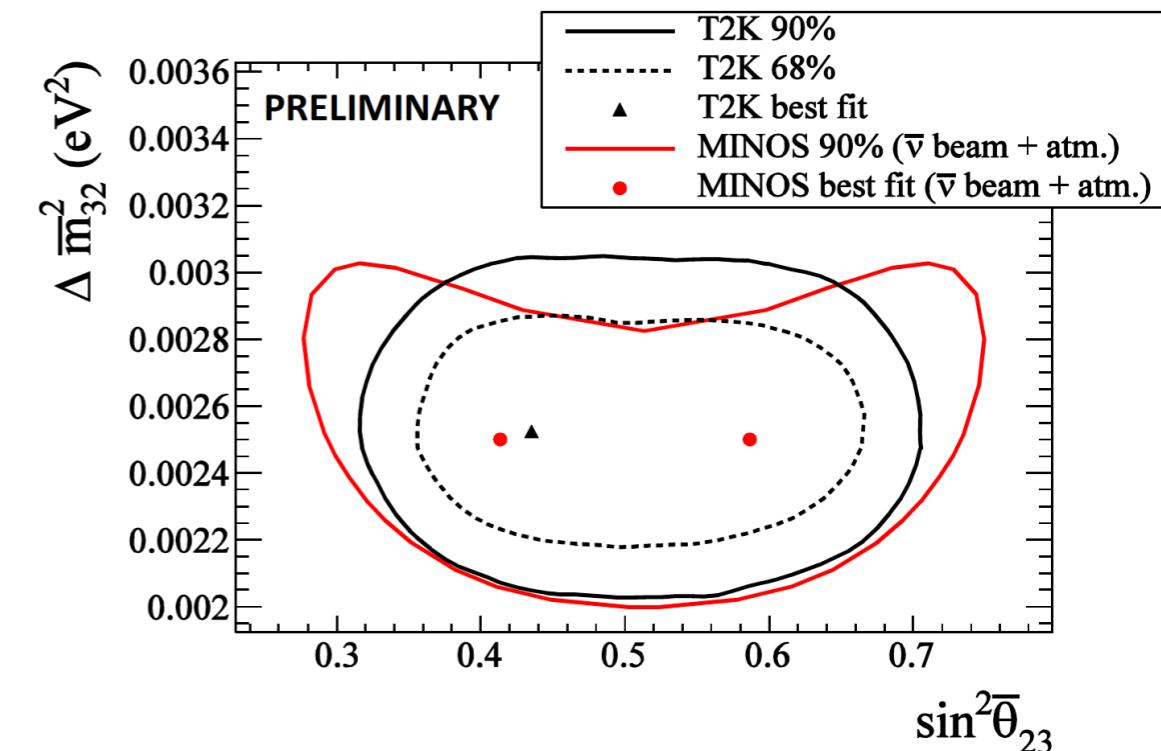
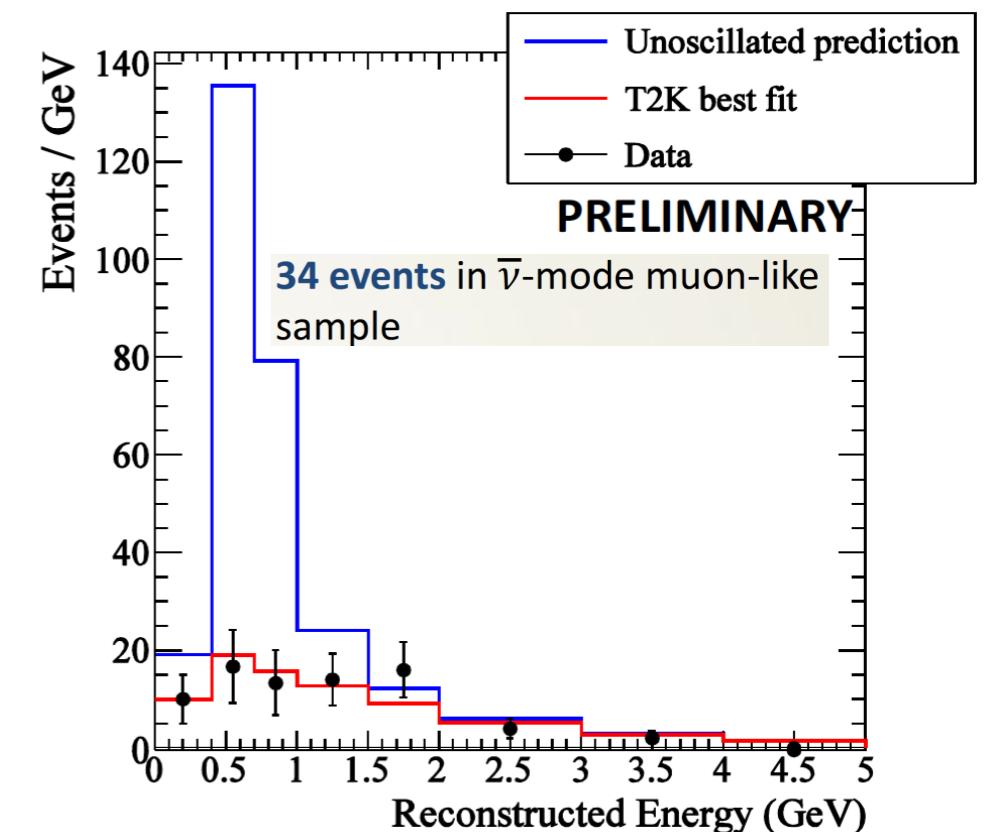
$4.04 \times 10^{20}$  POT in  
 $\bar{\nu}$ -mode

- The current data set contains 3 events
- Prediction (using T2K  $\nu$ -mode oscillation parameters) is **3.7 events** under  $\beta = 1$  and **1.3 events** under  $\beta = 0$

$$P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) = \beta \times P_{PMNS}(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$$

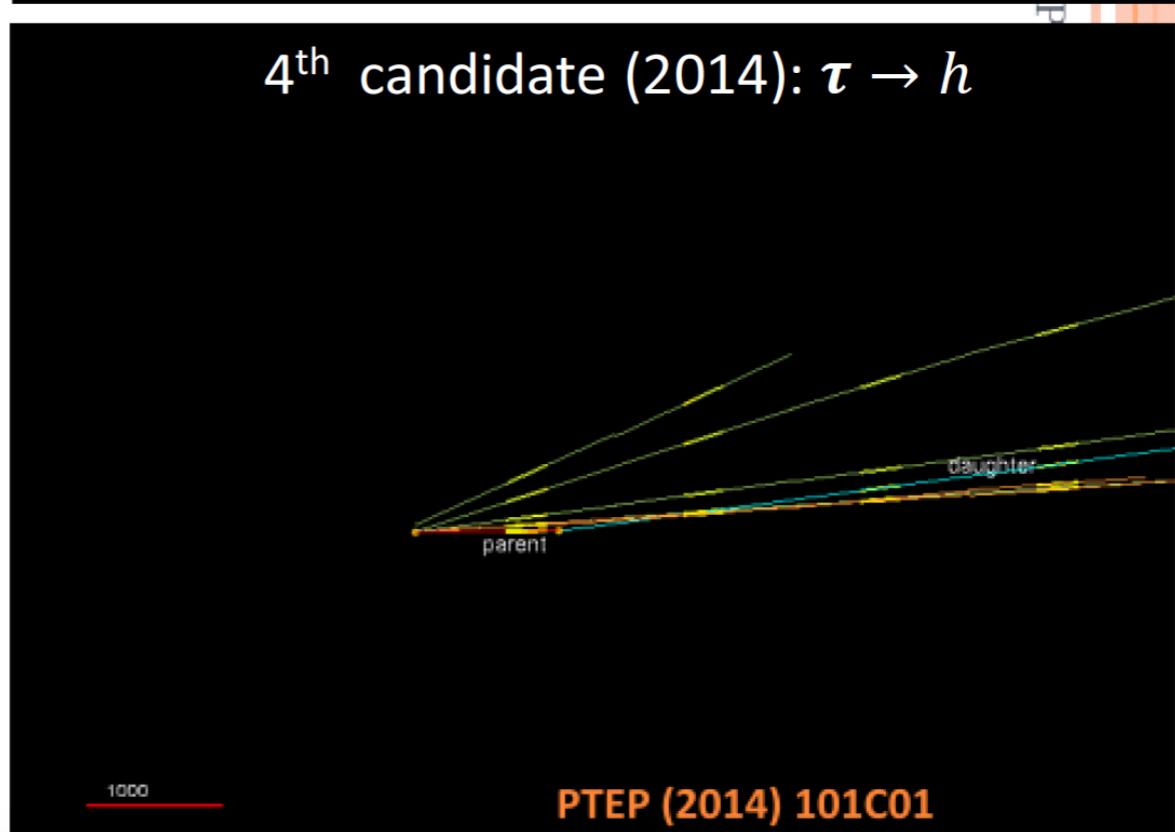
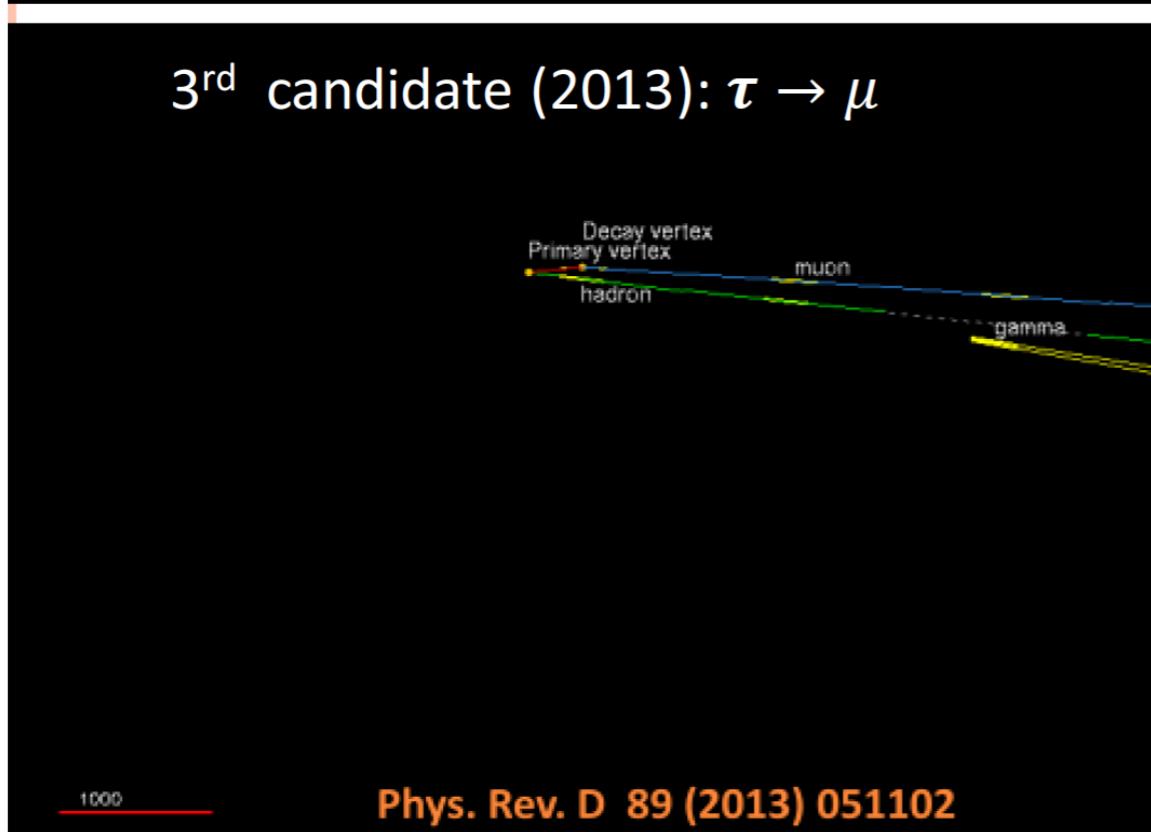
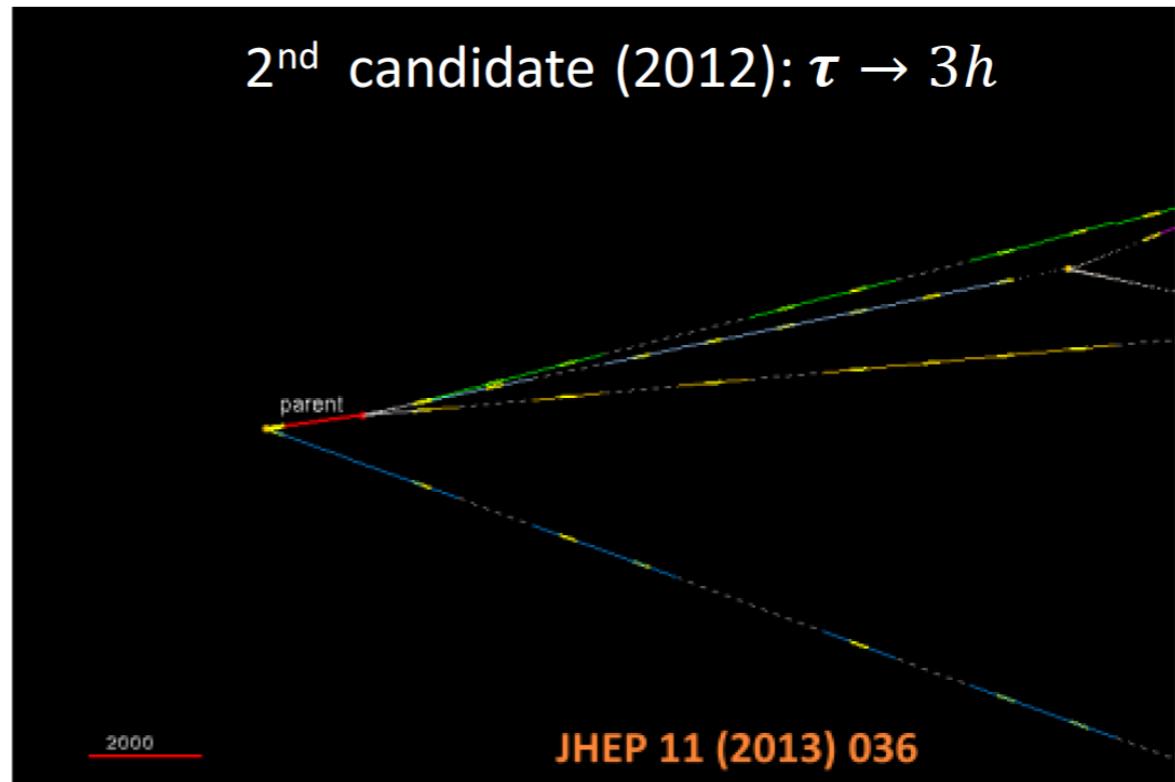
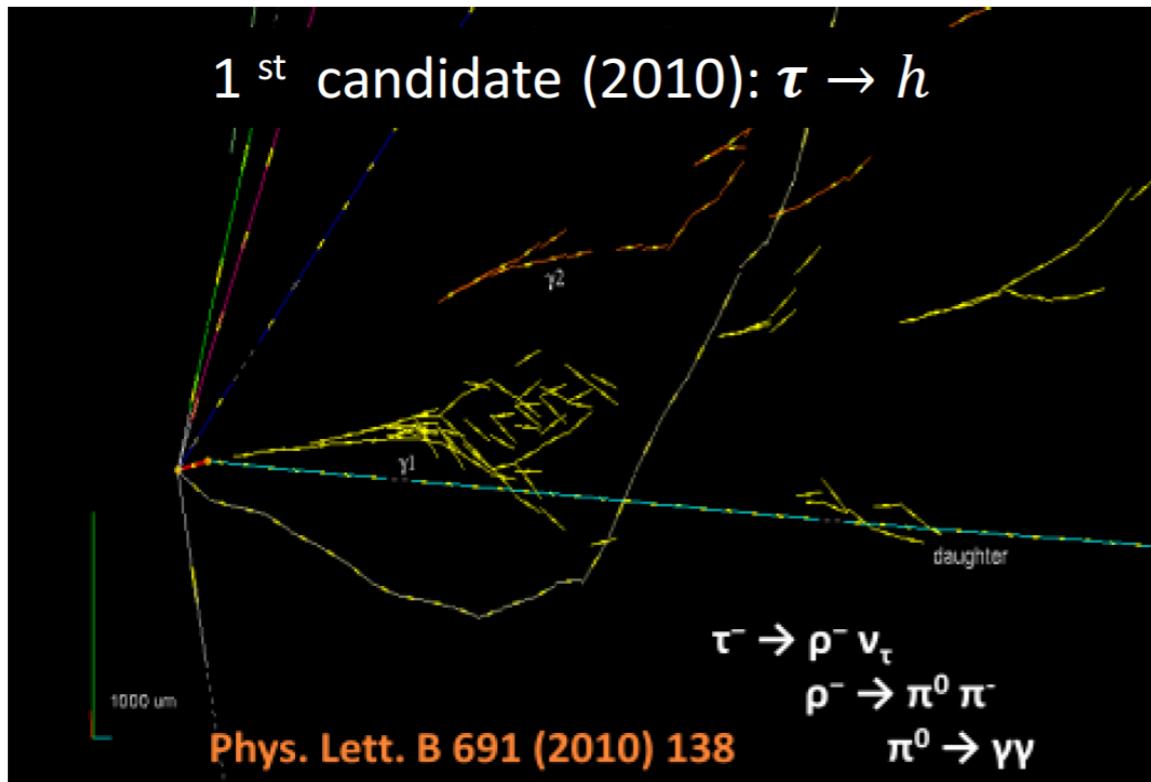
- Current data set does not provide sufficient evidence to support  $\beta=1$  over  $\beta=0$
- Consistent with NH,  $\delta = -\pi/2$  (with large statistical uncertainty)

### Anti-Numu Disappearance



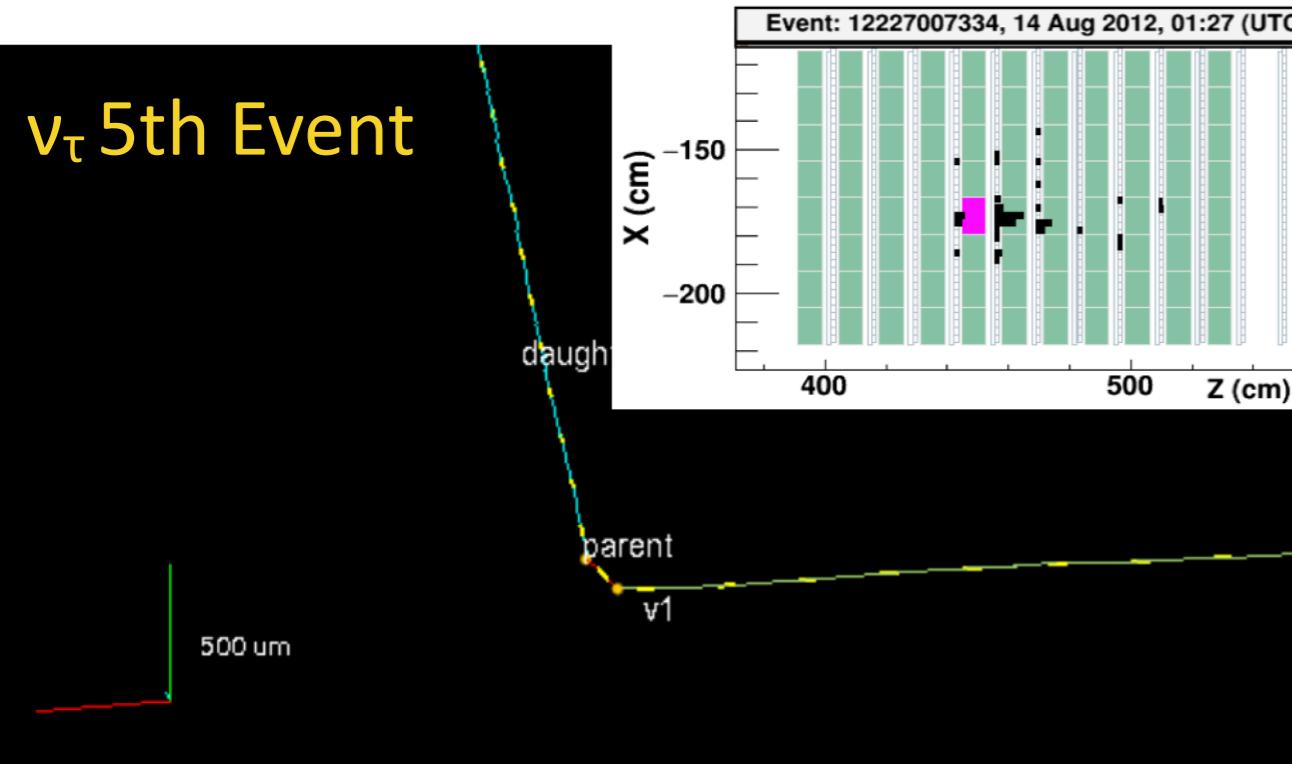
# New and Current Results

## ► OPERA - Chiara Sirignano



# New and Current Results

## ► OPERA - Chiara Sirignano

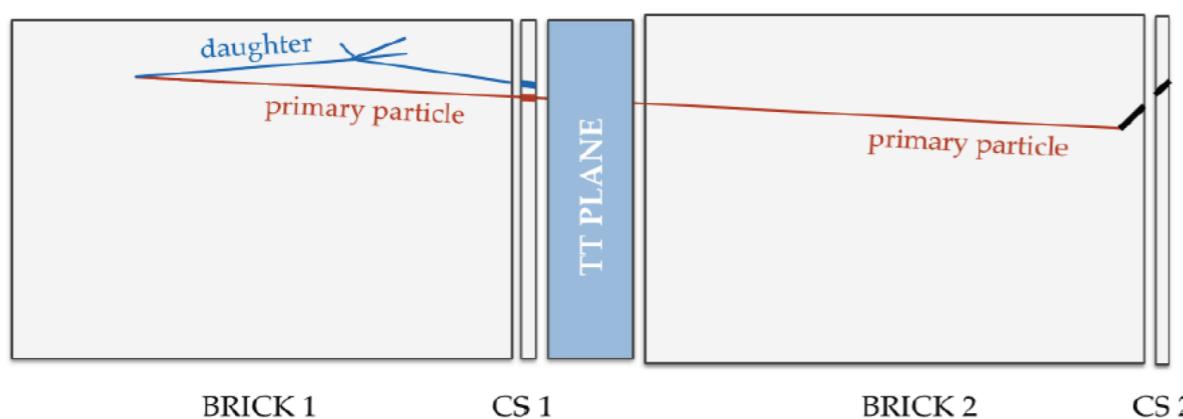


Channel	Expected background	Expected signal	Observed
$\tau \rightarrow 1h$	$0.04 \pm 0.01$	$0.52 \pm 0.10$	3
$\tau \rightarrow 3h$	$0.17 \pm 0.03$	$0.73 \pm 0.14$	1
$\tau \rightarrow \mu$	$0.004 \pm 0.001$	$0.61 \pm 0.12$	1
$\tau \rightarrow e$	$0.03 \pm 0.01$	$0.78 \pm 0.16$	0
Total	$0.25 \pm 0.05$	$2.64 \pm 0.53$	5

$$P\text{-value} = 1.1 \cdot 10^{-7}$$

Exclusion of background-only hypothesis:  $5.1 \sigma$

→ discovery of  $\nu_\tau$  appearance in the CNGS beam



**Primary particle**

Followed in the downstream brick  
Hadronic re-interaction: 1 visible particle

**Daughter**

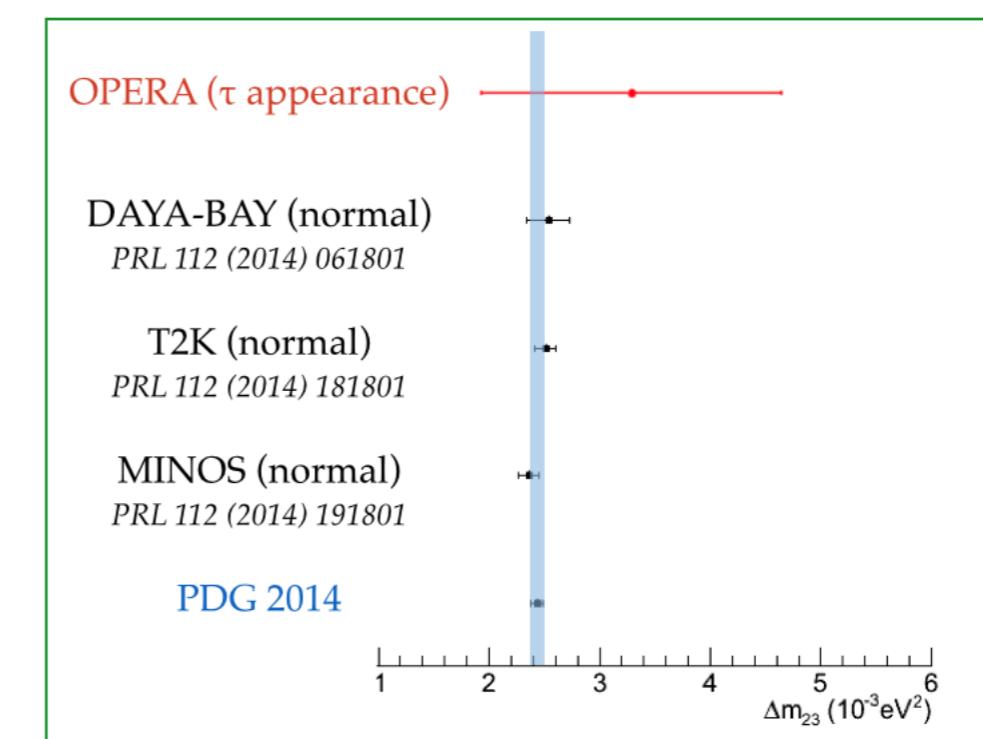
Hadronic re-interaction in the first brick



**Charm hypothesis discarded**



**Hadronic decay channel**



$$\Delta m_{23}^2 = [2.0 - 4.7] \cdot 10^{-3} \text{ eV}^2 \\ (\text{assuming full mixing})$$

# New and Current Results

## ► MINOS/MINOS+ - João Coelho

- MINOS+ has accumulated  $6.4 \times 10^{20}$  POT NuMI beam data. Focusing on exotic searches.
- Combined MINOS/MINOS+ fit consistent with reported MINOS results

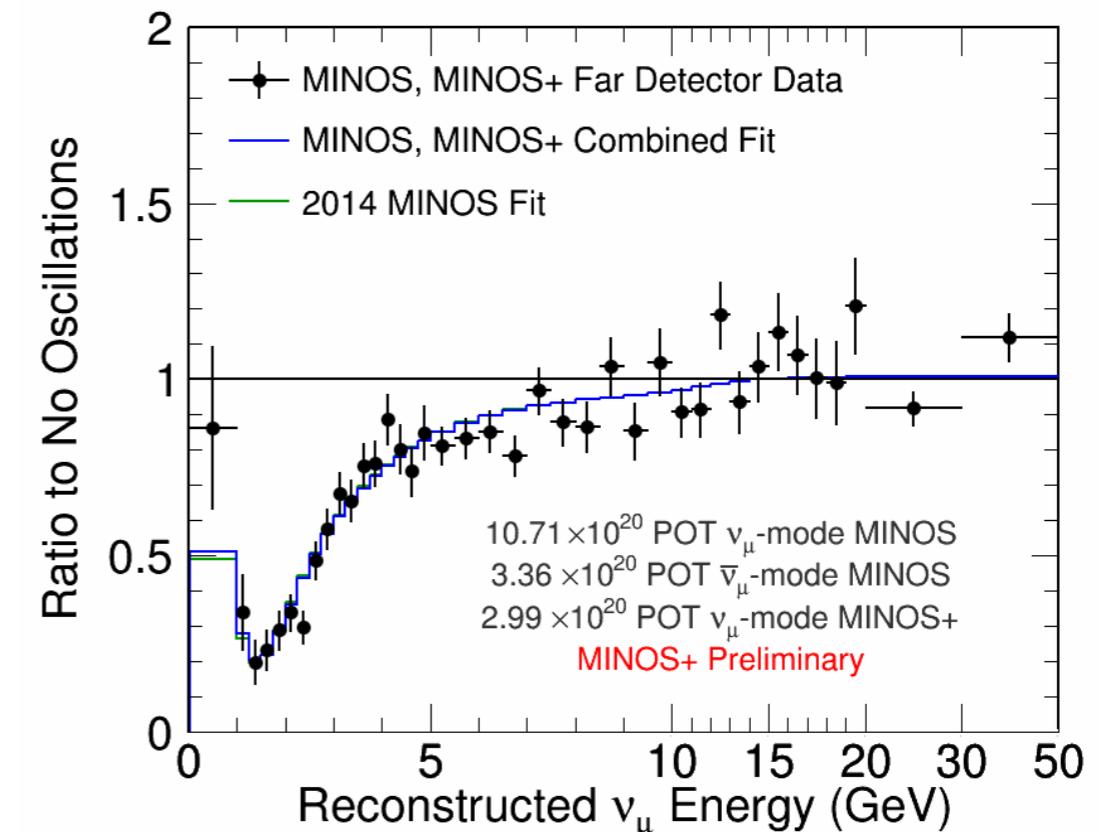
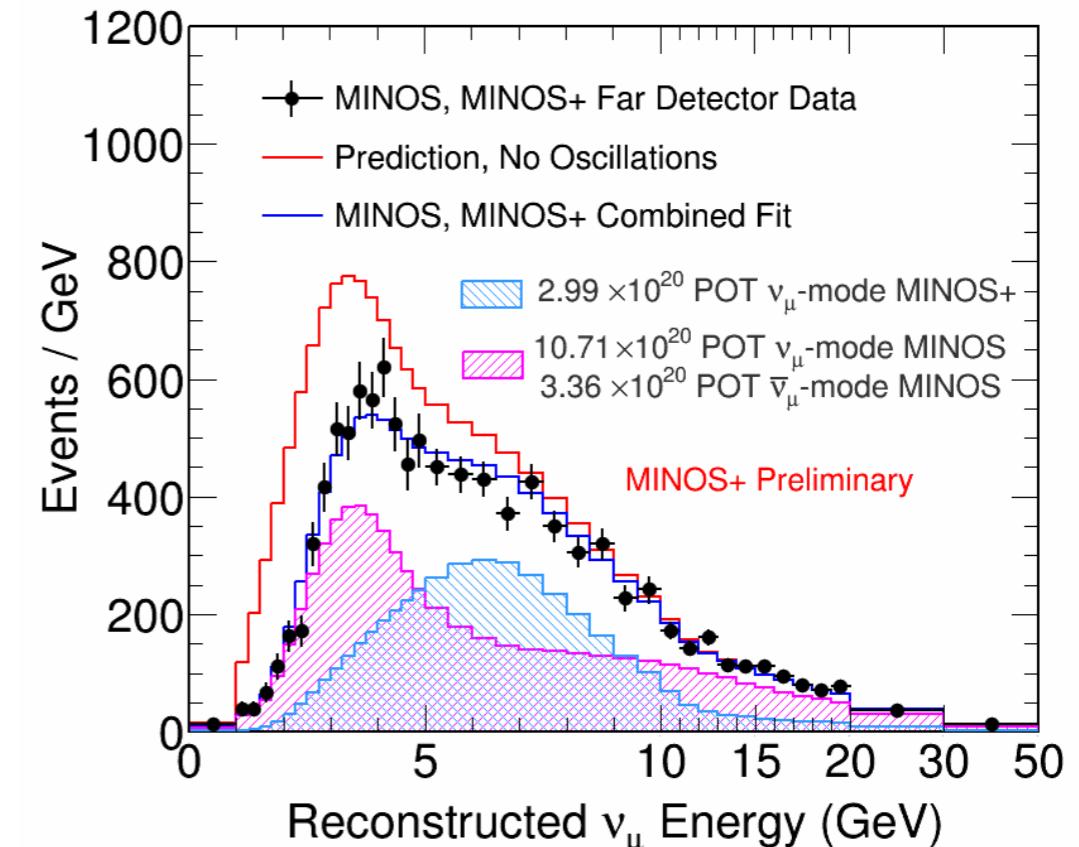
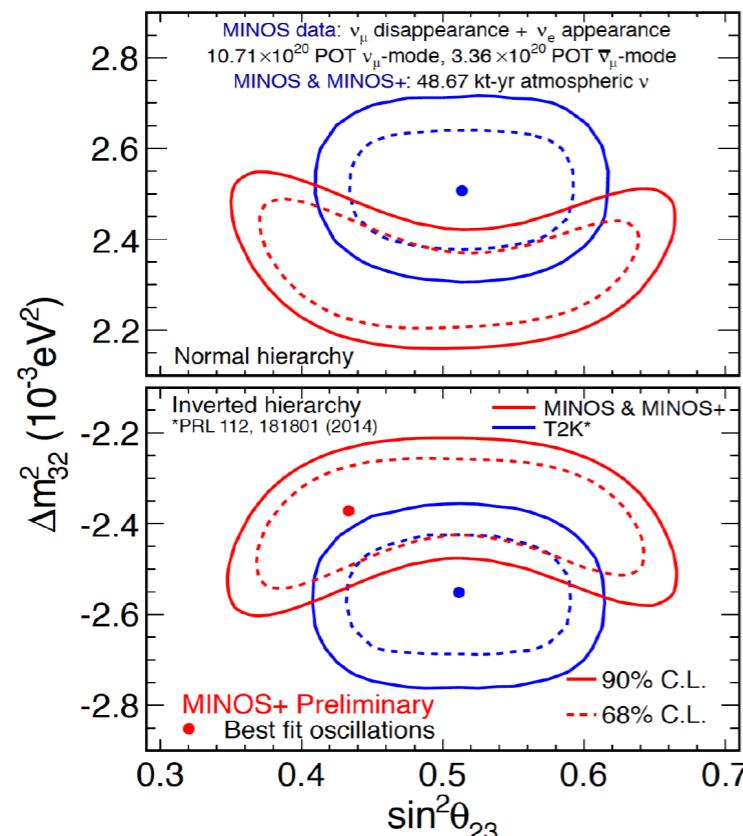
### Three-Flavor Oscillations Best Fit

Inverted Hierarchy

$$|\Delta m_{32}^2| = 2.37^{+0.11}_{-0.07} \times 10^{-3} \text{ eV}^2$$

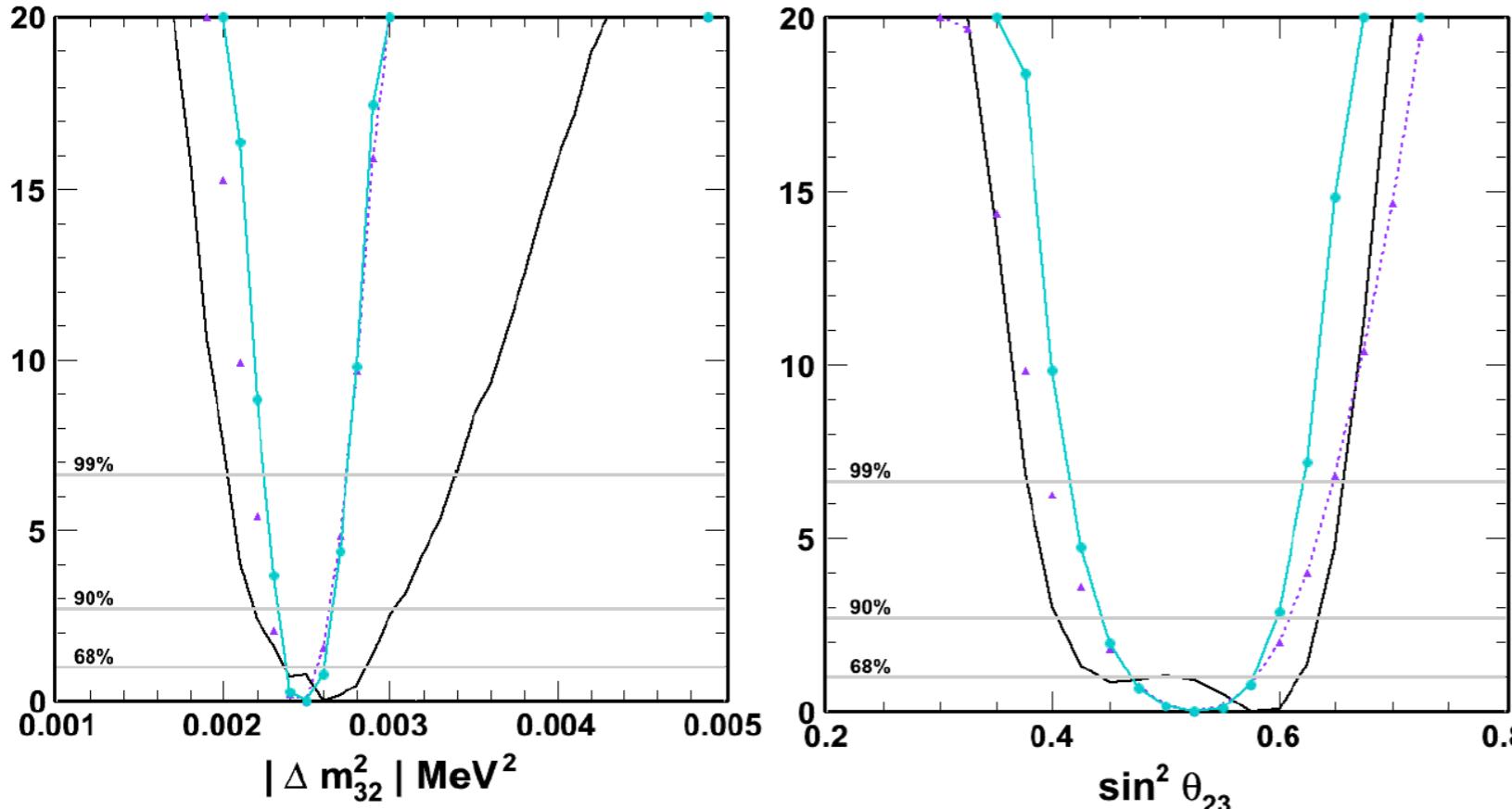
$$\sin^2 \theta_{23} = 0.43^{+0.19}_{-0.05}$$

$$0.36 < \sin^2 \theta_{23} < 0.65 \text{ (90% C.L.)}$$



# New and Current Results

## ► Super-Kamiokande - Jun Kameda



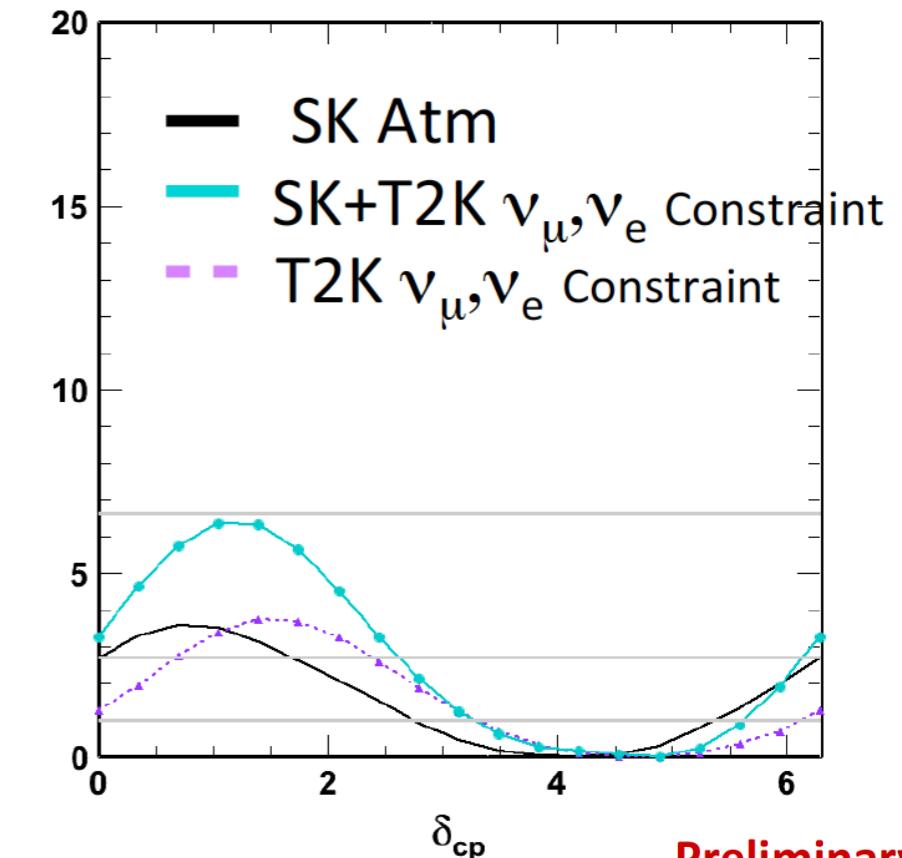
Preliminary

- Normal hierarchy favored at:  $\chi^2_{\text{NH}} - \chi^2_{\text{IH}} = -3.2$  (-3.0 SK only)
- CP Conservation ( $\sin\delta_{\text{cp}} = 0$ ) allowed at (at least) 90% C.L. for both hierarchies

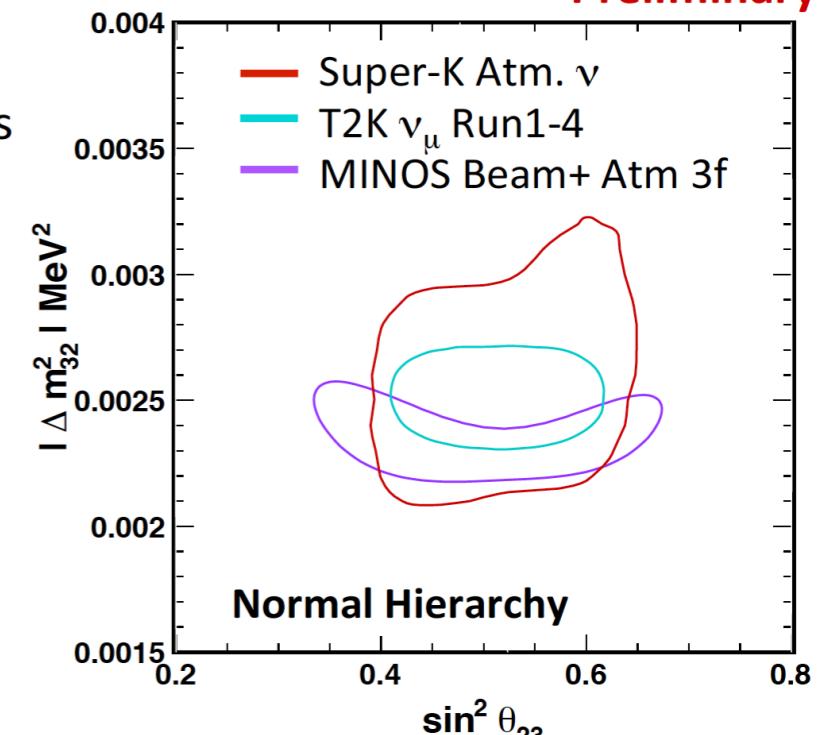
new result with 4972 days SK data

SK-I (1996-2001) SK-II (2003-2005)

SK-III(2005-2008) SK-IV(2008-Present) → total 4972days



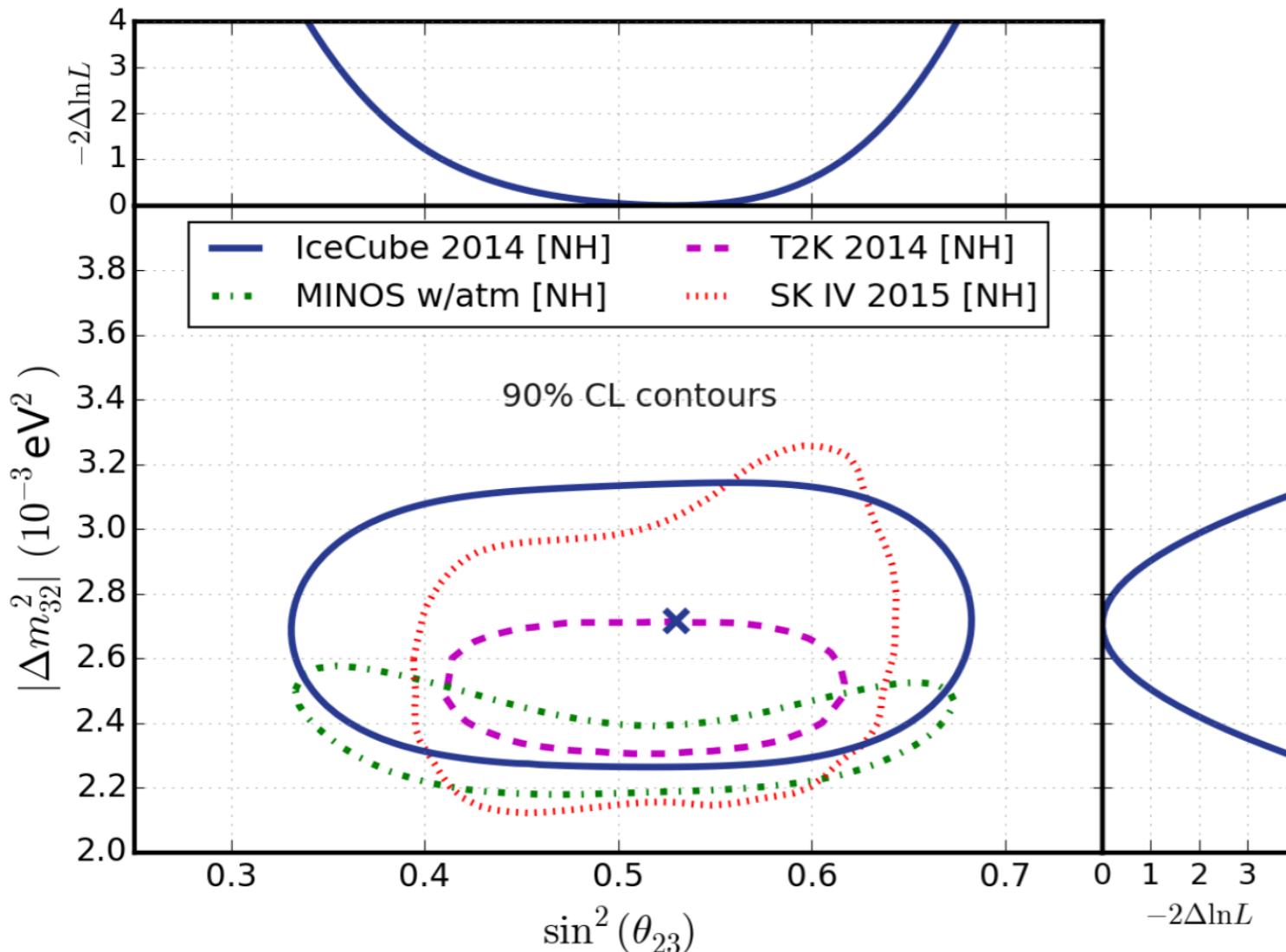
Preliminary



# New and Current Results

## ► IceCube/DeepCore - JP Athayde Marcondes de André

PRD 91, 072004 (2015) with SK result updated



$$|\Delta m_{32}^2| = 2.72^{+0.19}_{-0.20} 10^{-3} \text{ eV}^2$$

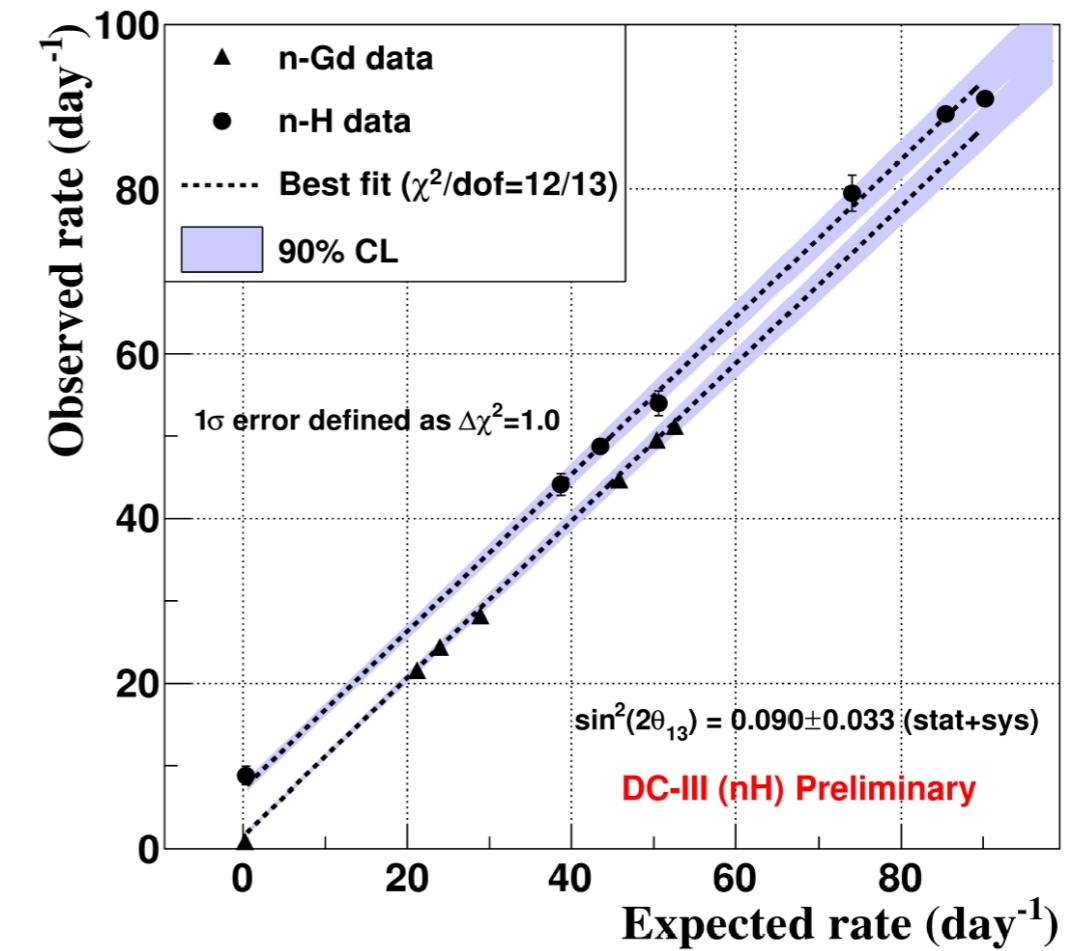
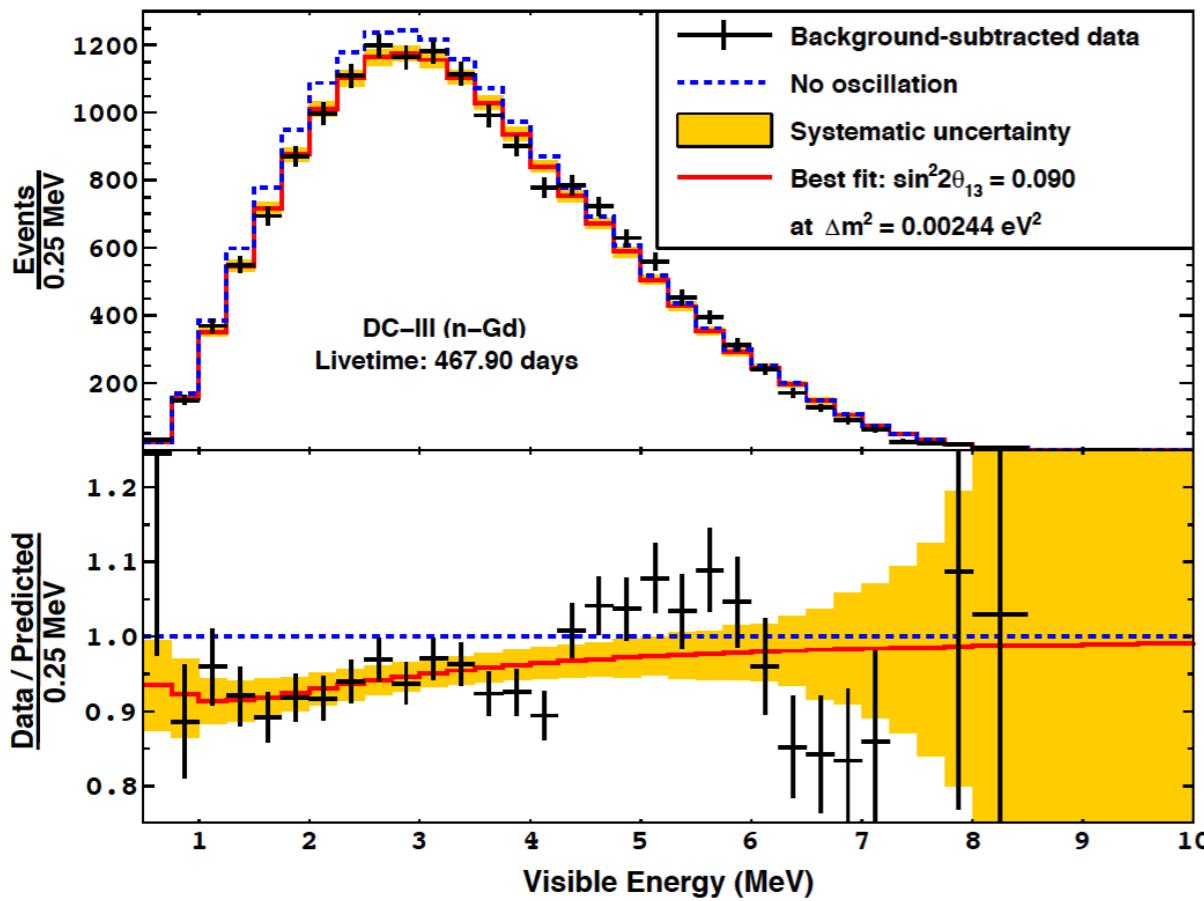
$$\sin^2(\theta_{23}) = 0.53^{+0.09}_{-0.12}$$

This measurement is still statistics limited!

- Still working on strategy ②: relax “golden event” requirements  $\Rightarrow$  expected increase by an order of magnitude of number  $\nu$  in sample

# New and Current Results

## ► Double Chooz - Guillaume Pronost

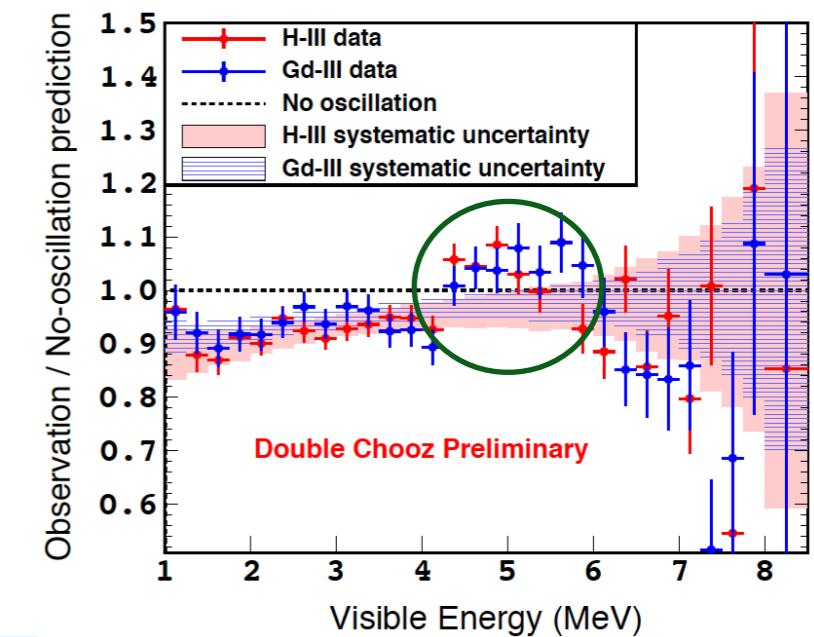


- New improved H-n analysis combined with 2014 Gd-n analysis improves statistical significance

$$\sin^2 2\theta_{13} = 0.090 \pm 0.033$$

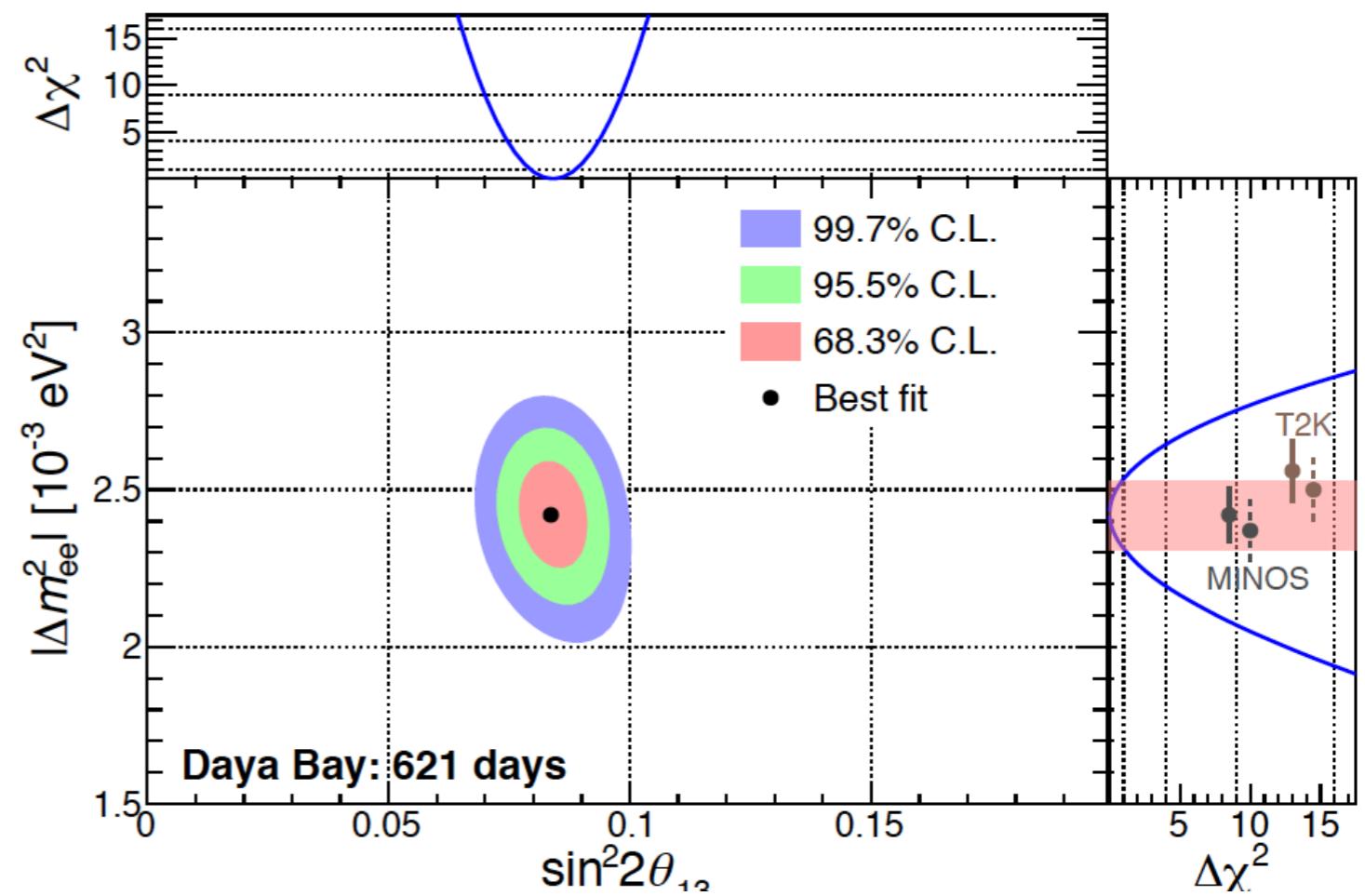
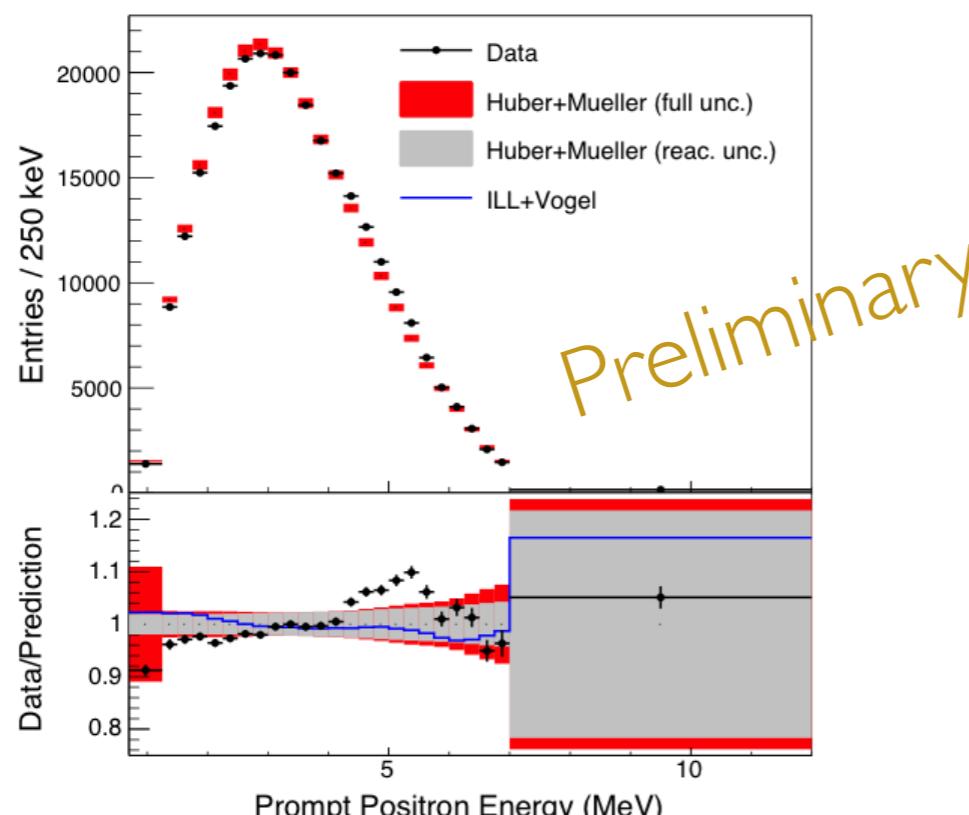
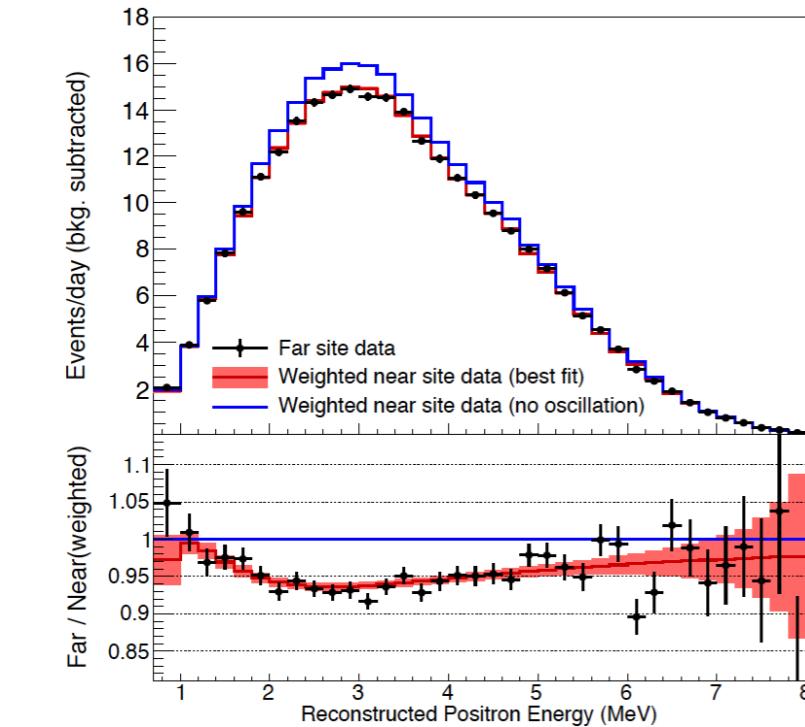
H only:  $\sin^2 2\theta_{13} = 0.098^{+0.038}_{-0.039}$ , Gd only:  $\sin^2 2\theta_{13} = 0.090^{+0.034}_{-0.035}$

- Working on understanding 5 MeV bump
- 6 months data taking with ND, Two-detector analysis in progress



# New and Current Results

## ► Daya Bay - Wei Wang



$$\left\{ \begin{array}{l} \sin^2 2\theta_{13} = 0.084 \pm 0.005 \\ |\Delta m_{ee}^2| = (2.42 \pm 0.11) \times 10^{-3} \text{ eV}^2 \end{array} \right.$$

to appear on **Phys. Rev. Lett.**, arXiv:1505.03456

- Most precise measurement of  $\theta_{13}$  and  $\Delta m^2_{ee}$
- 5 MeV bump also seen by Daya Bay

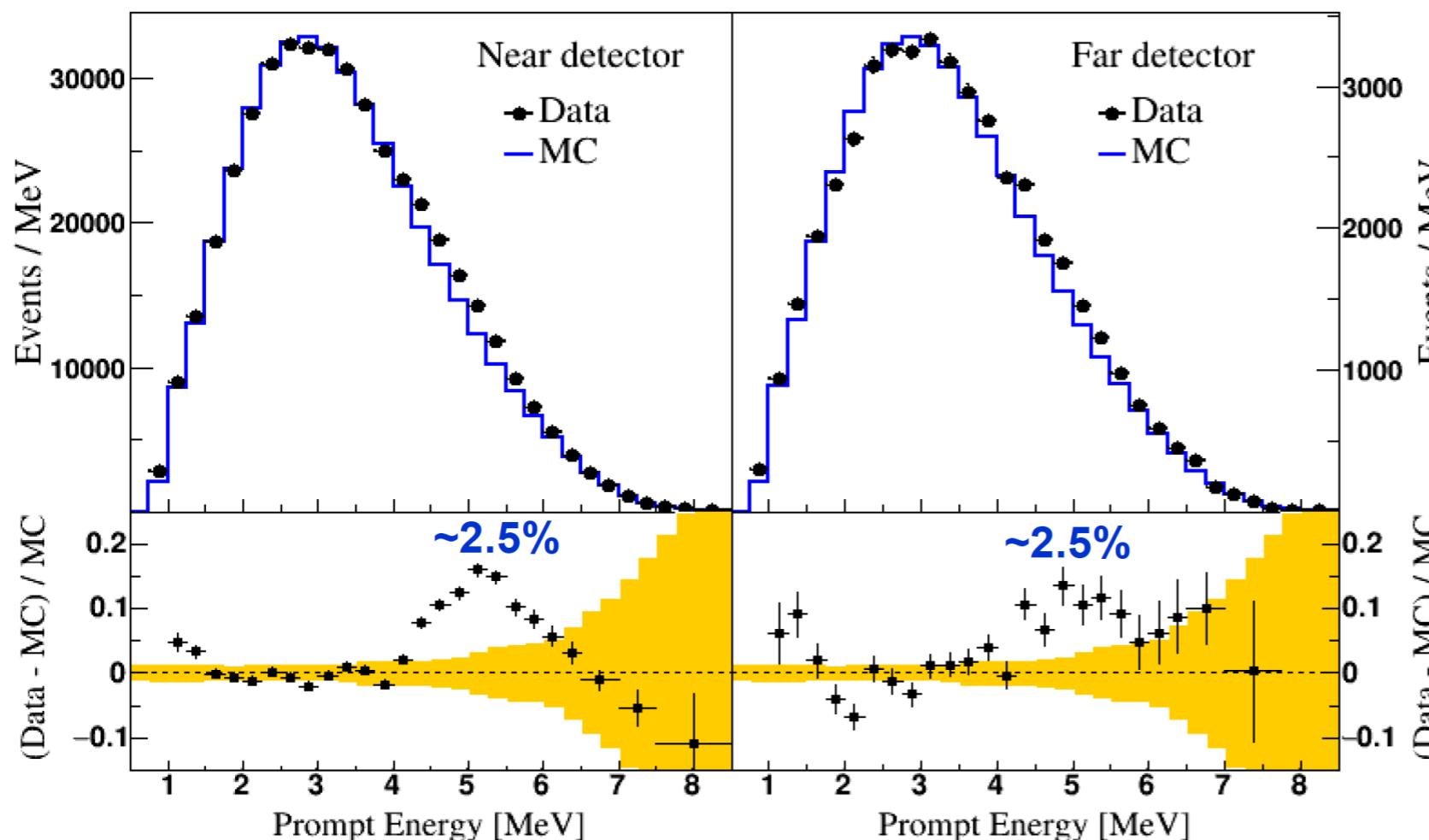
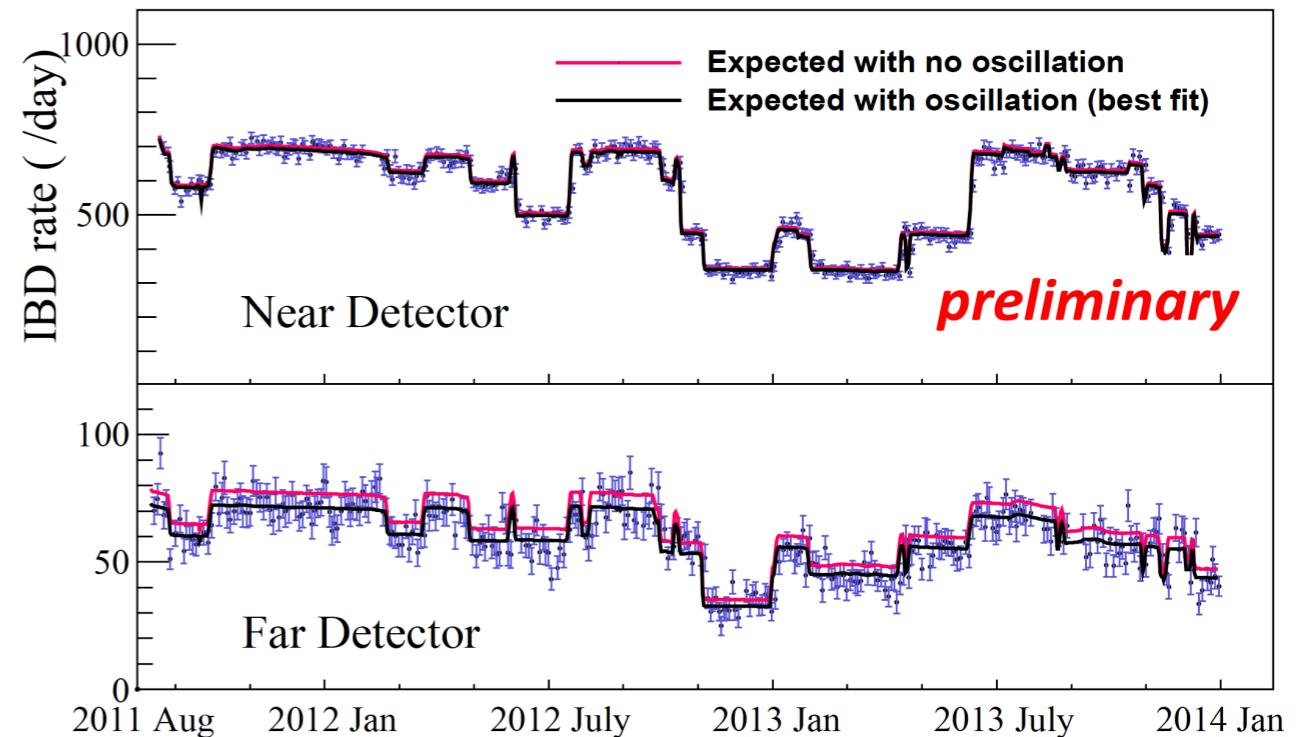
# New and Current Results

## ► RENO - Kyungkwang Joo

- Analyzed 800 days of data
- New rate-only measurement of  $\theta_{13}$

$$\sin^2 2\theta_{13} = 0.087 \pm 0.008(\text{stat}) \pm 0.008(\text{syst})$$

- Shape analysis coming soon!

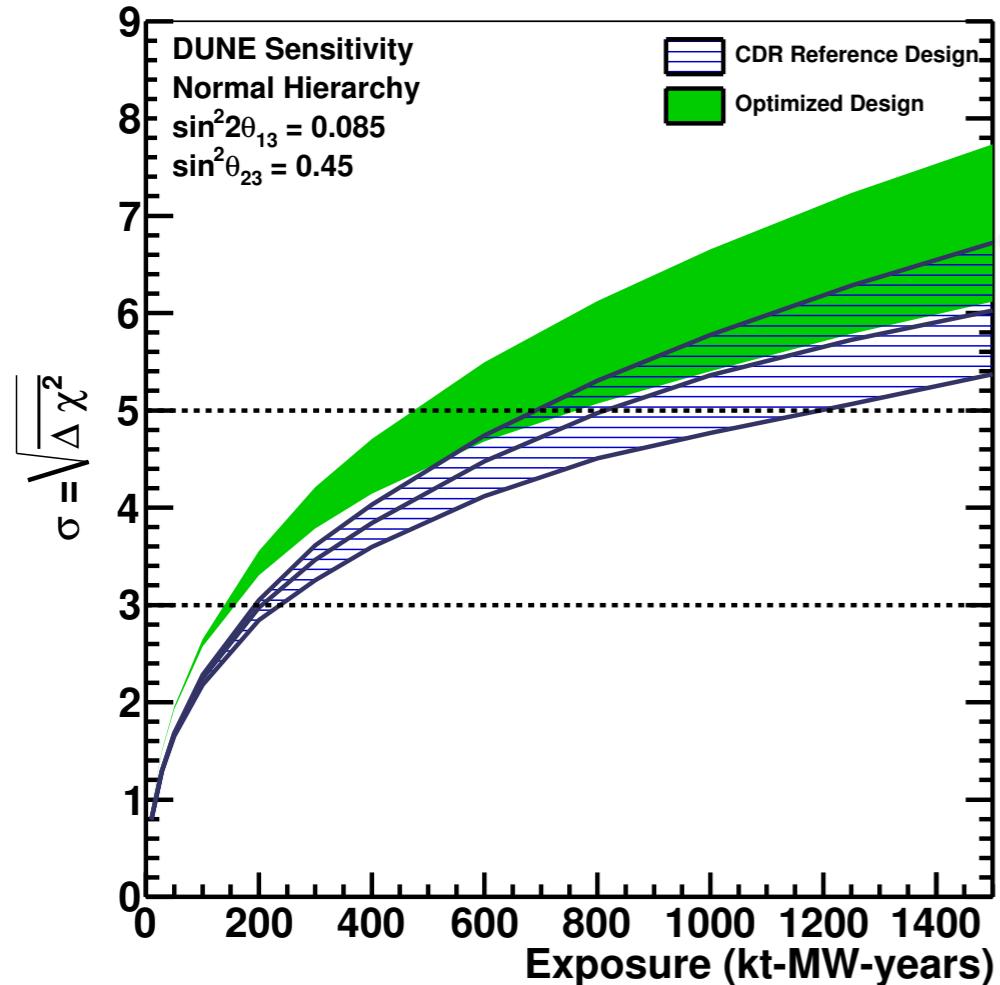


- Clear excess at E=5 MeV
- Recent ab initio calculation [D. Dwyer and T.J. Langford, PRL 114, 012502 (2015)] :
- Excess may be explained by addition of eight isotopes, such as  $^{96}\text{Y}$  and  $^{92}\text{Rb}$

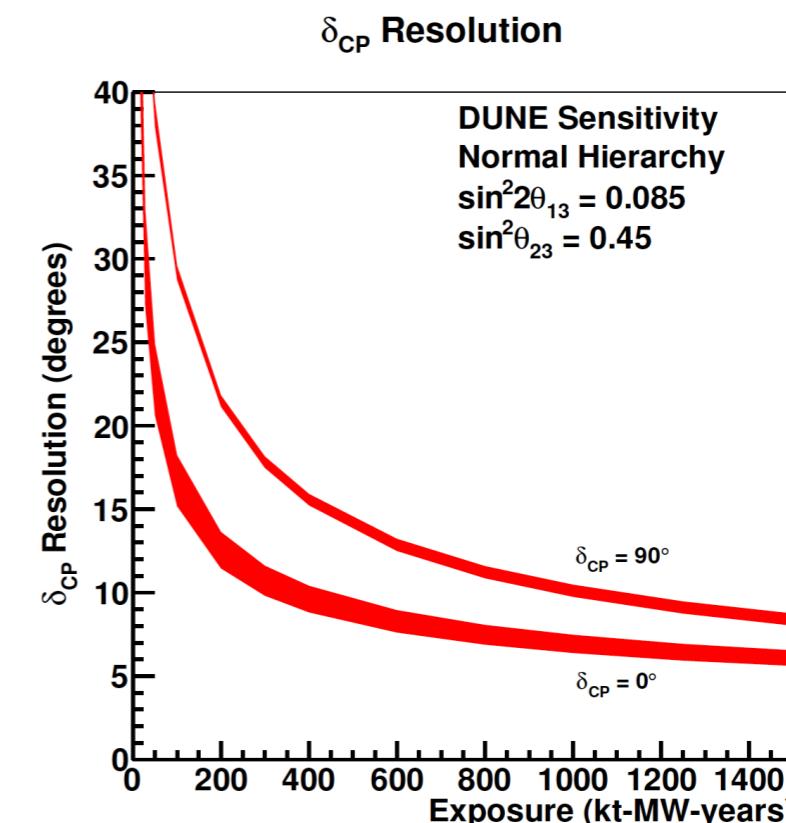
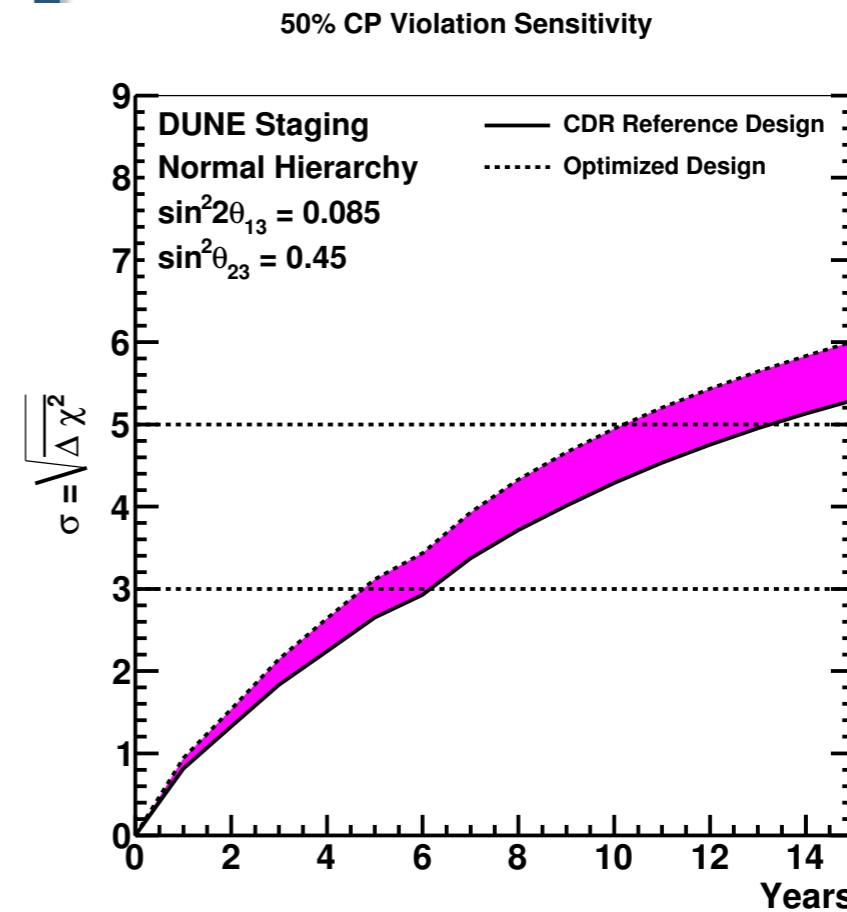
# Future Prospects

## DUNE - Jim Strait

### 50 % CP Violation Sensitivity



Significance	CDR Reference Design	Optimized Design
3σ for 75% of $\delta_{CP}$ values	1320 kt · MW · year	850 kt · MW · year
5σ for 50% of $\delta_{CP}$ values	810 kt · MW · year	550 kt · MW · year

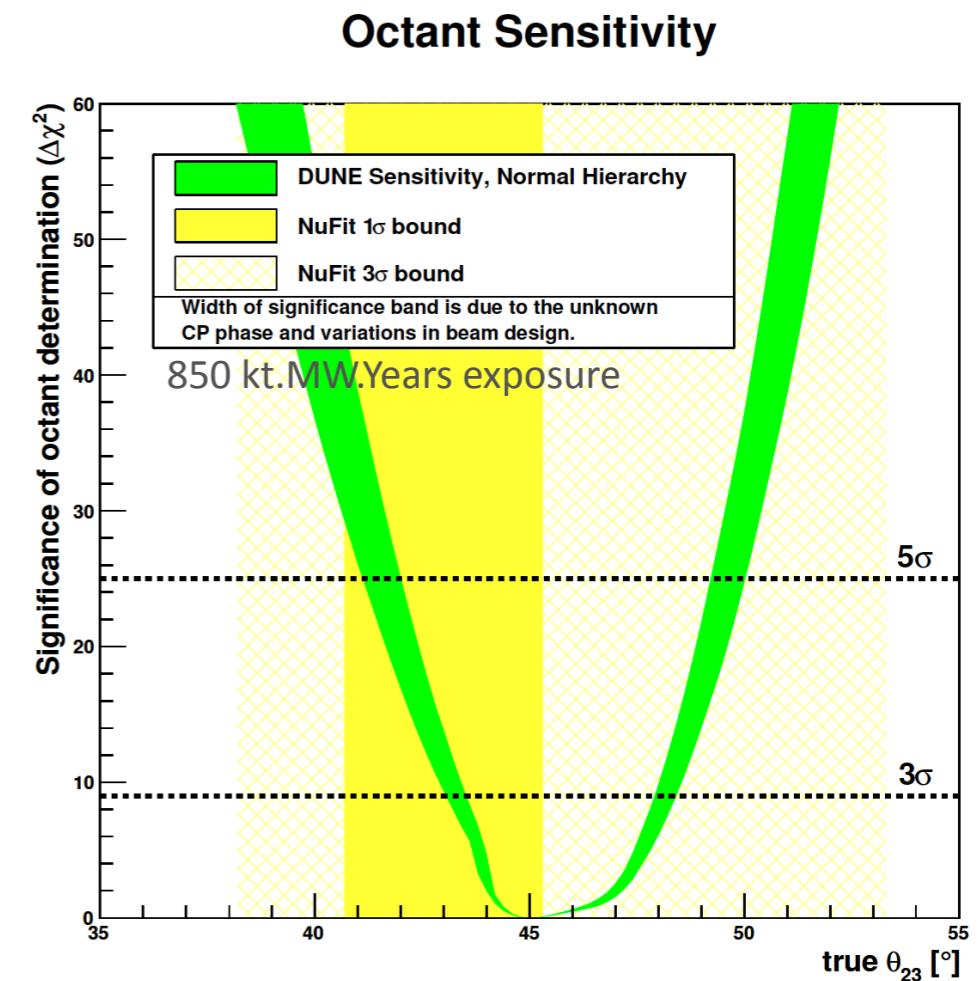
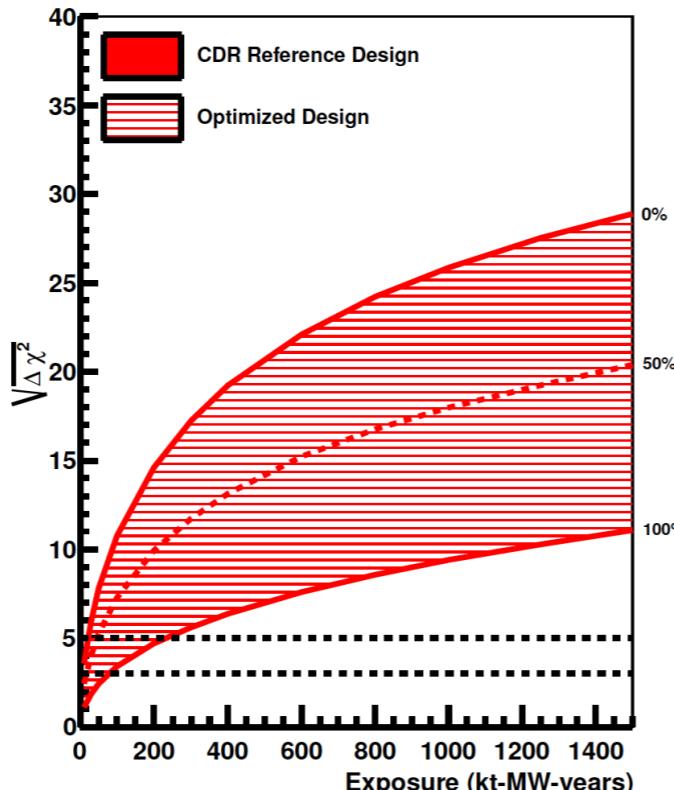
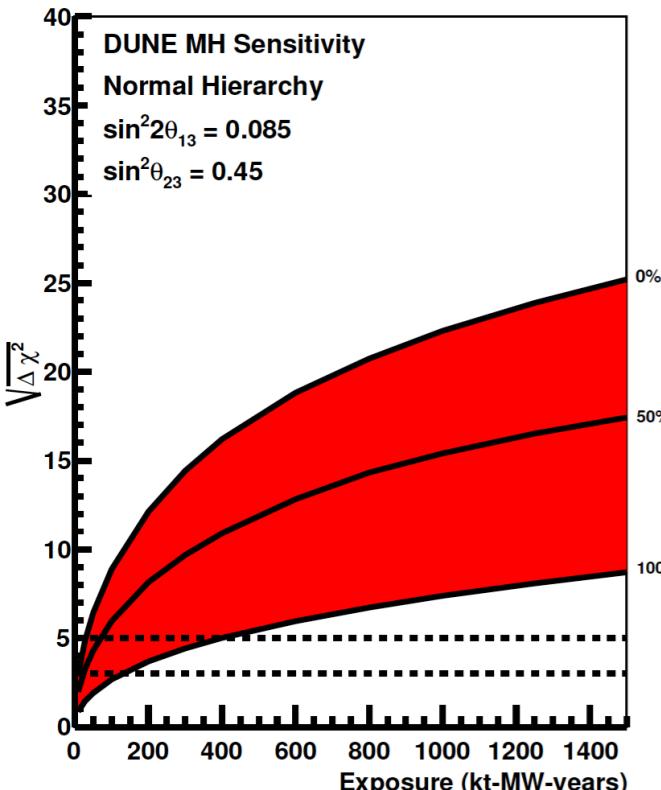


### Sensitivity to CP violation

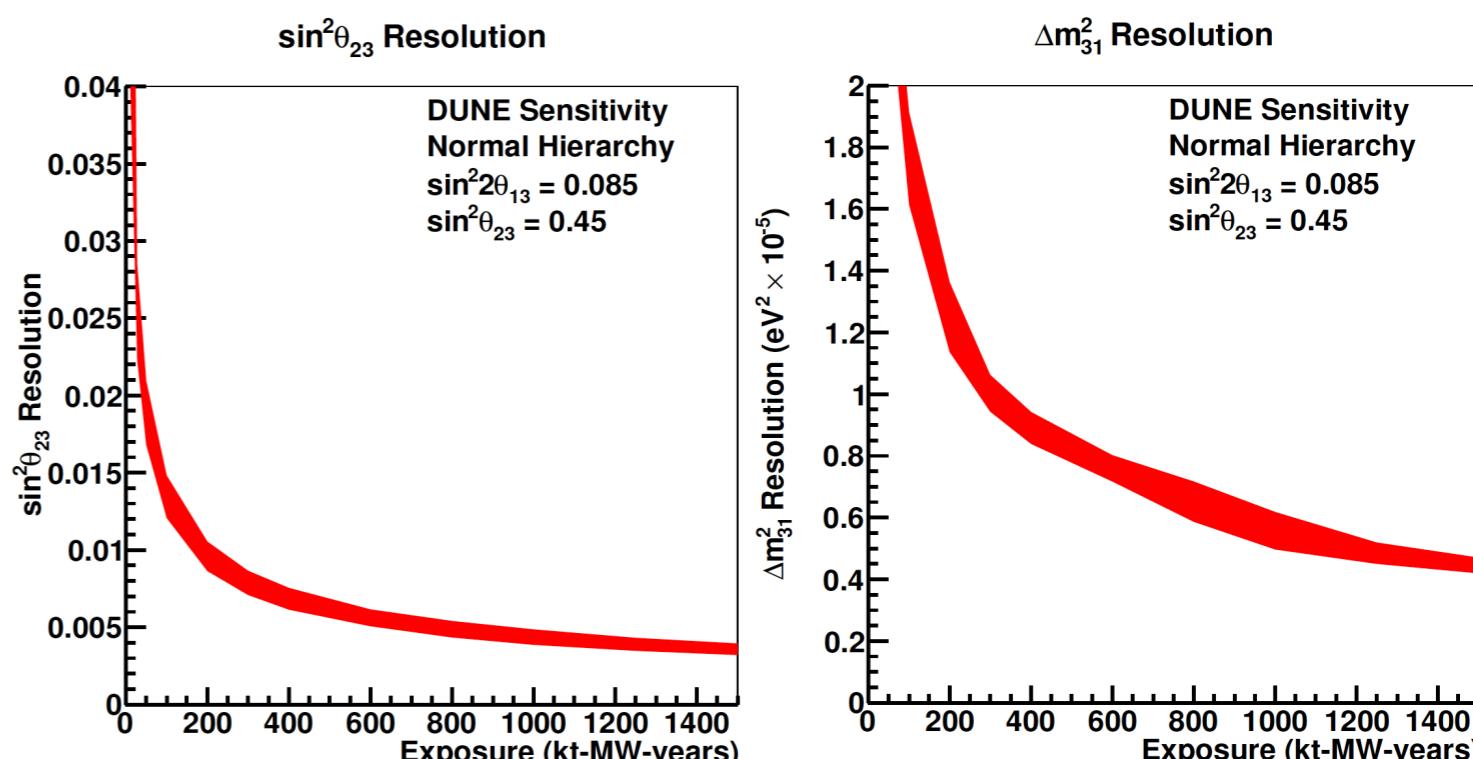
- With present staging plan and optimized design, can reach 5σ for 50% CP space in ~10 years
- $\Delta\delta \sim 10^\circ$  ( $15^\circ$ ) for  $\delta=0^\circ$  ( $90^\circ$ ) after 10 years

# Future Prospects

## DUNE - Jim Strait

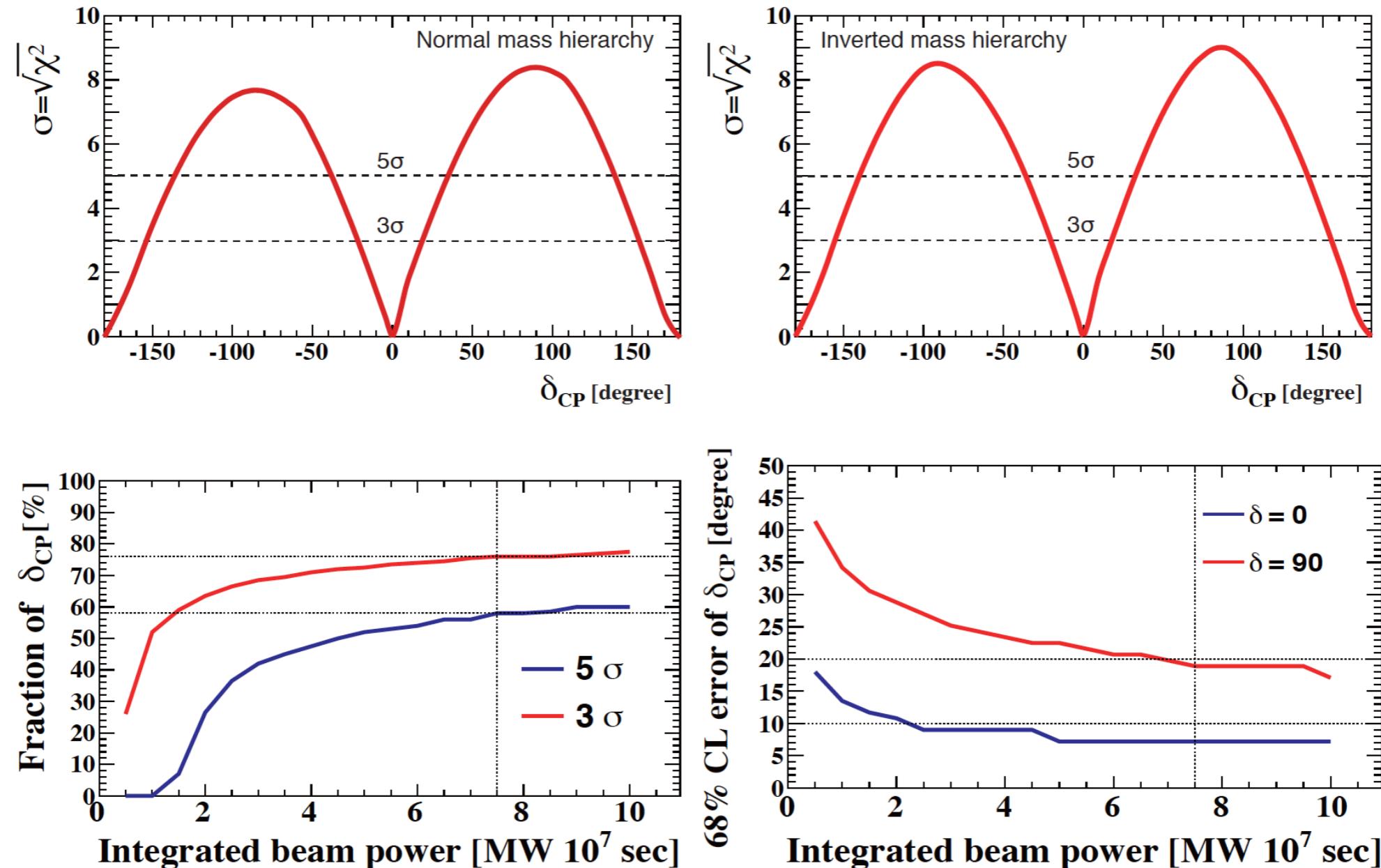


- Sensitivity to mass hierarchy and octant of  $\theta_{23}$
- < 1% precision on atmospheric oscillation parameters



# Future Prospects

## ► HyperK - Tom Feusels

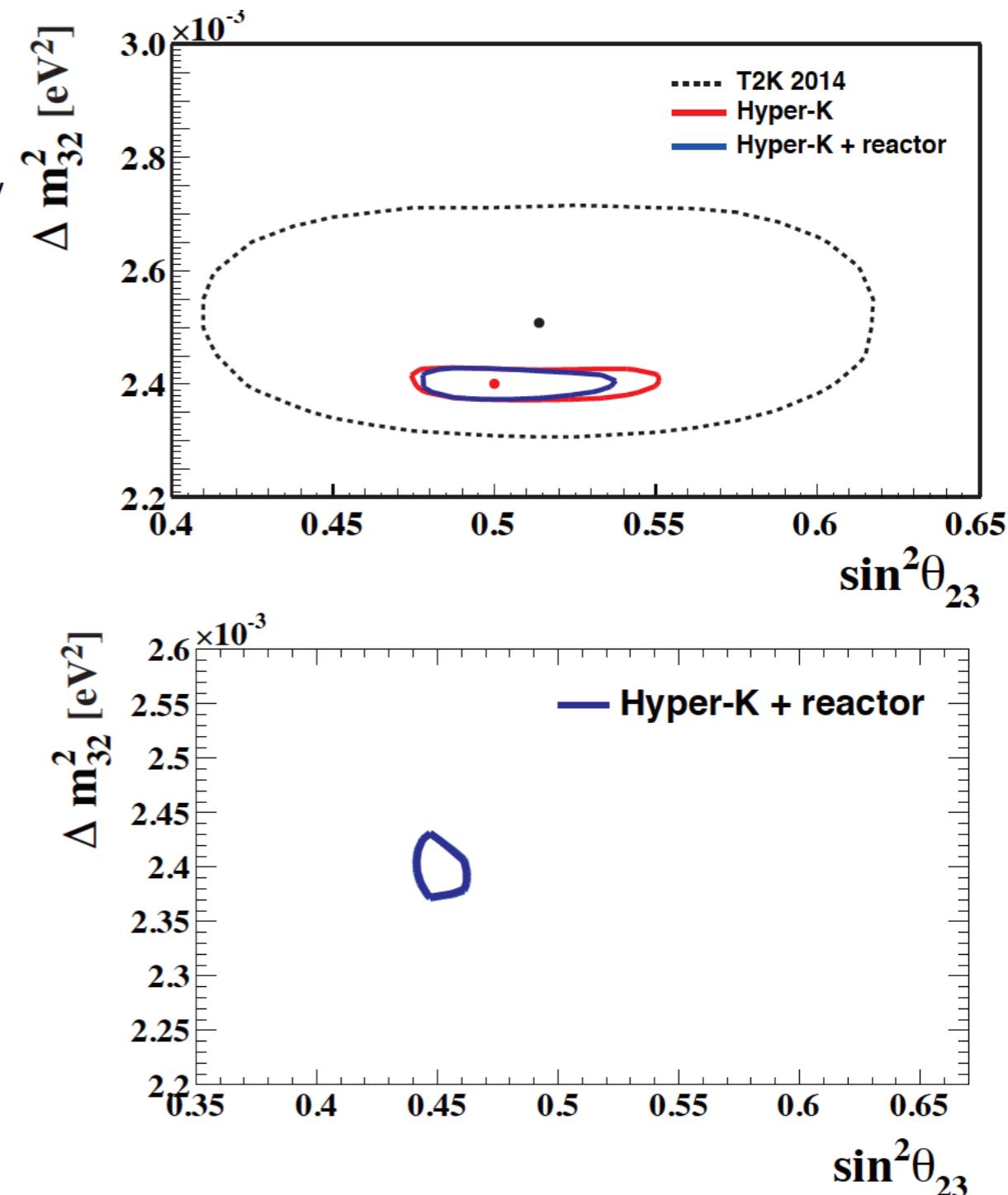
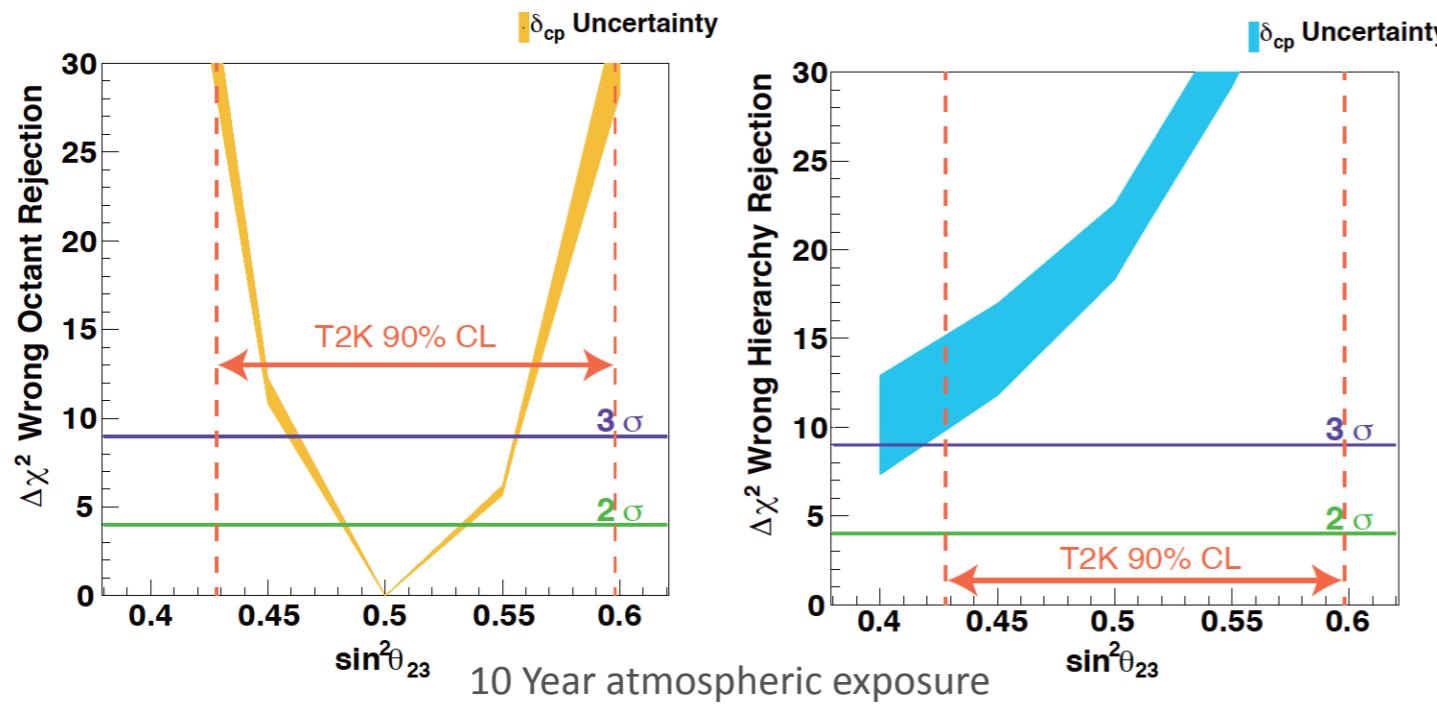


If Mass Hierarchy known:

- Fraction for which  $\sin\delta_{CP}=0$  can be excluded with more than  $3(5)\sigma$  is 76%(58%).
- $\delta_{CP}$  precision is better than  $19^\circ$  for  $7.5MW \times 10^7 \text{ s}$ .

# Future Prospects

## ► HyperK - Tom Feusels

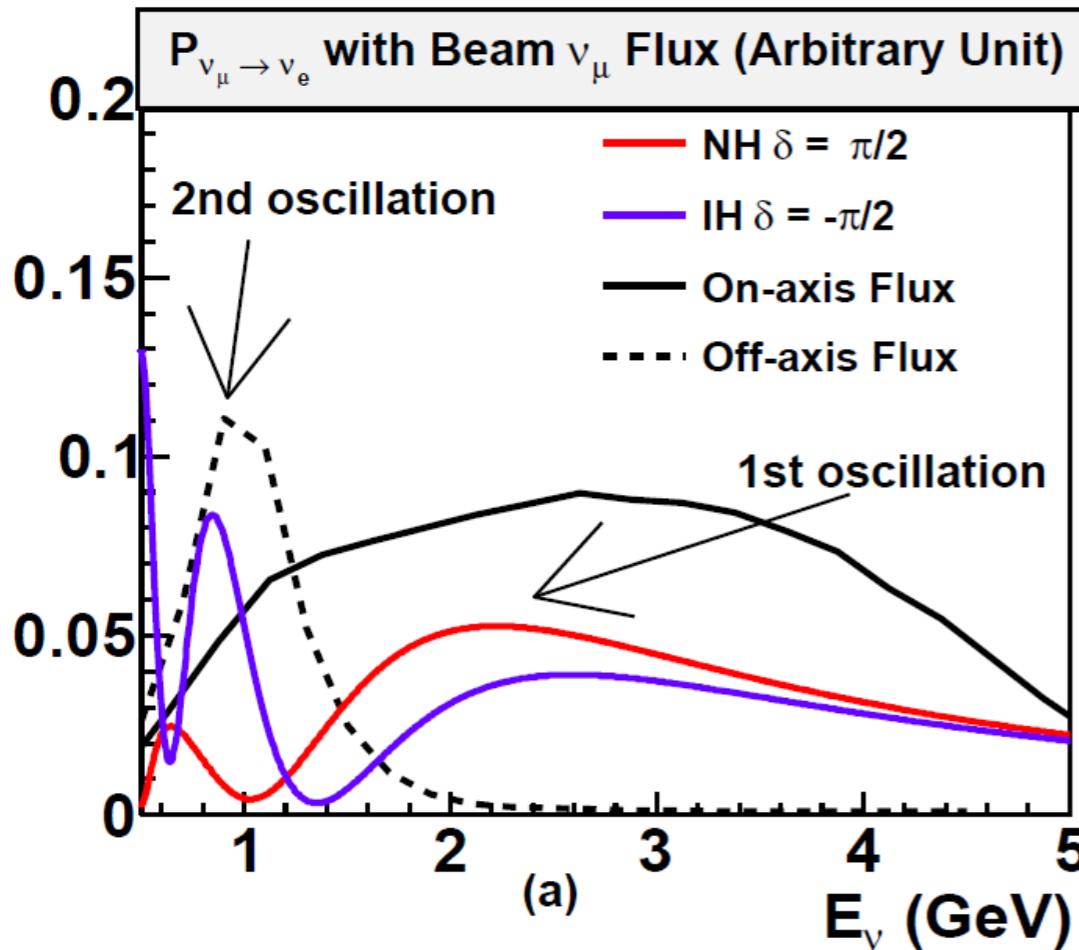


- Can determine octant of  $\theta_{23}$  with constraint from reactor
- Also sensitivity to octant of  $\theta_{23}$  and mass hierarchy (atmospheric sample)
- < 1% precision on atmospheric oscillation parameters

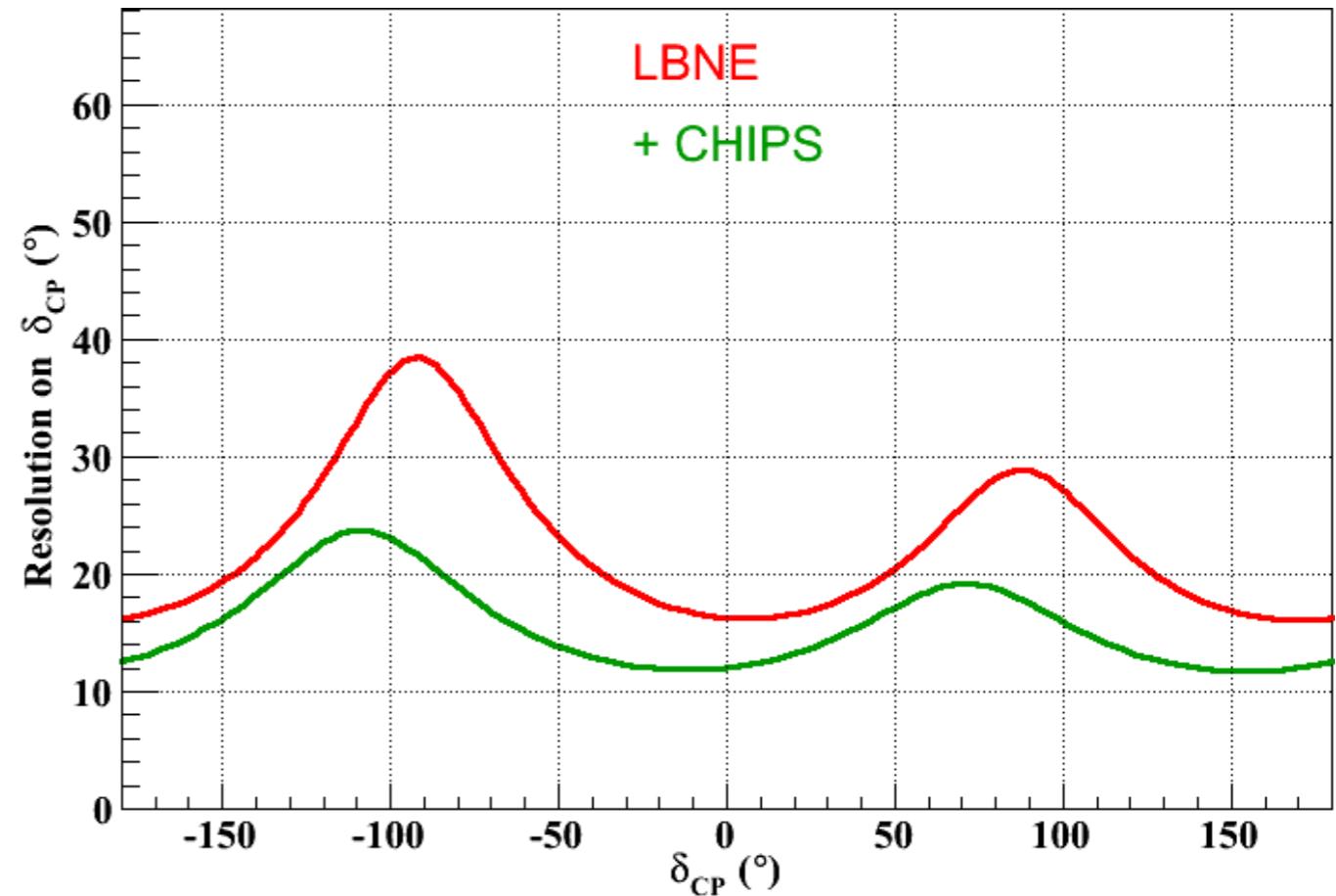
True $\sin^2 \theta_{23}$	$\sin^2 \theta_{23}$	$\Delta m_{32}^2$ ( $10^{-5}$ eV $^2$ )
0.45	0.006	1.4
0.50	0.015	1.4
0.55	0.009	1.5

# Future Prospects

## ► CHIPS - João Coelho



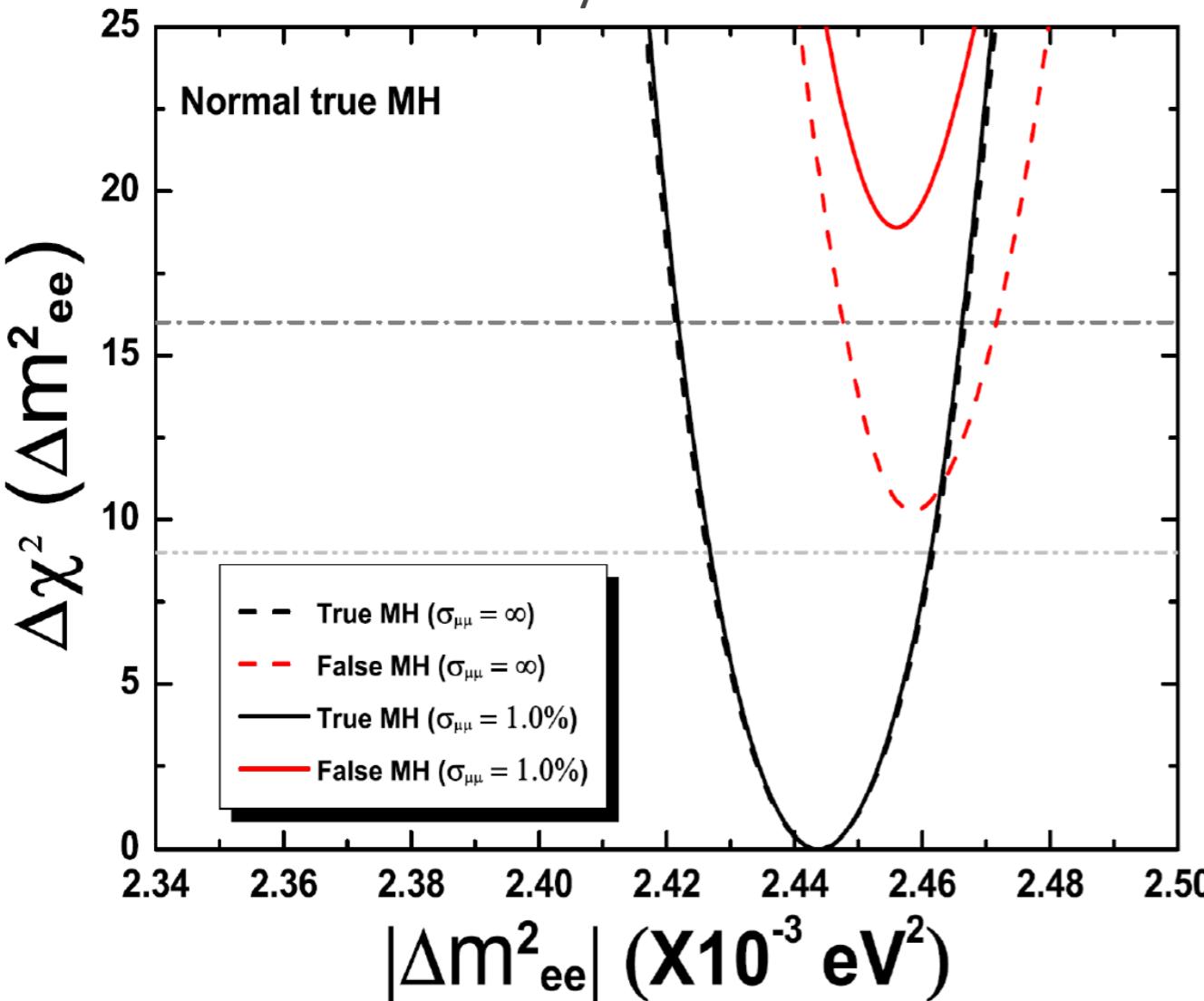
- ▶ Placing a WC detector in a lake 20 mrad off LBNF beam axis provides sensitivity to second osc. maximum and would improve DUNE CPV sensitivity.
- ▶ Can also be placed at mine pit 7 mrad off the NuMI beam axis
- ▶ Ongoing R&D with 30 ton prototype



# Future Prospects

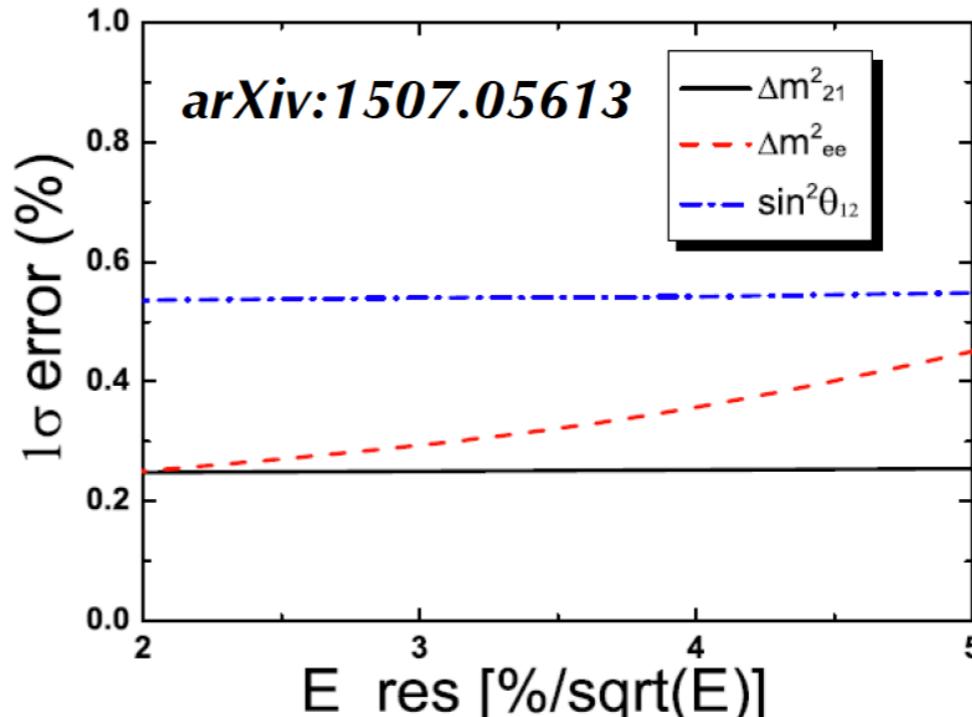
## ► JUNO - Wei Wang

Mass hierarchy determination



- ~3-sigma if only a relative spectral measurement without external atmospheric mass-squared splitting
- ~4-sigma with an external  $\Delta m^2$  measured to ~1% level in  $\nu_\mu$  beam oscillation experiments

► <1% measurements of  $\theta_{12}$ ,  $\Delta m^2_{12}$  and  $\Delta m^2_{ee}$



JUNO: 100k evts  
arXiv:1507.05613

	Nominal	+ B2B (1%)	+ BG	+ EL (1%)	+ NL (1%)
$\sin^2 \theta_{12}$	0.54%	0.60%	0.62%	0.64%	0.67%
$\Delta m^2_{21}$	0.24%	0.27%	0.29%	0.44%	0.59%
$ \Delta m^2_{ee} $	0.27%	0.31%	0.31%	0.35%	0.44%

As of July 6, digging 422 meters (of 1340.6 meters)  
Roughly 4 meters/day    491m on July 26  
Rock type-III  
Little underground water leakage

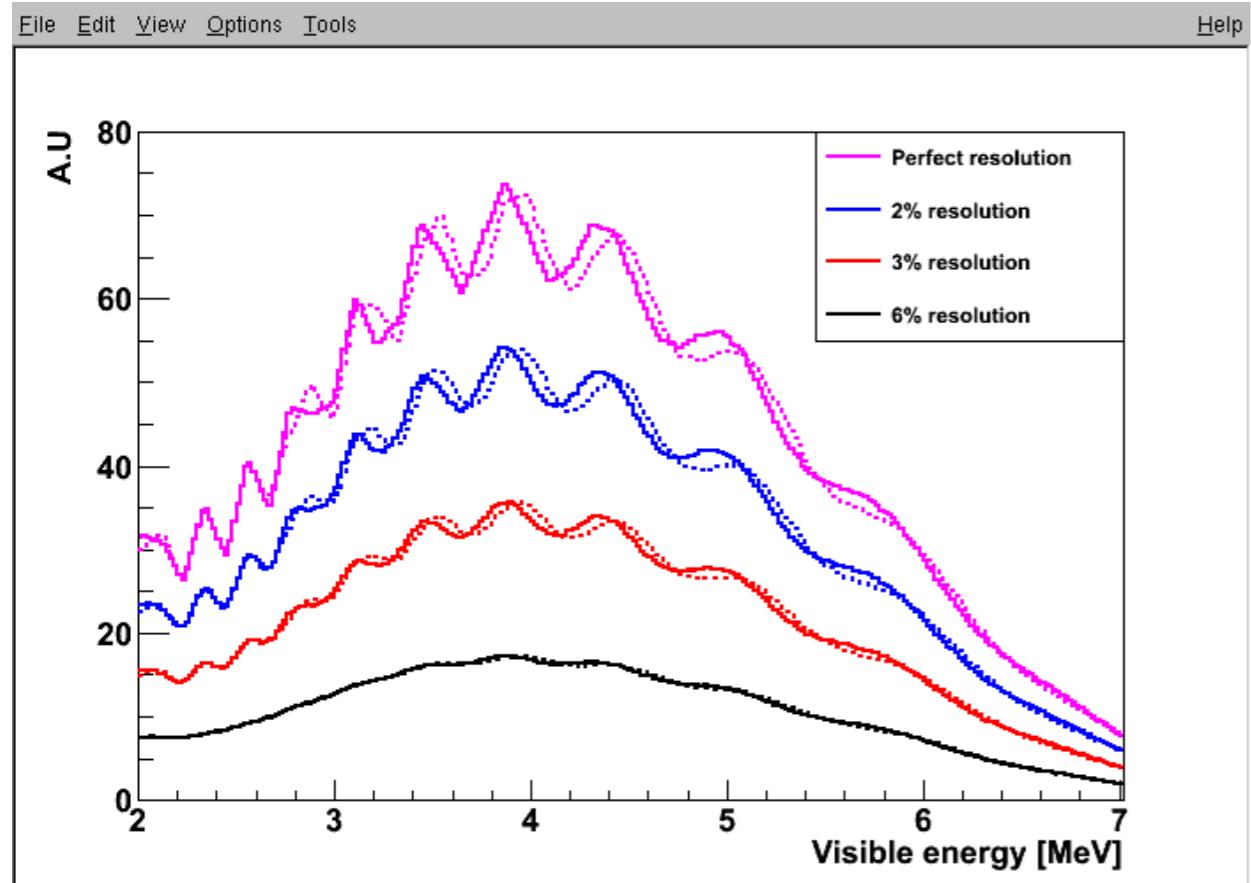
Slope Tunnel



► Construction underway!

# Future Prospects

## ► RENO-50 - Kyungkwang Joo



**3% energy resolution essential for distinguishing the oscillation effects between normal and inverted mass hierarchies**

► Proposal submitted for full funding

- Precise measurement of  $\theta_{12}$ ,  $\Delta m^2_{21}$  and  $\Delta m^2_{32}$

$$\frac{\delta \sin^2 \theta_{12}}{\sin^2 \theta_{12}} < 1.0\%(1\sigma) \quad (\leftarrow 5.4\%) \quad \frac{\delta \Delta m^2_{21}}{\Delta m^2_{21}} < 1.0\%(1\sigma) \quad (\leftarrow 2.7\%) \quad \frac{\delta \Delta m^2_{32}}{\Delta m^2_{32}} < 1.0\%(1\sigma) \quad (\leftarrow 5.2\%)$$

- Neutrino burst from a Supernova in our Galaxy

- ~5,600 events (@8 kpc)

- A long-term neutrino telescope

- Geo-neutrinos : ~ 1,000 geo-neutrinos for 5 years

- Study the heat generation mechanism inside the Earth

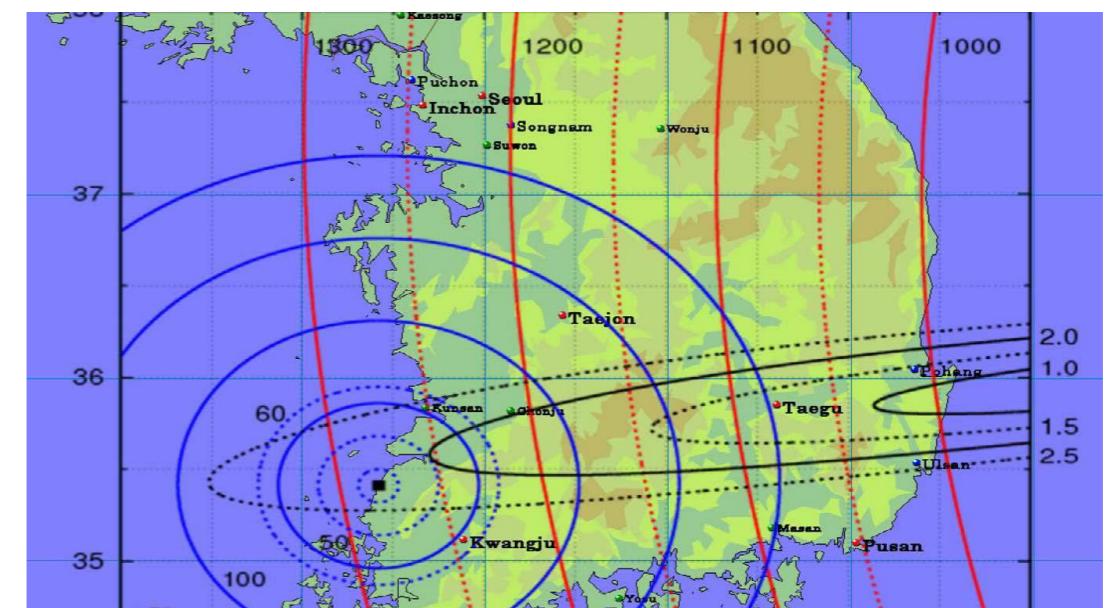
- Solar neutrinos : with ultra low radioactivity

- MSW effect on neutrino oscillation

- Probe the center of the Sun and test the solar models

- Neutrinoless double beta decay search : possible modification like KamLAND-Zen

- Detection of J-PARC beam : ~200 events/year

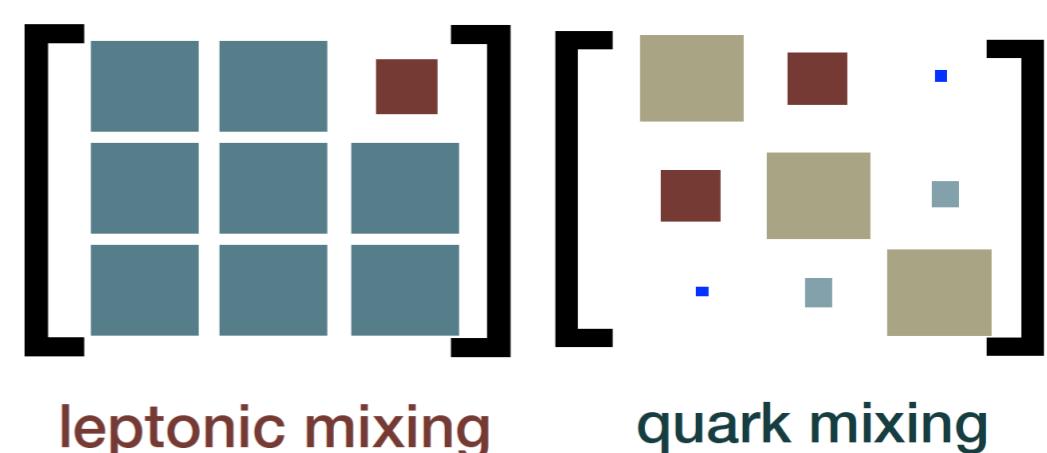
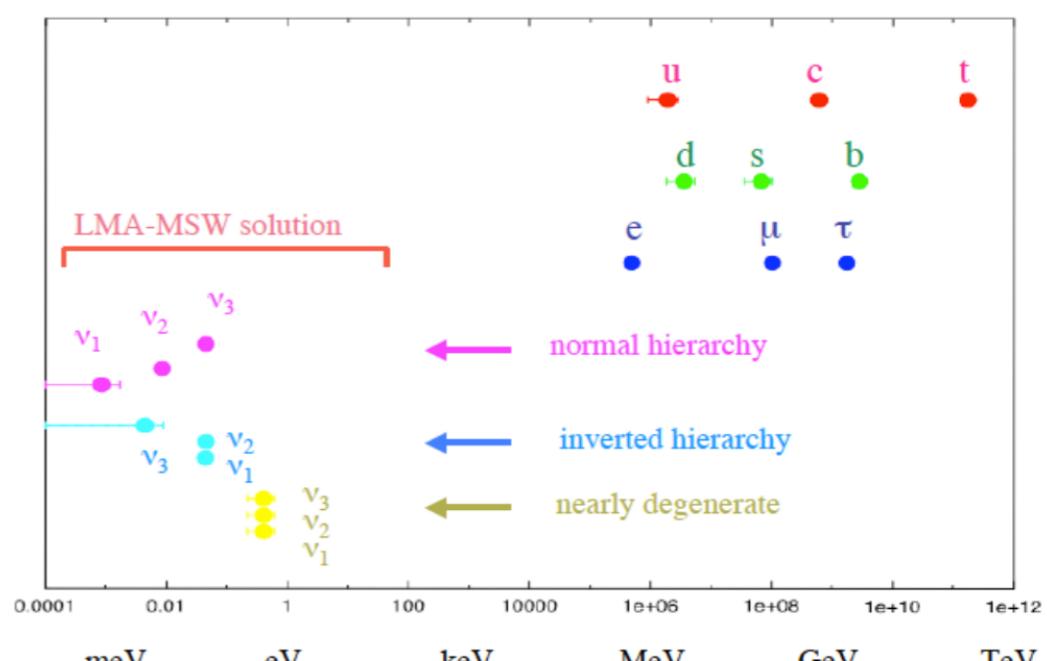


# Answer to Q1

**Question from NuFact '14:** What are the new developments and predictions from flavor models on neutrino oscillation parameters? What precision do we need to achieve to probe them? Which parameters (or combinations of them) are more powerful to test them?

- ▶ Non-zero value of  $\theta_{13}$  rules out models like tribimaximal mixing and golden ratio models
- ▶ Perturbative corrections to TBM result in sum rules correlating mixing angles and CP phase, which can be tested by precision measurements of  $\theta_{12}, \theta_{23}, \delta_{CP}$
- ▶ Present experimental data still allows for a wide variety of models
- ▶ <1% precision on mixing angles achievable in next 10-15 years. 10°-15° precision on  $\delta_{CP}$  achievable after several years of DUNE and/or HyperK running.

$$m_\nu \ll M_e, u, d$$

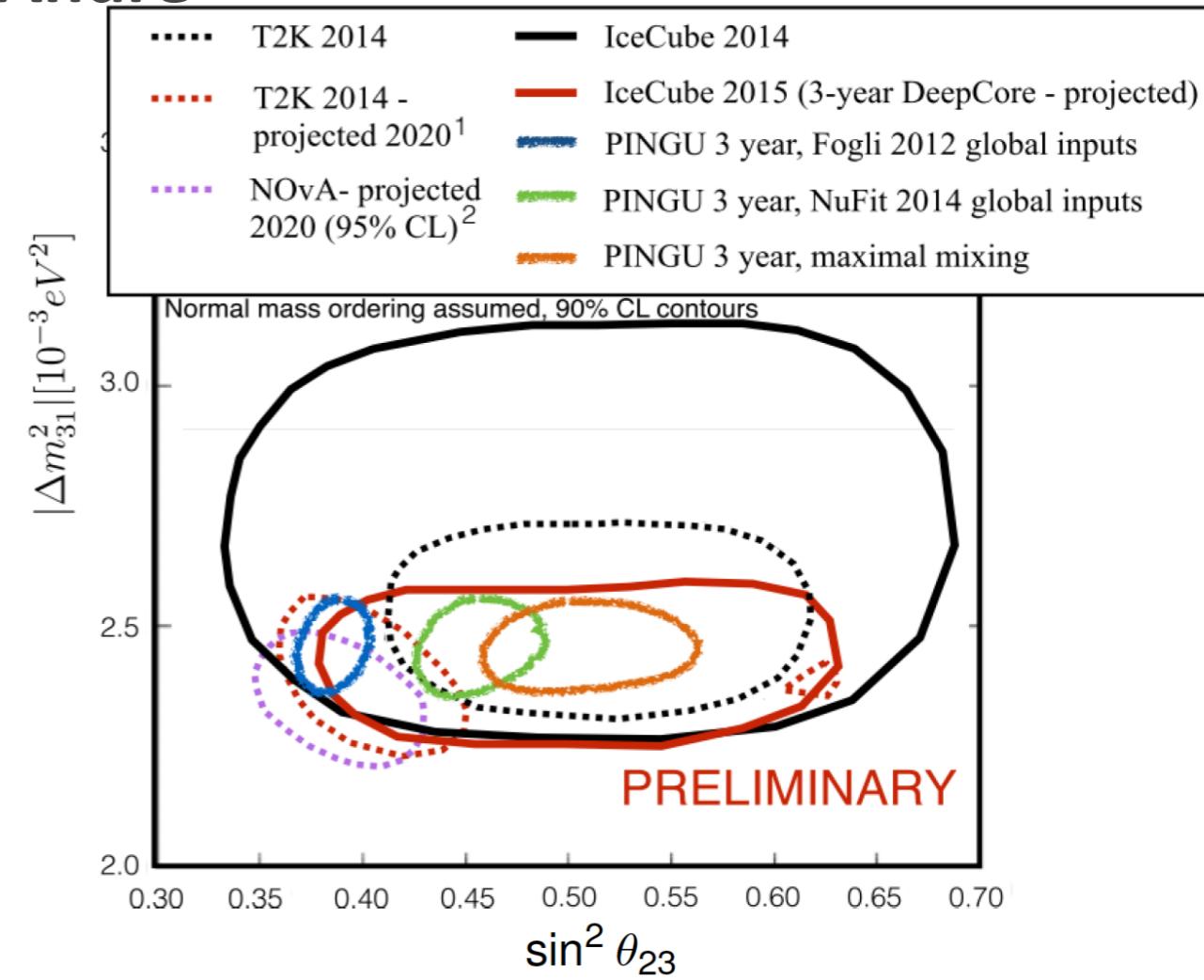
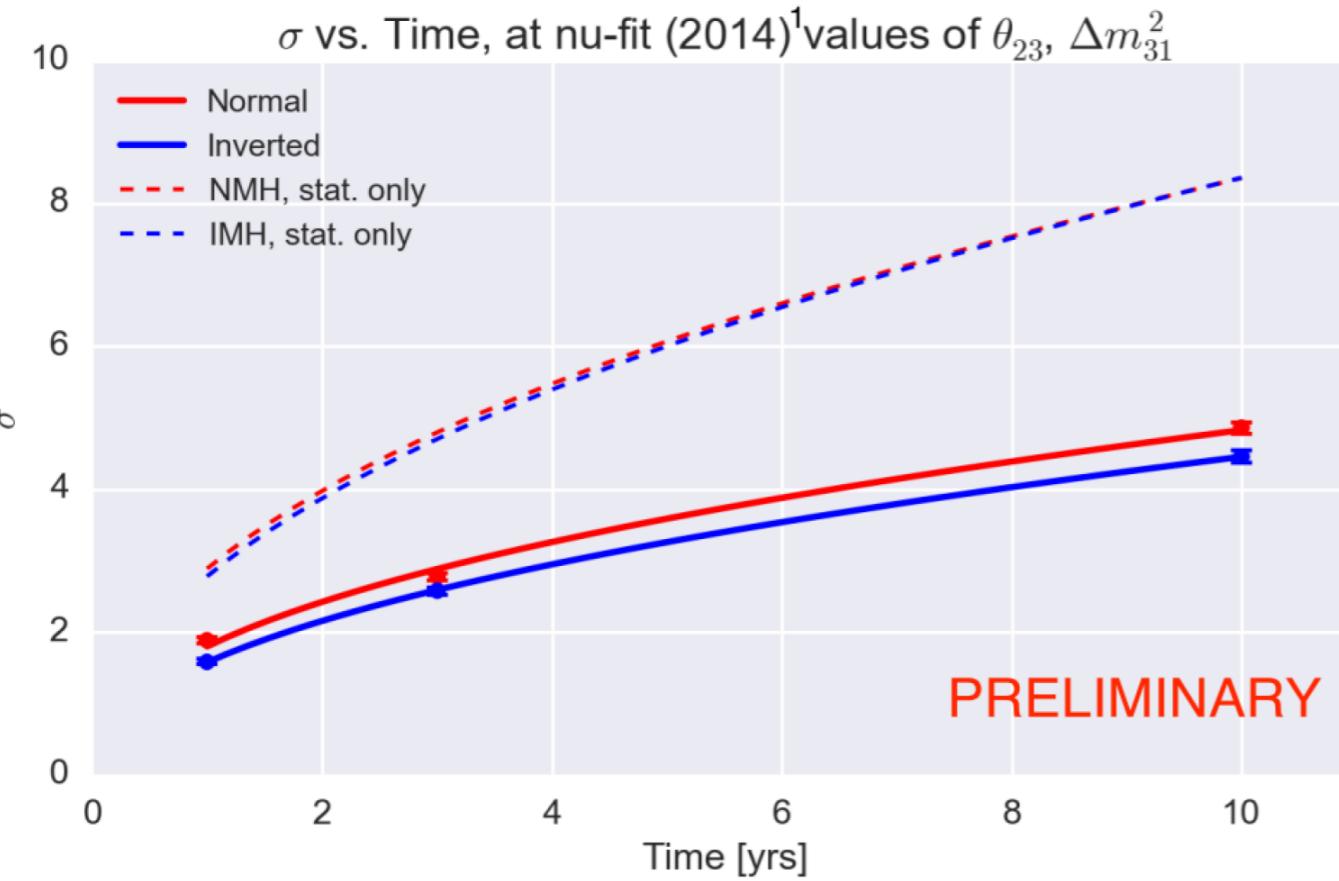


# Q2: Atmospheric and LBL

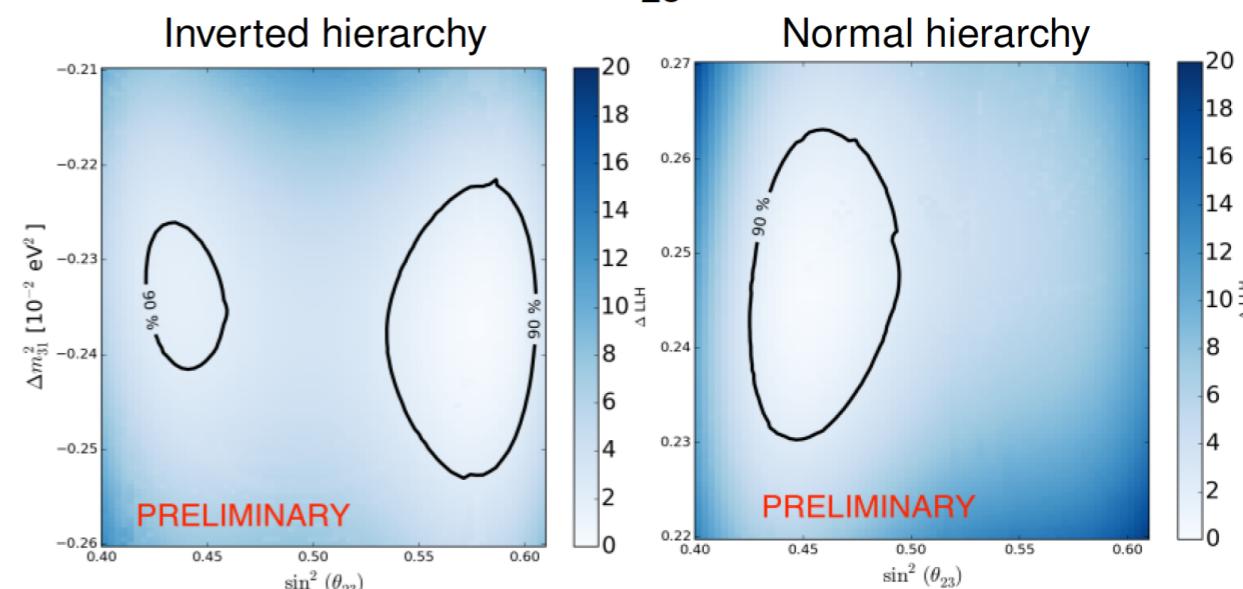
**Question from NuFact '14:** Are atmospheric neutrino measurements competitive with next generation long-baseline facilities in the determination of the mass hierarchy? And the octant of  $\theta_{23}$ ? How much complementarity is there between them?

# Atmospheric v Future Projects

## ► PINGU - JP Athayde Marcondes de André

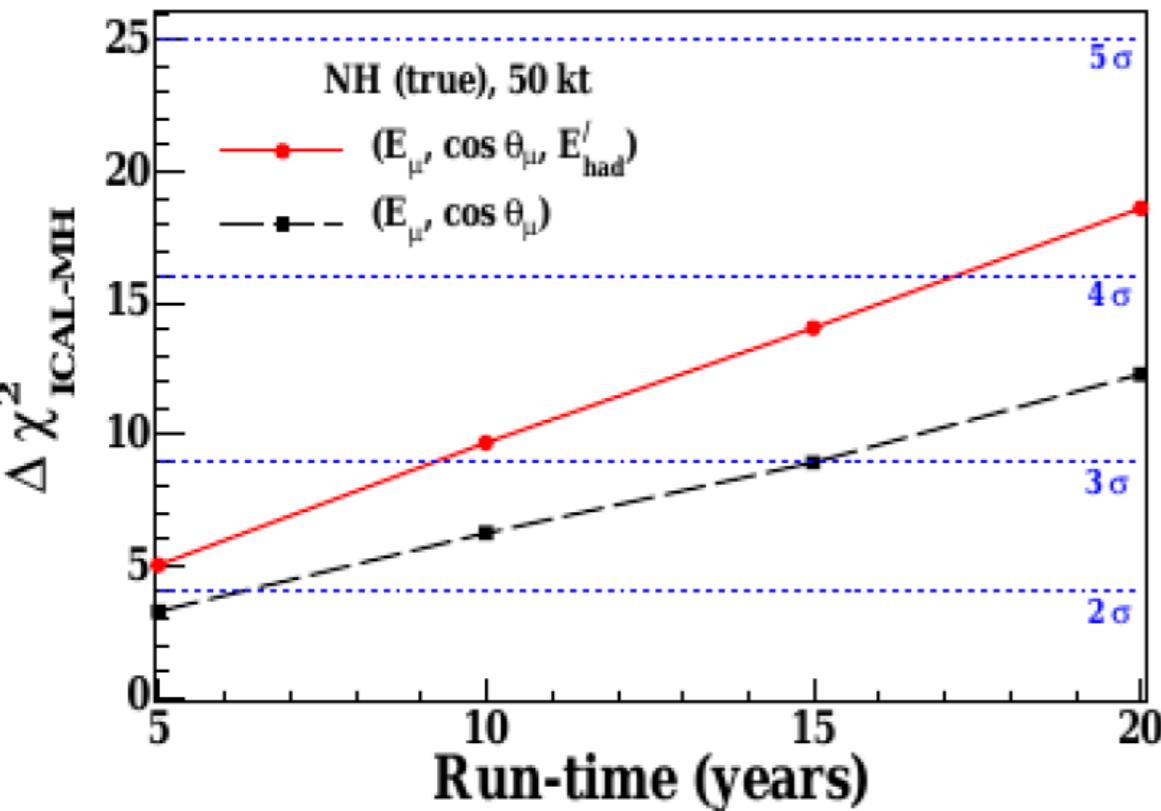


- Can reach  $3\sigma$  resolution of mass hierarchy after 3 years of running.
- Mass hierarchy sensitivity strongly dependent on value of  $\theta_{23}$
- Expect to achieve similar  $\theta_{23}$  and  $\Delta m^2_{32}$  precision as NOvA and T2K

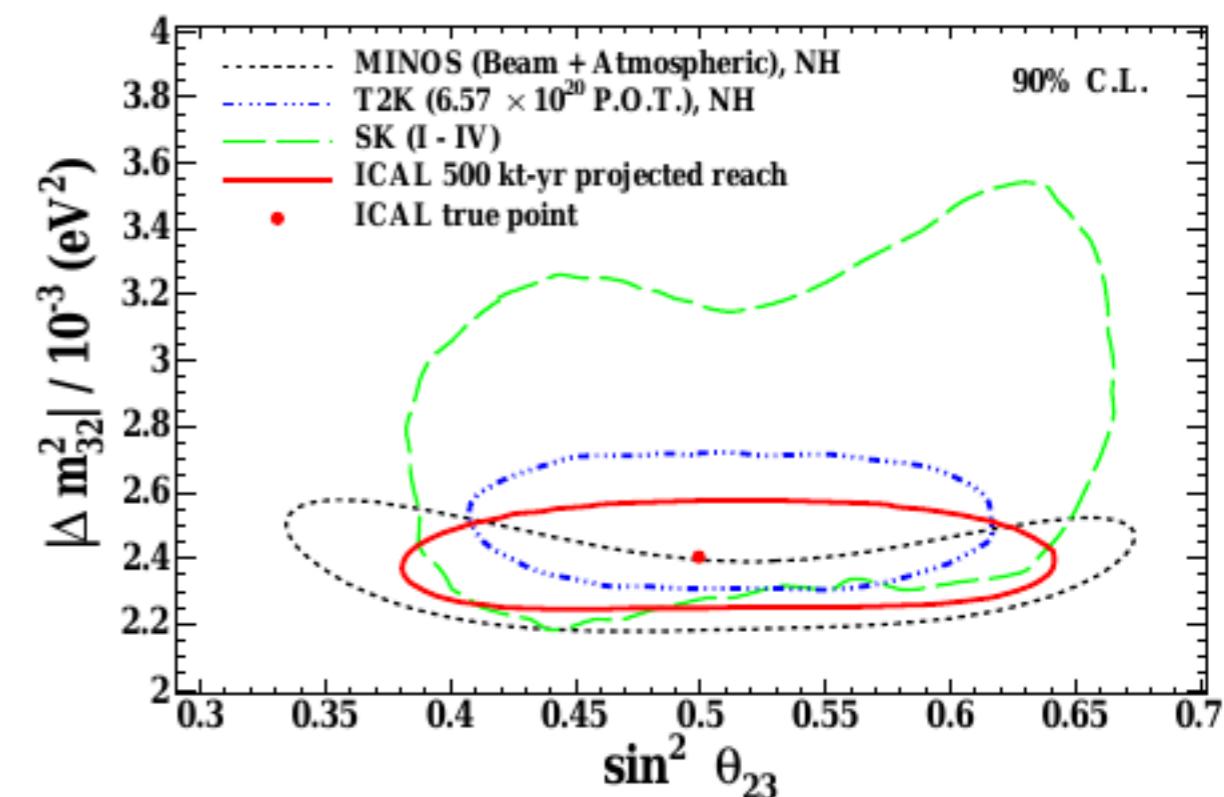
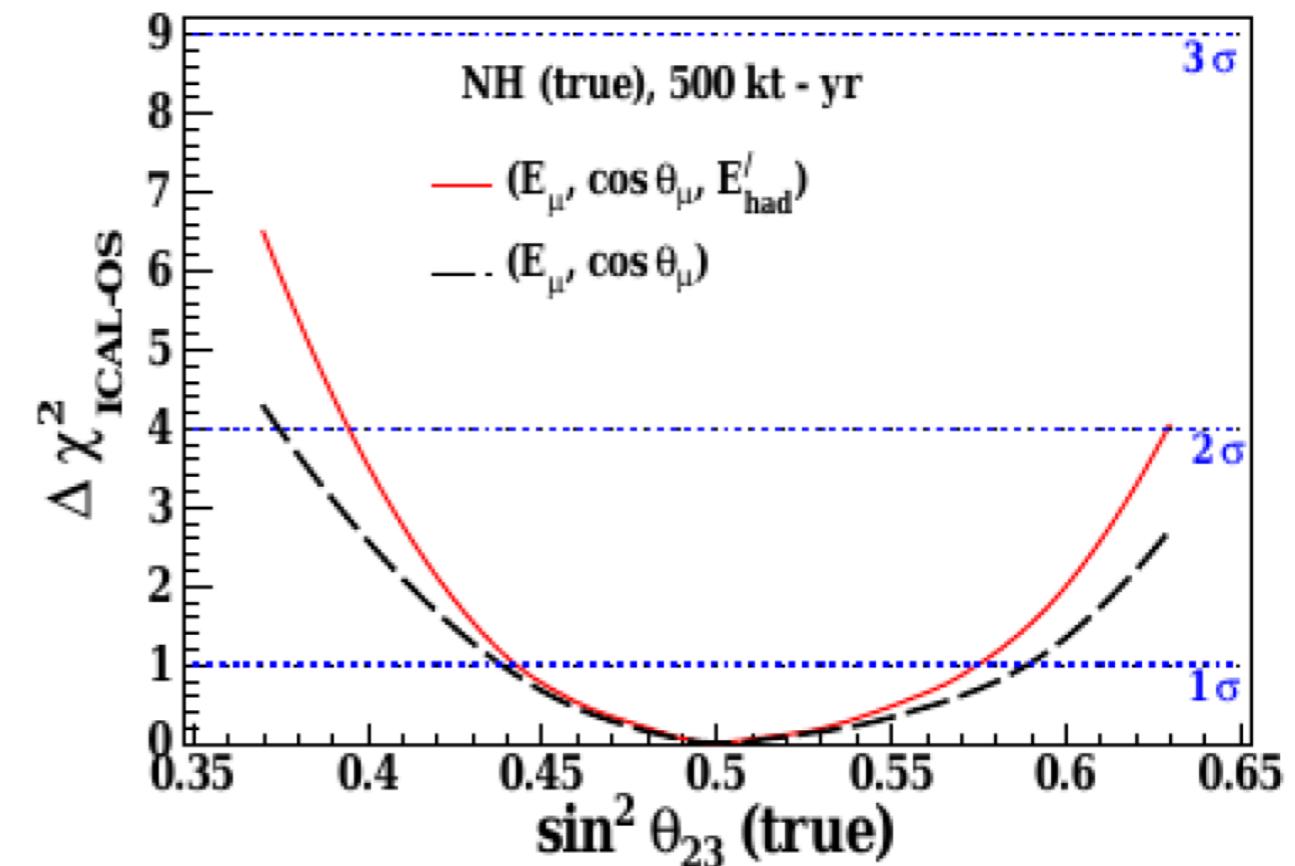


# Atmospheric v Future Projects

## ► INO - Sanjeev Kumar Veerma

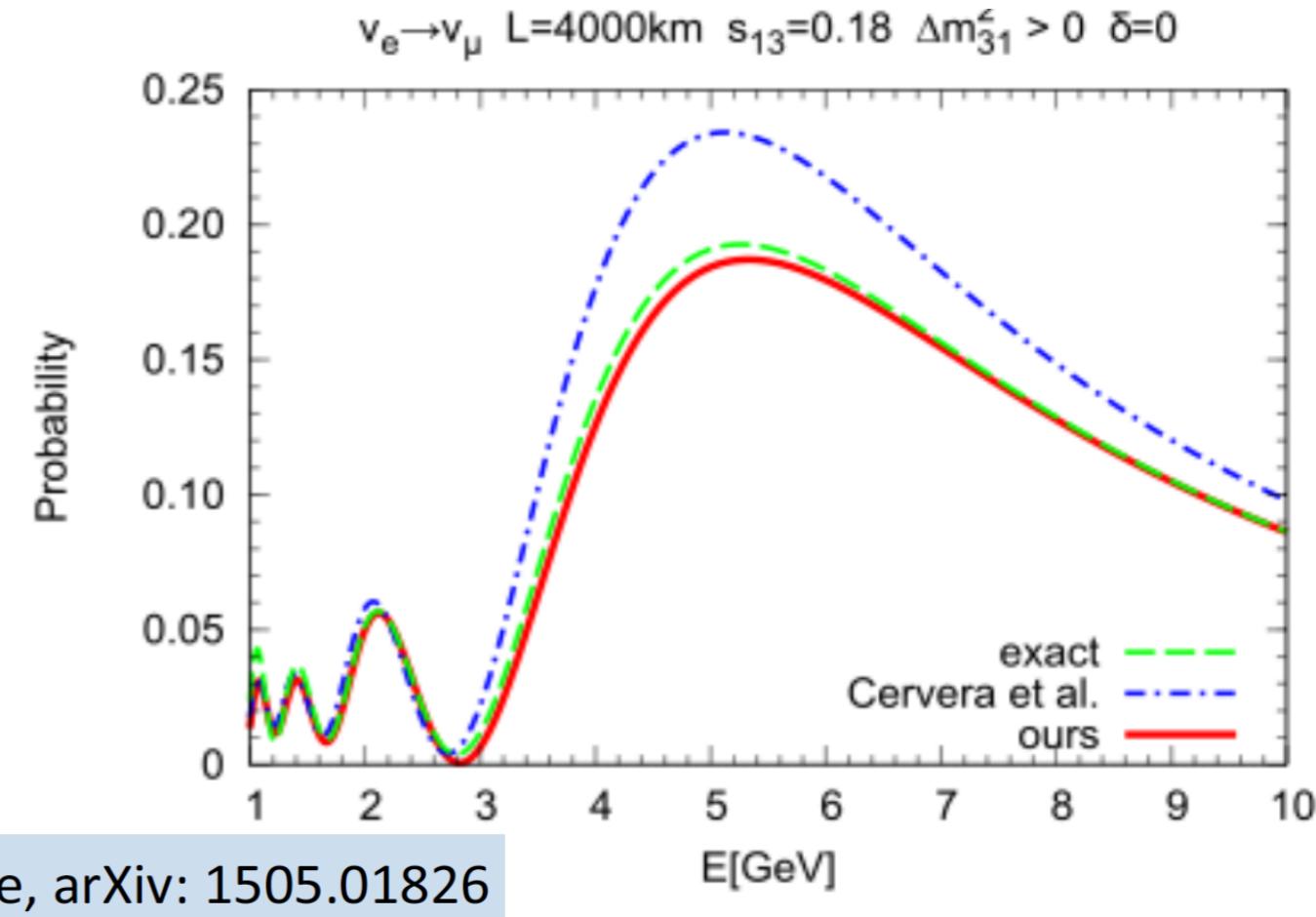
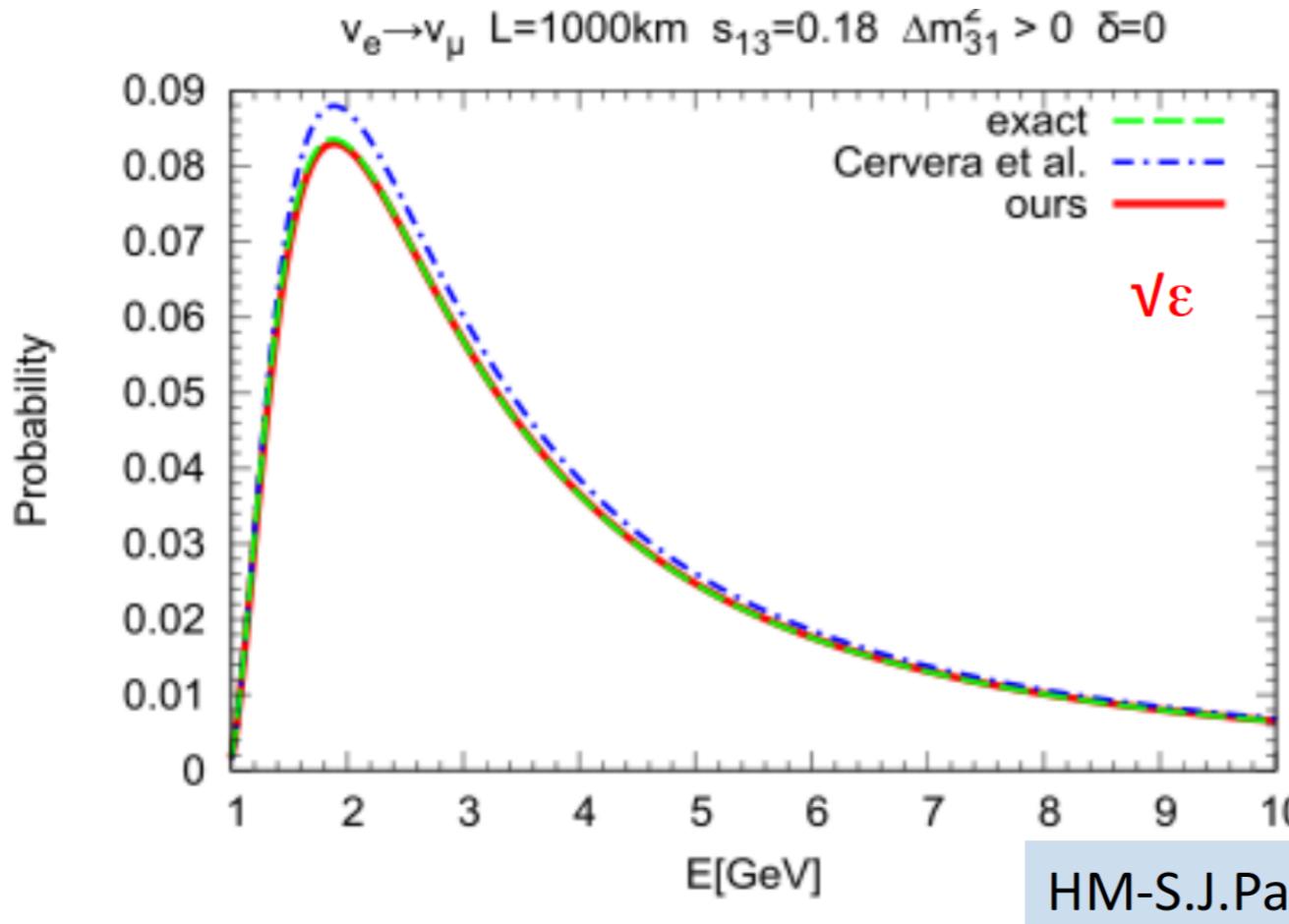


- Can reach 3 $\sigma$  resolution of mass hierarchy after 10 years of running. Improved to 6 years if combined with NOvA and T2K
- 2 $\sigma$  sensitivity to octant of  $\theta_{23}$  after 10 years of running.



# $\nu$ Osc. Formulae

## ► Hisakazu Minakata - Compact formulas for $\nu$ oscillations in matter



HM-S.J.Parke, arXiv: 1505.01826

- Expansion of 3-flavor formulae in  $\epsilon = \Delta m_{21}^2 / \Delta m_{31}^2$
- Good match to exact formulation for large distances relevant to atmospheric neutrino propagation
- Should improve computational time w.r.t to exact formulation
- Evaluating validity for expansions in  $O(\epsilon^2)$

# Answer to Q2

**Question from NuFact '14: Are atmospheric neutrino measurements competitive with next generation long-baseline facilities in the determination of the mass hierarchy? And the octant of  $\theta_{23}$ ? How much complementarity is there between them?**

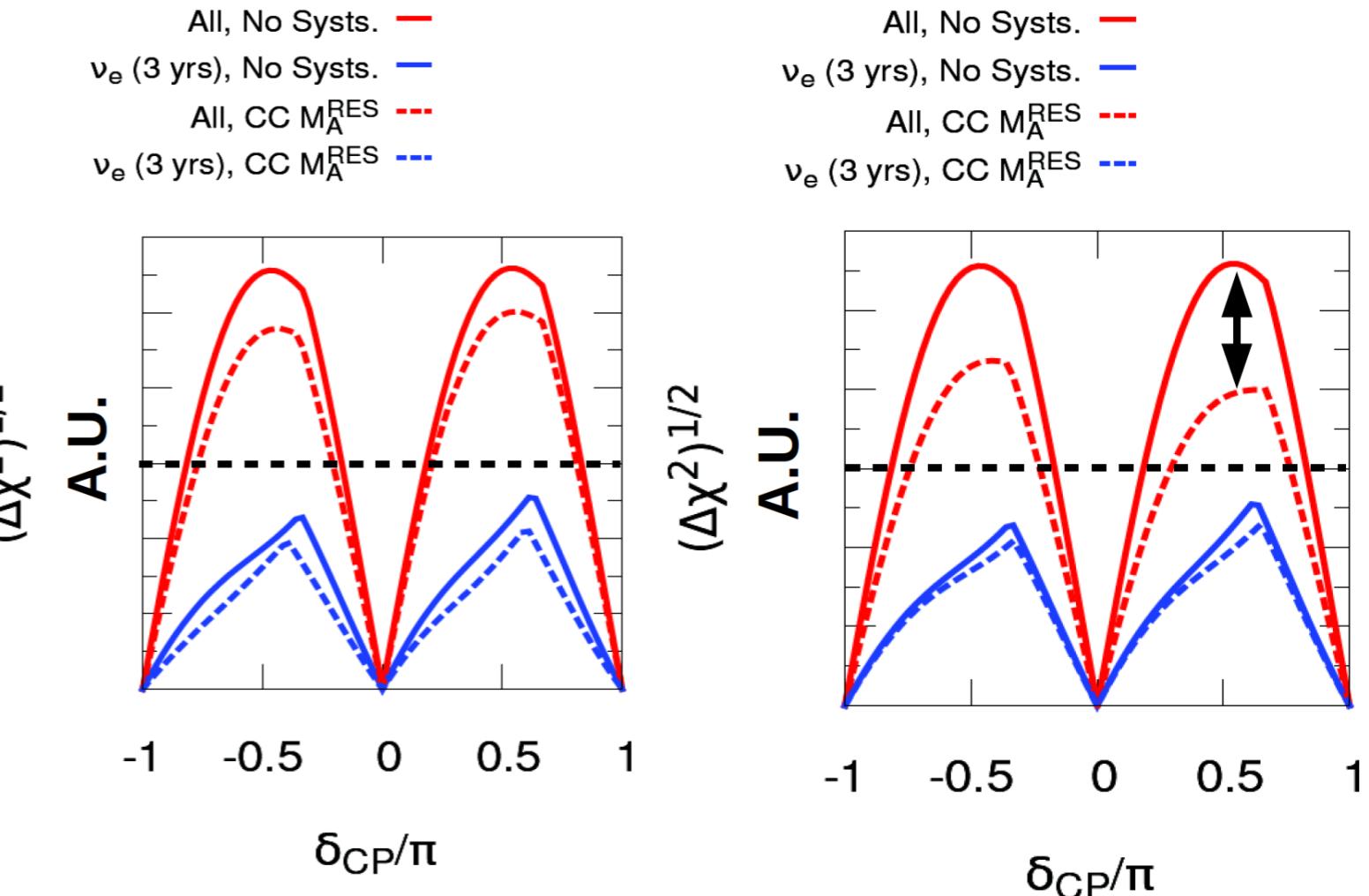
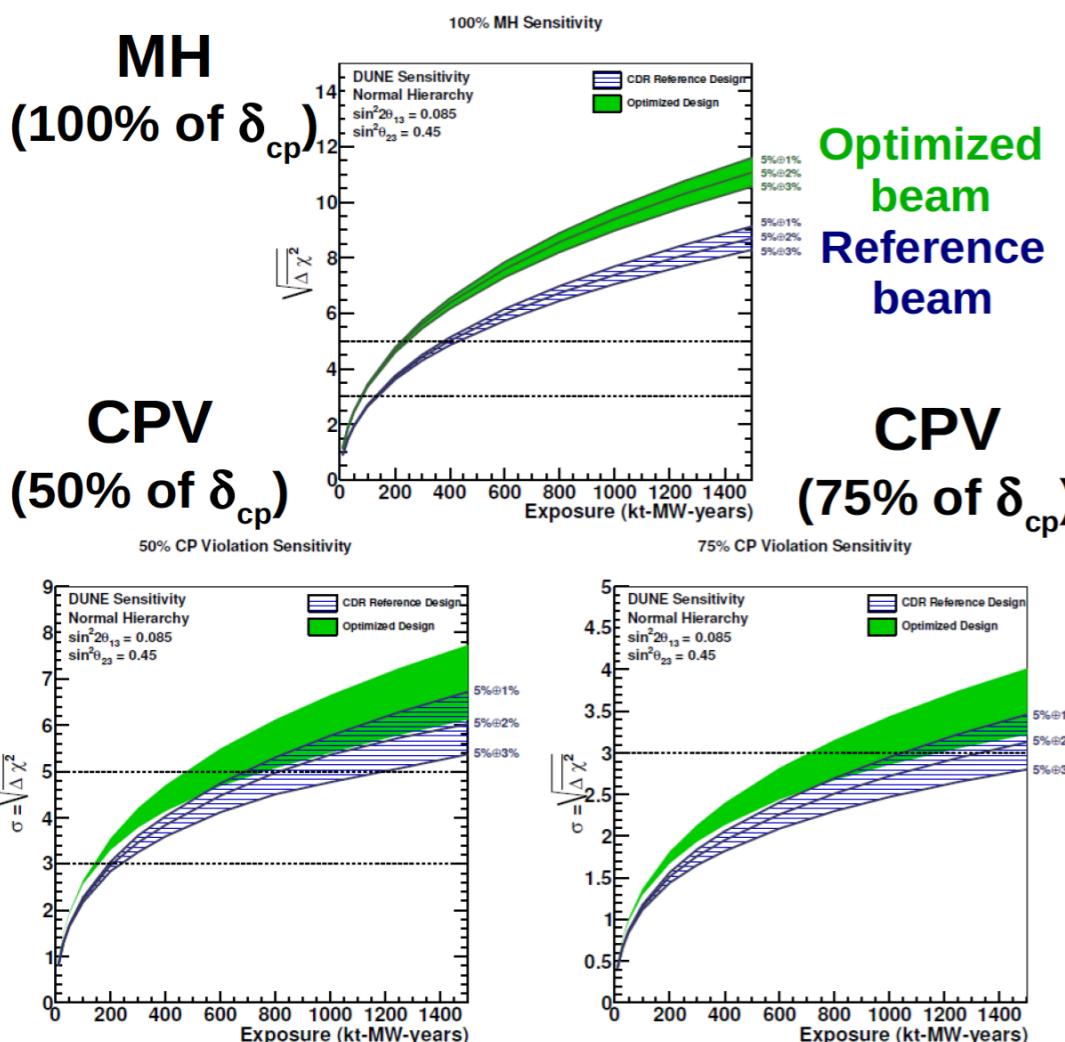
- ▶ Projects like PINGU or ORCA may potentially reach over  $3\sigma$  in mass hierarchy determination before turn-on of next generation long-baseline facilities
- ▶ In combination with INO, and atmospheric data from HyperK and DUNE, they will reinforce global measurements of mass hierarchy
- ▶ Determination of octant of  $\theta_{23}$  from atmospheric data difficult for proposed atmospheric detectors. To be continued at future NuFacts

# Q3: Systematics

**Question from NuFact '14:** What is the target for the systematic error budget of next generation facilities? What do we need to reach this level? How much improvement in constraining flux uncertainties can we expect from dedicated hadron production efforts?

# Systematics

- DUNE - Dan Cherdack
- Showing effect of cross-section systematics (left) + flux uncertainties (right) on CP Violation sensitivity for FD-only fits
- Smaller effect on mass hierarchy



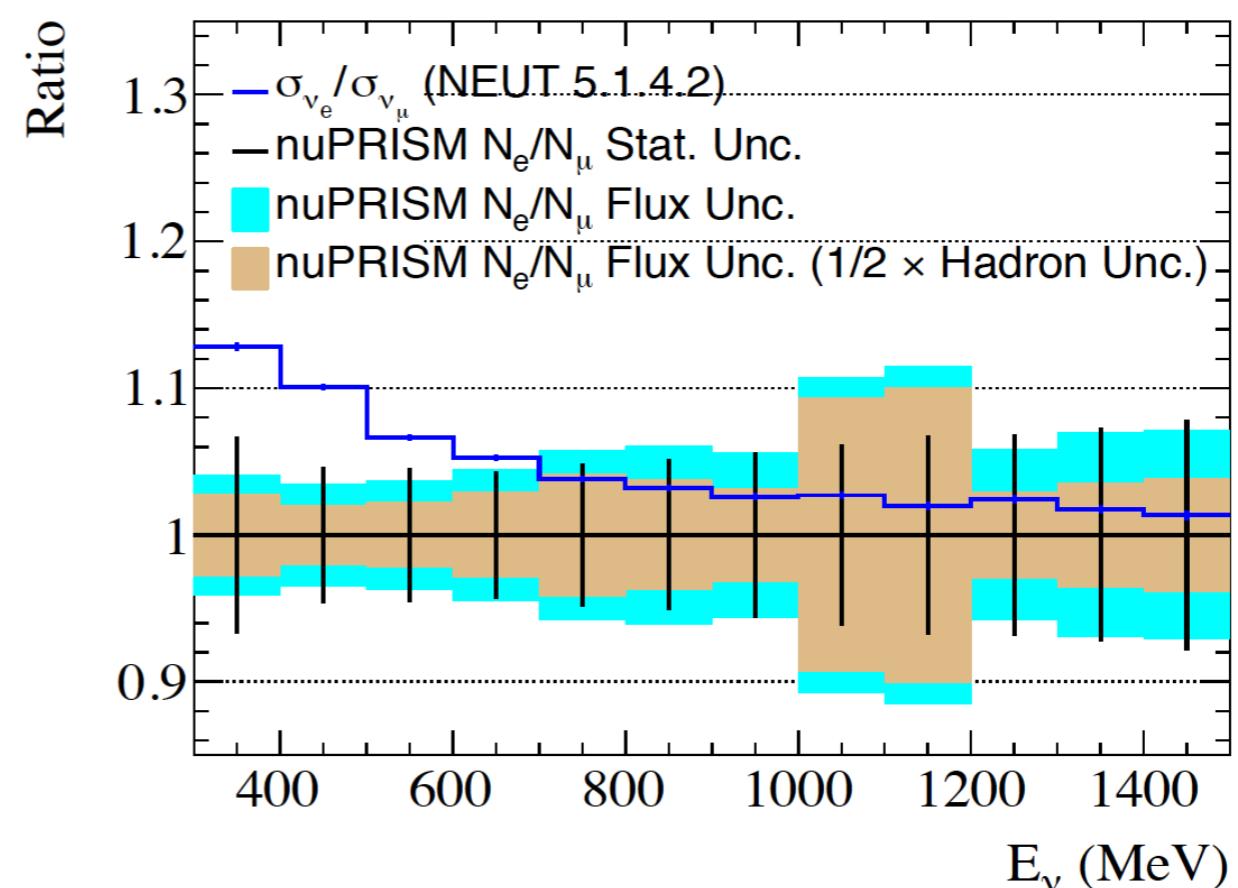
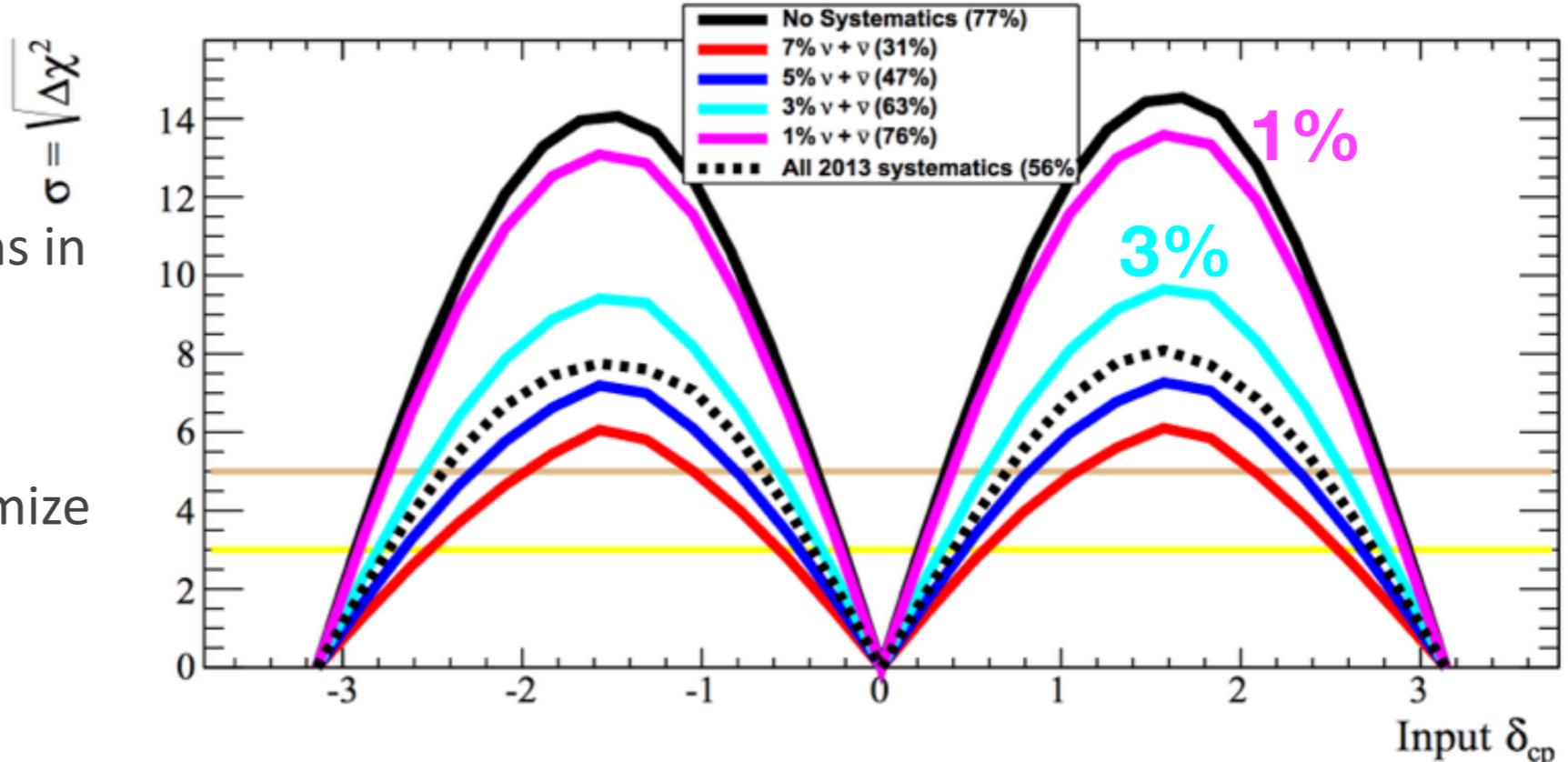
- For systematics to be dangerous they must be able to replicate the effects of shifting  $\delta_{cp}$  in all 4 analysis samples
- Absolute flux normalization and shape
  - Secondary and tertiary hadron production
  - Flux shape differences at the Near and Far detectors
- Uncertainties from cross section models and nuclear initial state models need to be factorized
- A coherent picture of nuclear initial state effects is required
- Cross section flavor differences and rates for exclusive final state channels require theoretical input
- The convolution of flux, cross section, FSI and detector effects in determining energy scale will be difficult to untangle
  - Both FSI and detector effects can be different for  $\nu$  and  $\bar{\nu}$
  - Relative  $\bar{\nu}/\nu$  uncertainties currently provide freedom to mimic  $\delta_{cp}$ -like effects
- Biases in the energy scale from mis-reconstruction and/or poorly modeled/constrained missing energy (neutrons) must be eliminated

# Systematics

## ► HyperK - Mark Hartz

- ▶ Showing impact of cross sections in measurement of CPV
- ▶ Need <1-2% systematic to minimize impact.
- ▶ Achievable by nuSTORM. With nuPRISM concept, 3% level on  $\sigma_{\nu_e}/\sigma_{\nu_\mu}$  is plausible

	Flux Error	Hadron x1/2	Stat. Error
300-600 MeV	3.2%	1.7%	2.9%
600-900 MeV	5.2%	3.4%	2.7%



# Systematics

## ► Alessandro Bravar - Prospects for reduction of beam flux uncertainties

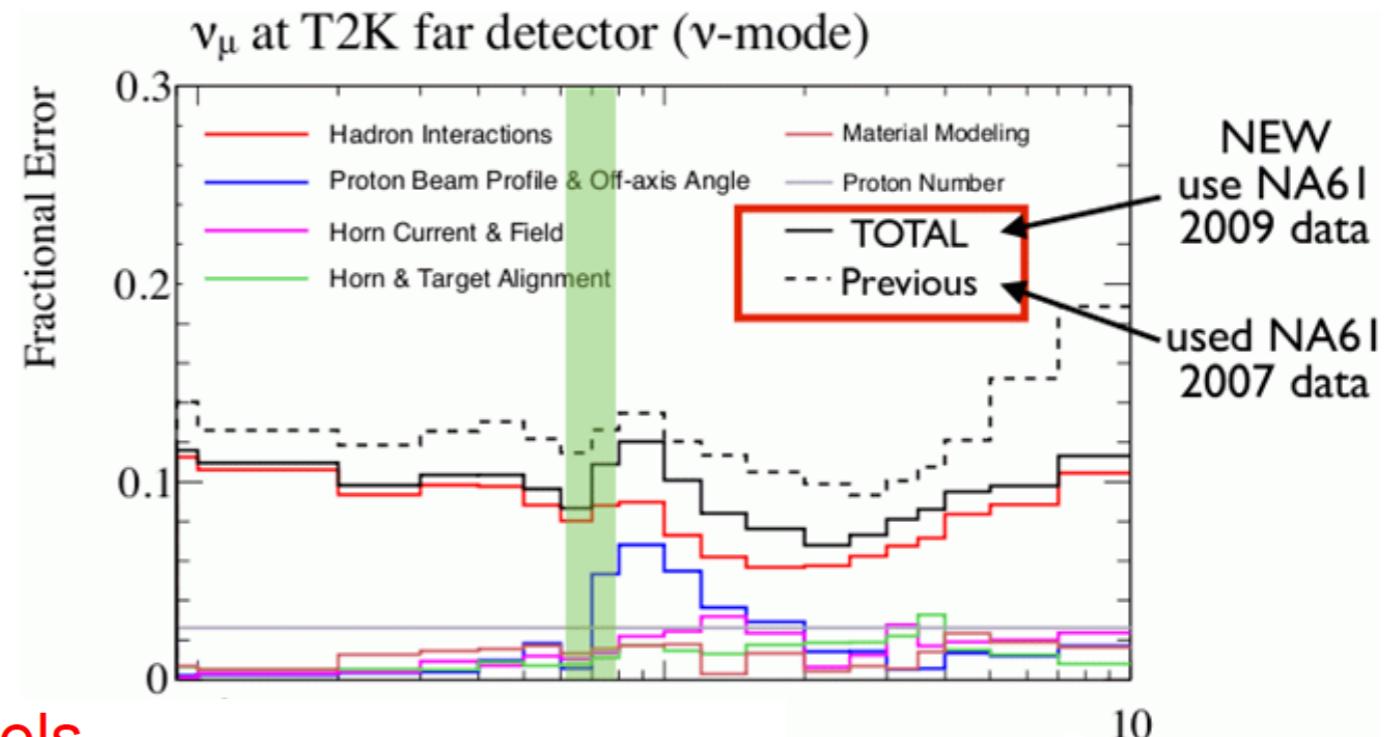
To improve on NA61 results:

increase statistics by a factor of 10

better understanding of interaction and production cross sections

forward acceptance (upgrades under way)

vertexing (replica targets)



None of the existing hadroproduction models describes satisfactorily the ensemble of NA61 data (same for MIPP)!

Use “in situ” measurements to further constrain the flux

neutrino – electron elastic scattering (only “standard candle” in neutrino scattering)  
muon monitors

Expect that uncertainties on neutrino fluxes will decrease down to 5% from the current ~10 % over the next 5+ years

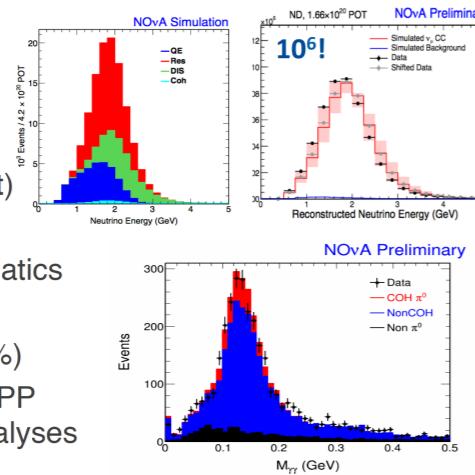


# Systematics

## ► Debbie Harris - Prospects for precision cross section measurements

### Upcoming results from NOvA

- Analyses Underway (J. Paley, NuFact15)
  - $\nu_\mu$  CC
  - $\nu_e + A$
  - $\nu + e$  (flux constraint)
  - Coherent  $\pi^0$
- Dominant  $\nu_e + A$  Systematics
  - Flux (21%)
  - Energy scale (10-15%)
  - Improve both with MIPP and supporting  $\nu_\mu$  analyses



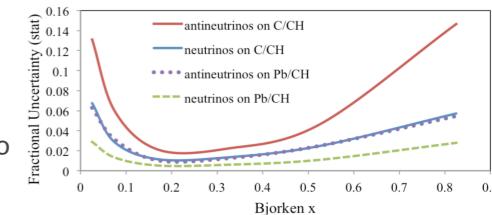
### Upcoming results from T2K

- On Axis
  - QE two-track versus QE-like
    - Multi-nucleon searches
  - QE double differential
  - Charged pion double differential
  - Neutral Pions
  - Coherent charged (neutral?)
  - More  $\nu_e$  cross sections
  - Nuclear Ratios (w/ $H_2O$ , Pb, CH) for model comparisons
  - Pion multiplicities

- On Axis
  - Energy dependent  $\nu_\mu$  CC inclusive on Fe
  - $\nu_\mu$  CC Coherent pion production on C
  - $\nu_\mu$  CC  $0\pi$  differential on C

### Upcoming results from MINERvA

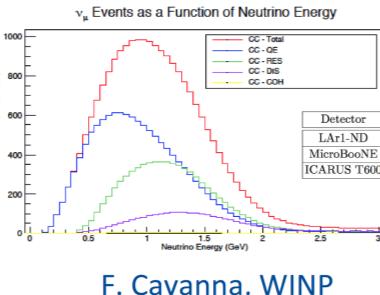
- $\nu_e$  CCQE cross sections (Ghosh)
- Nuclear target ratios for DIS events (Bravar)
- $\nu_\mu$  CCQE double differential (Carneiro)
- CCQE on Fe, Pb, C compared to CH in Low Energy Beam
- Medium Energy results will feature much higher statistics
  - Higher flux and cross section, higher numbers of protons on target collected
  - Exclusive channel ratio results for Fe, Pb, C, compared to CH
  - Chance to look at nuclear effects in DIS at few per cent level!



DH, FNAL PAC 6/15

### Next Step for 1GeV LAr Measurements

- 3 detectors, statistics at 5M!
- Mix of CCQE, Resonance
- Events below for 6.6E20POT
- LAr1-ND: 3M CCQE, >1M pion production events
- All detectors have fine granularity

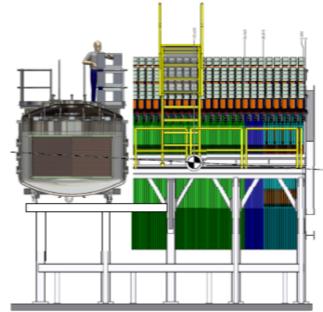


F. Cavanna, WIMP

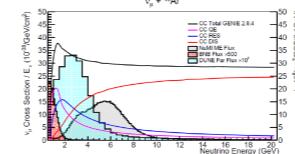
Detector	Distance	Fiducial Mass	CC $\nu_\mu$ events	CC $\nu_e$ events
MicroBooNE	470m	61t	122k	800
LAr1-ND	110m	112t	5.2M	38k
ICARUS	600m	476t	550k	2k

### Next step for 3-6GeV LAr precision: CAPTAIN MINERvA

- Install the CAPTAIN detector in MINERvA to study neutrino-argon interactions in the medium-energy NuMI beam
- CAPTAIN-MINERvA can measure cross section ratios (i.e., argon to carbon)
  - More stringent tests of the models can be performed with ratios due to cancellation of large systematic uncertainties such as the neutrino flux
- Stage I approval from Fermilab Director

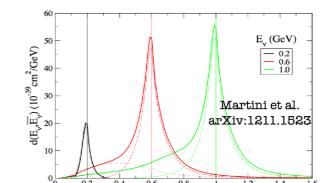
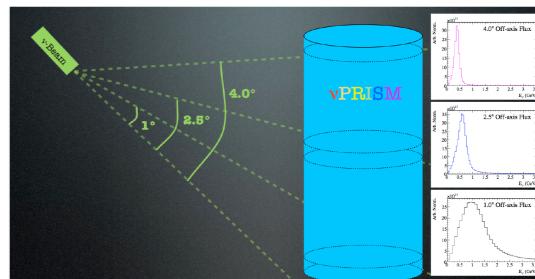


@6x1020 POT	Events with reco. $\mu$	Events with reco. $\mu$ + chg
CCQE-like	900k	800k
CC $1\pi^\pm$	2000k	1000k
CC $1\pi^0$	1600k	600k



### Getting to a MonoChromatic Neutrino Beams

- NUPRISM:
  - Take advantage of 2-body decay kinematics to "create" monochromatic energy beams
  - Not for the faint of heart, will need a lot of statistics to do subtractions from different locations
  - Best chance at directly quantifying this picture
- What about other targets?

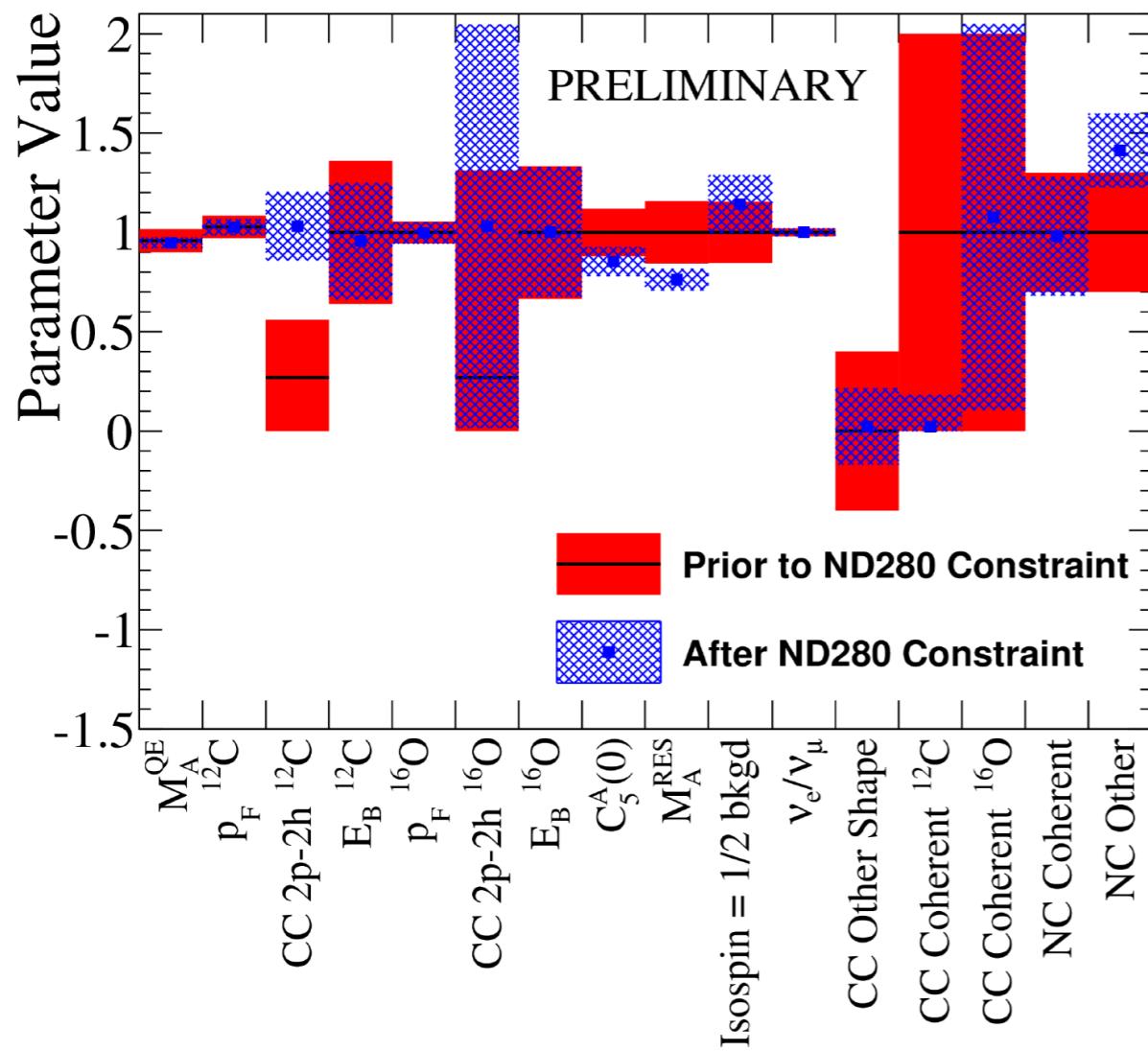


Fermilab

- We need to harvest the impressive data that is already recorded
- Need to get good statistics in anti-neutrino mode and harvest that data as well!

# Systematics

## ► ND Session - T2K - Kendall Mahn



Some cross section parameters (2p2h on C,  $M_A^{RES}$ ) changed significantly compared to external data prior

$\bar{\nu}_\mu$ disappearance analysis		w/o ND measurement	w/ ND measurement
v flux and cross section	flux	7.1%	3.5 %
	cross section cmn to ND280	5.8%	1.4 %
	(flux) $\times$ (cross section cmn to ND280)	9.2%	3.4 %
	cross section (SK only, include ↓)	10.0 %	
	multi-nucleon effect on oxygen	9.5%	
	total	13.0%	10.1%
Final or Secondary Hadronic Interaction			2.1%
Super-K detector			3.8%
total			14.4% 11.6%

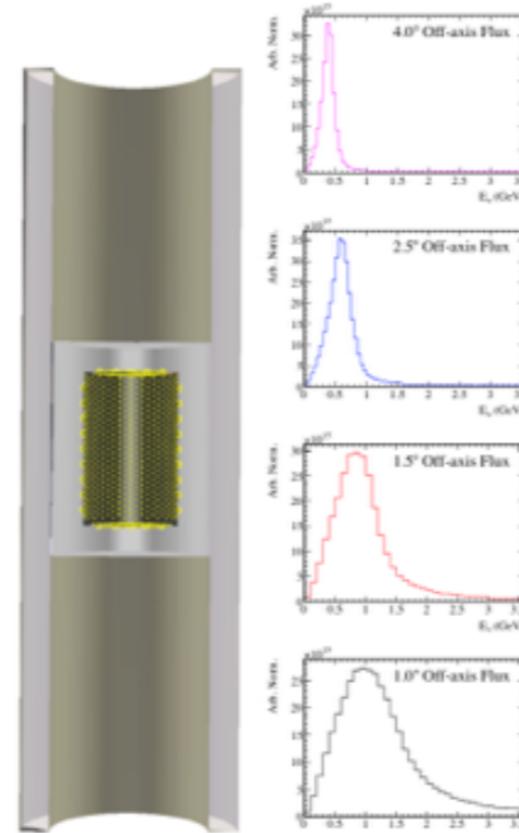
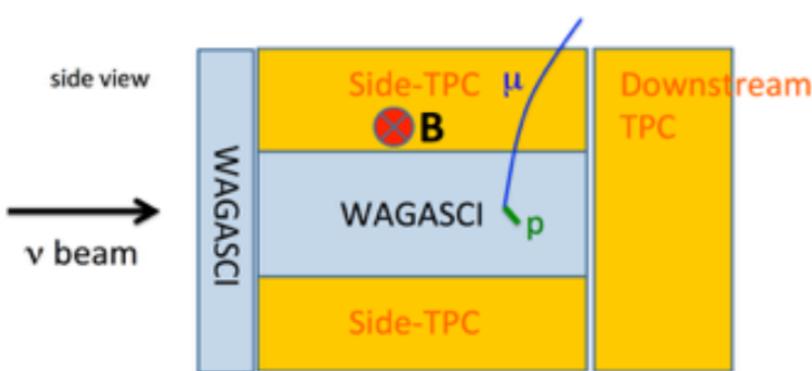
Fractional error on number-of-event prediction

- ND impact on T2K's analysis
  - Tuning of cross sections
  - Large reduction in flux uncertainties

# Systematics

## ► ND Session - HyperK - Akihiro Minamino; DUNE - Kirk McDonald NuPRISM

### WAGASCI in ND280 magnet

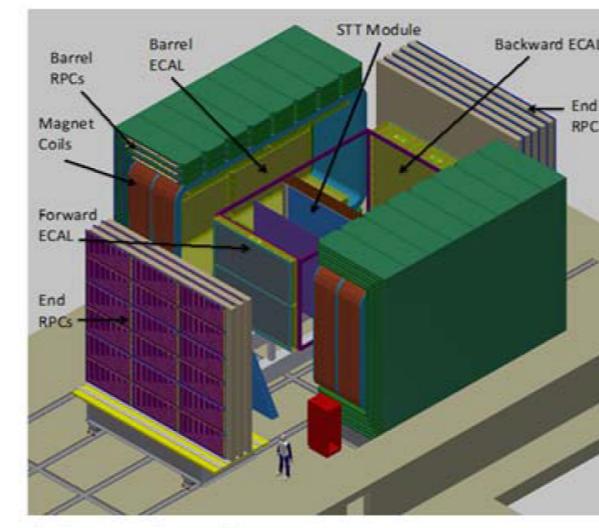


### The NOMAD-inspired Fine-Grained Tracker (FGT)

- It consists of:**



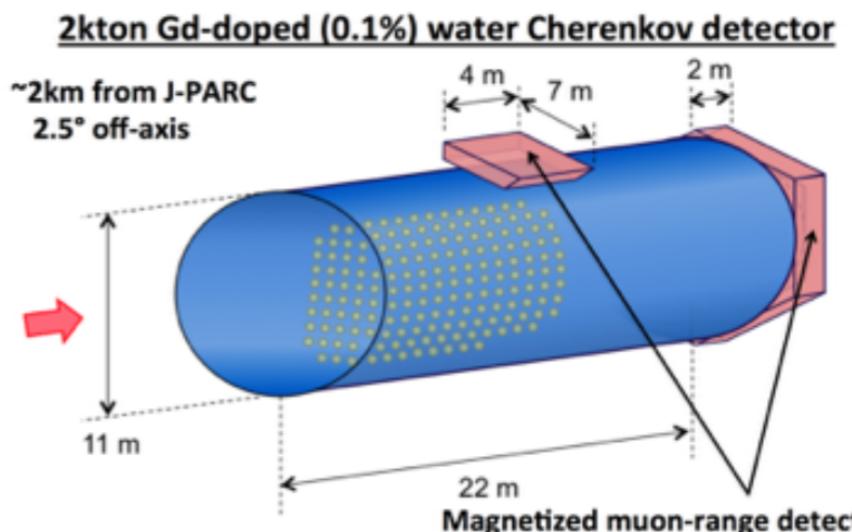
- Central straw-tube tracking system
- Lead-scintillator sampling ECAL
- Large-bore warm dipole magnet
- RPC-based muon tracking systems



- It provides:**

- Constraints on cross sections and the neutrino flux
- A rich self-contained non-oscillation neutrino physics program

### TITUS



	$\nu_e$ cross section	H <sub>2</sub> O target	4π accep.	Wrong sign BG	NC, Int. $\nu_e$ BG	Muon FS vs. $\nu$	Hadronic FS	# of neutron (Gd)	CC $\pi^0$
Current ND280	Yellow	Yellow	Yellow	Green	Yellow	Yellow	Yellow	Red	Yellow
ND280 (WAGASCI)	Yellow	Green	Green	Green	Yellow	Yellow	Yellow	Red	Yellow
ND280 (HP-TPC)	Red	Red	Green	Green	Yellow	Yellow	Green	Green	Red
ND280 (WbLS)	Yellow	Red	Yellow	Green	Yellow	Yellow	Yellow	Red	Yellow
ND280 (Emulsion)	Yellow	Green	Green	Green	Yellow	Yellow	Green	Red	Yellow
νPRISM	Green	Green	Yellow	Green	Green	Red	Red	Yellow	Yellow
TITUS	Green	Green	Green	Green	Yellow	Red	Red	Green	Green

= Good

= OK

= Not Good <sup>26</sup>

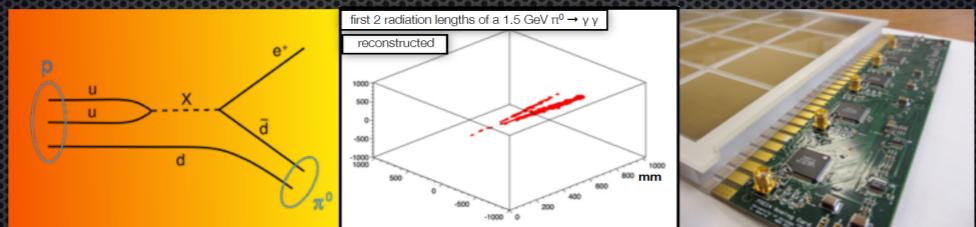
Quantifying the benefits of augmenting the ref. design with a LArTPC or high-pressure gaseous argon TPC

# Systematics

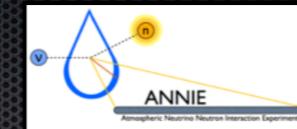
## ► ND Session - ANNIE - Mayly Sanchez; CAPTAIN and LArIAT - Jason

### The ANNIE experiment

- ANNIE, seeks to measure the abundance of final state neutrons from neutrino interactions in water, as a function of energy (see arXiv:1409.5864 and arXiv:1504.01480).
- It is also the first application in a HEP experiment of LAPPDs (Large-Area Picosecond Photo-Detectors).



- ▶ A key physics measurement in understanding the nature neutrino-nucleus interactions.
- ▶ Application of a promising technology for detecting neutrinos.



### MiniCAPTAIN

1m Ø LArTPC in neutron beam at Weapons Neutron Research facility



#### Physics goals:

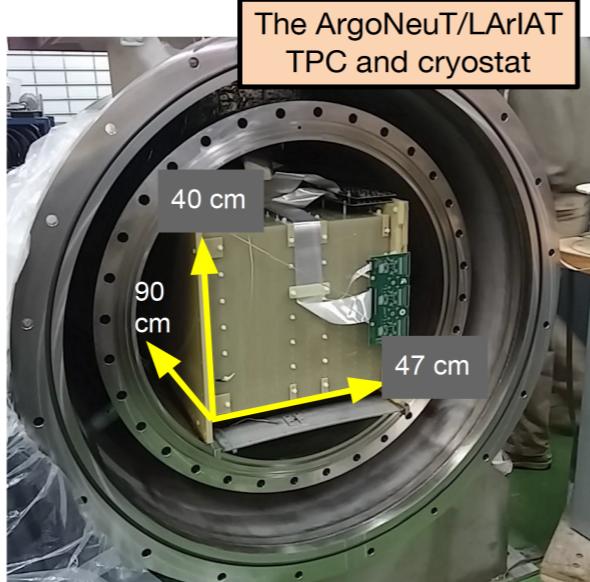
- Ar\* nuclear de-excitations
- Neutron scatters at known E<sub>n</sub>
- Neutron-induced π<sup>±</sup> production

### LArIAT

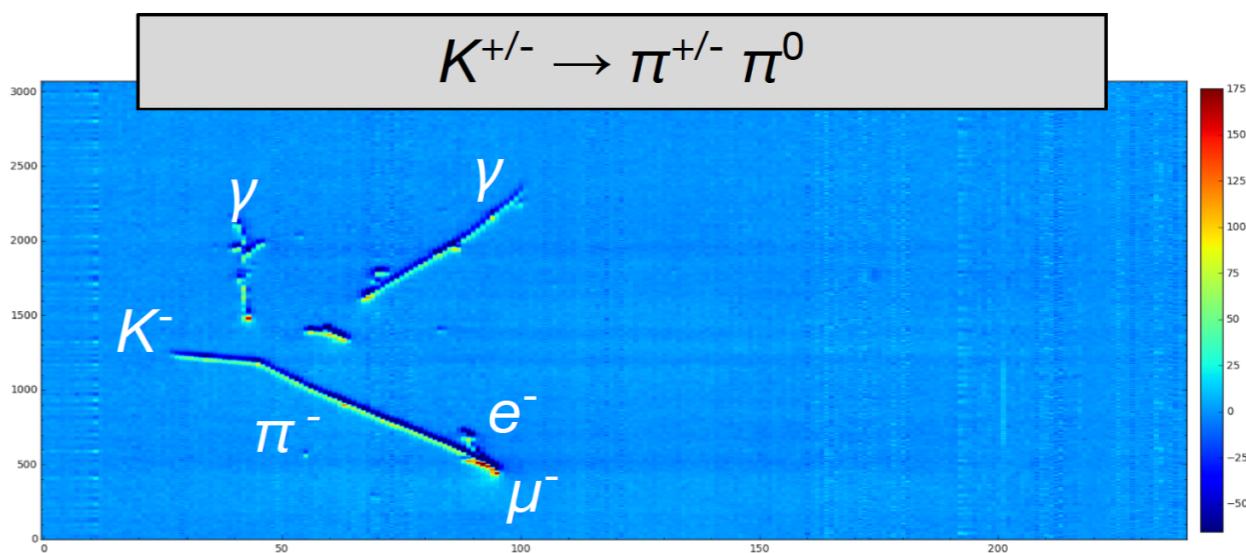
“Table-top” (170L) LArTPC in a test beam at Fermilab Test Beam Facility

- Repurposed ArgoNeuT detector
- **Physics goals:**
  - π-Ar interactions
  - e/γ shower ID
  - μ-Ar capture
  - non-magnetic charge determination
  - kaon studies
  - Geant4 validation
- **R&D goals:**

Optimize PID algorithm, calorimetry with charge & light, and 2D/3D event reconstruction



- ▶ Taking data since April 30, 2015



# Answer to Q3

***Question from NuFact '14: What is the target for the systematic error budget of next generation facilities? What do we need to reach this level? How much improvement in constraining flux uncertainties can we expect from dedicated hadron production efforts?***

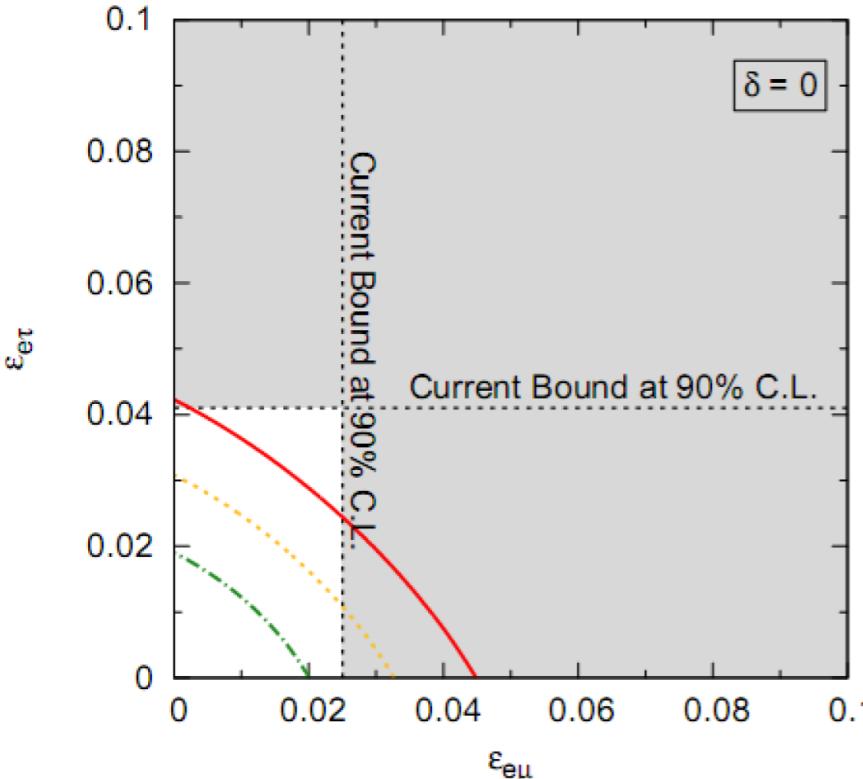
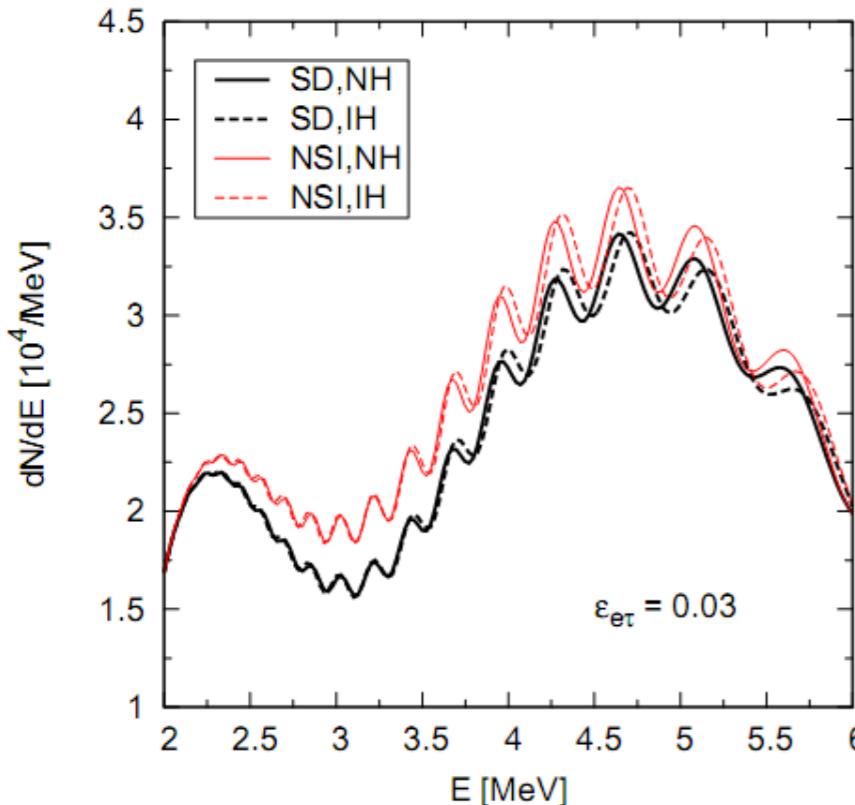
- ▶ Absolute beam flux uncertainties can be reduced from 10% to 5% in next 5 years
- ▶ <1%-2% precision on cross sections required for ultimate sensitivity to CPV
- ▶ nuSTORM would be ideal, but have comprehensive program of cross section measurements in H<sub>2</sub>O and Ar throughout the World over the next decade
- ▶ Multi-purpose Near Detectors essential in reducing flux and cross section uncertainties.
- ▶ To assess cross section systematics in more focused way, Hugh Gallagher proposed creation of mock data challenge samples by generator groups in consultation with experiments and theorists.

# Q4: New Physics and v CPV

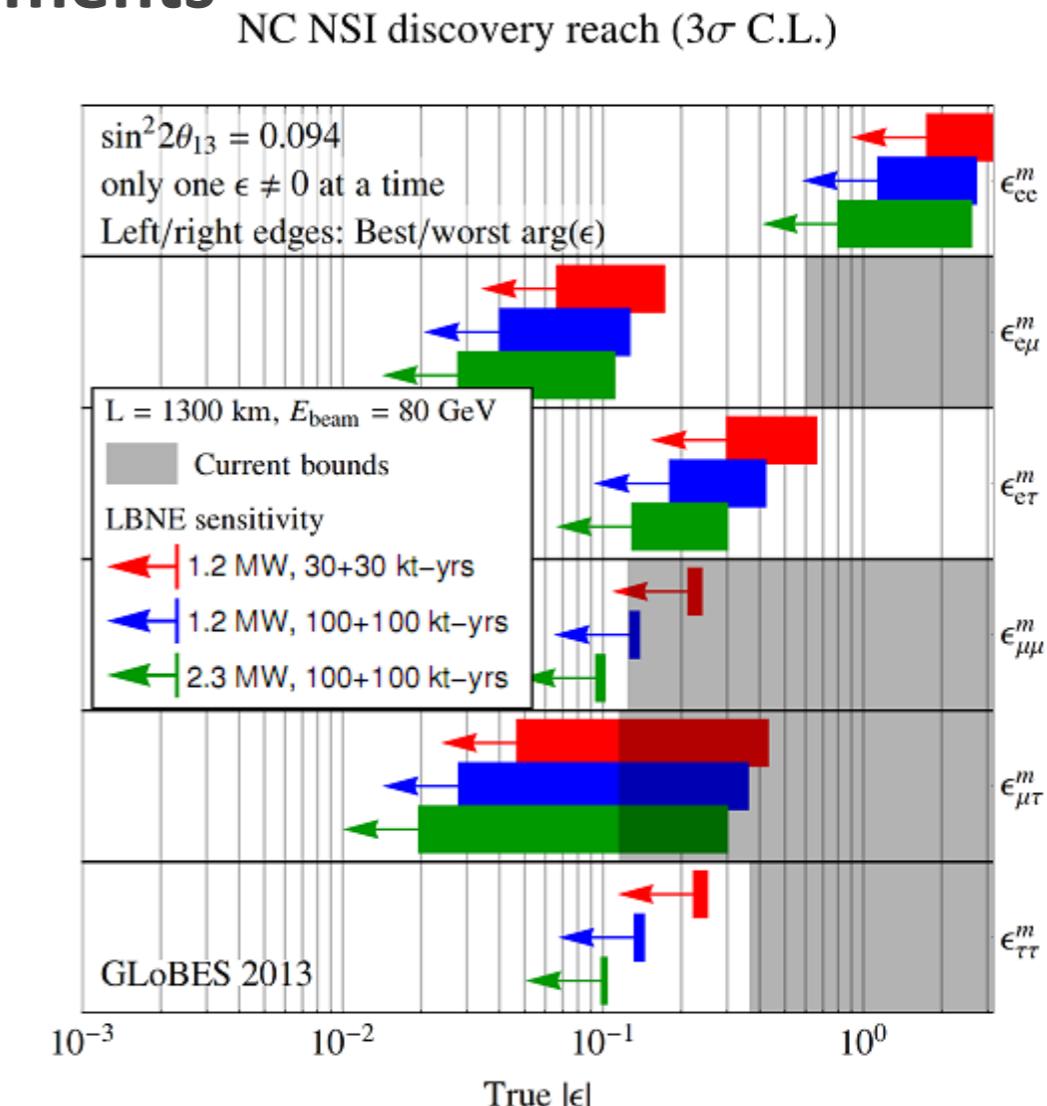
**Question from NuFact '14:** Do the current bounds on new physics in the neutrino sector (NSI, non-unitarity, steriles...) allow for effects large enough to interfere with CPV searches? Which experimental setups can improve these bounds?

# New Physics and $\nu$ CPV

## ► Sushant Raut - Effect of NSIs on CPV measurements



Matter NSI  
bounds:  
DUNE



Parameter	ESSvSB limits	Existing limits
$\epsilon_{\mu e}^s$	0.014	0.026
$\epsilon_{\mu e}^d$	0.014	0.025

# New Physics and $\nu$ CPV

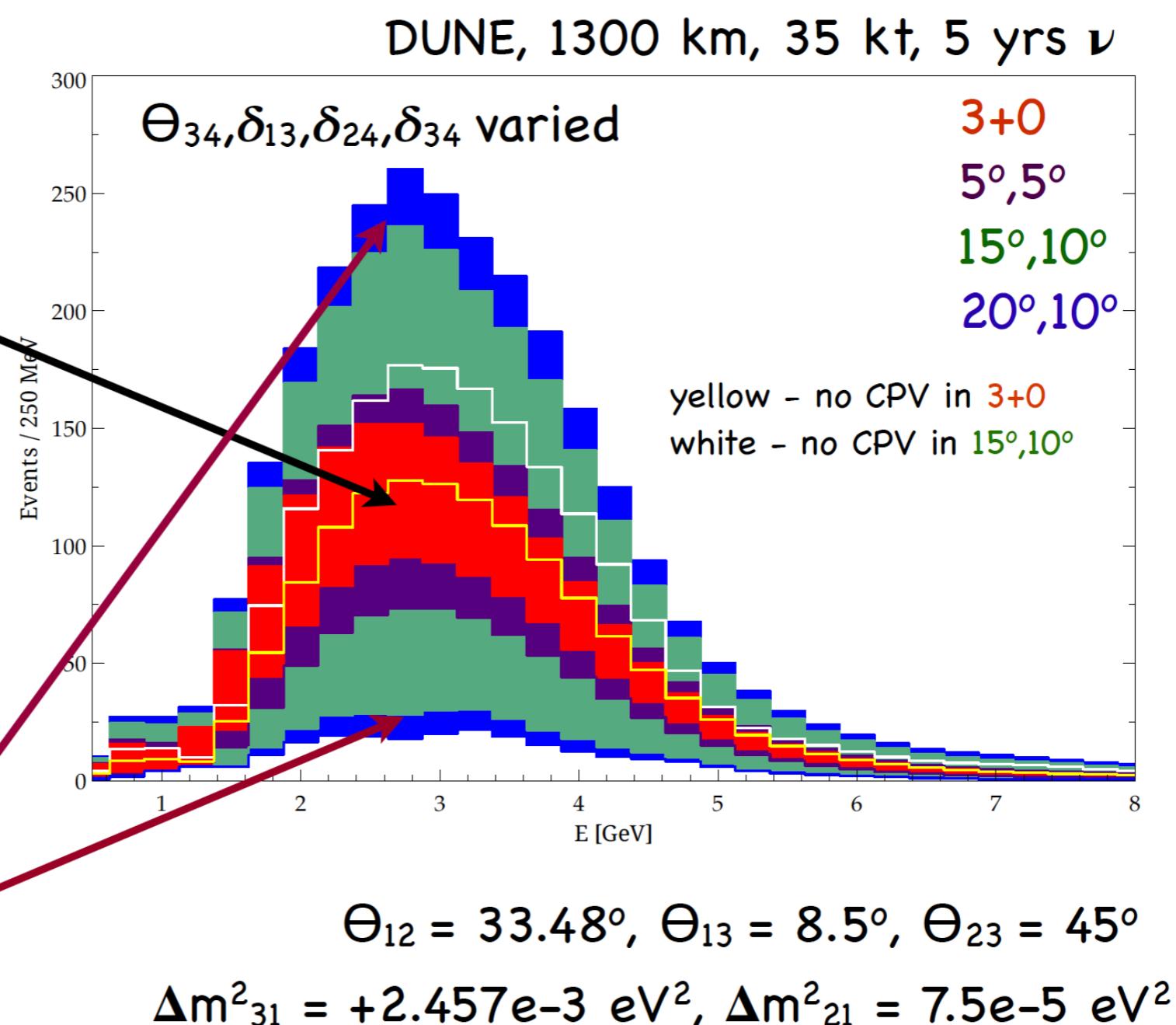
## ► Suprabh Prakash - Effect of sterile neutrinos on CPV measurements

We consider a  $O(1 \text{ eV}^2)$  mass sterile neutrino, heavier than the other 3 mass eigenstates  $\Rightarrow \Delta m^2_{41} = +1 \text{ eV}^2$ .  $\Delta m^2_{31}$  can be + or -

In producing these plots, we have averaged over the  $\Delta m^2_{4i}$  induced oscillations.

The 3+1 band can potentially encompass the 3+0 band, leading to substantial degeneracy.

For large active-sterile mixings, an excess or shortage of events, esp. at osc. max. will be pointers to the existence of new physics.



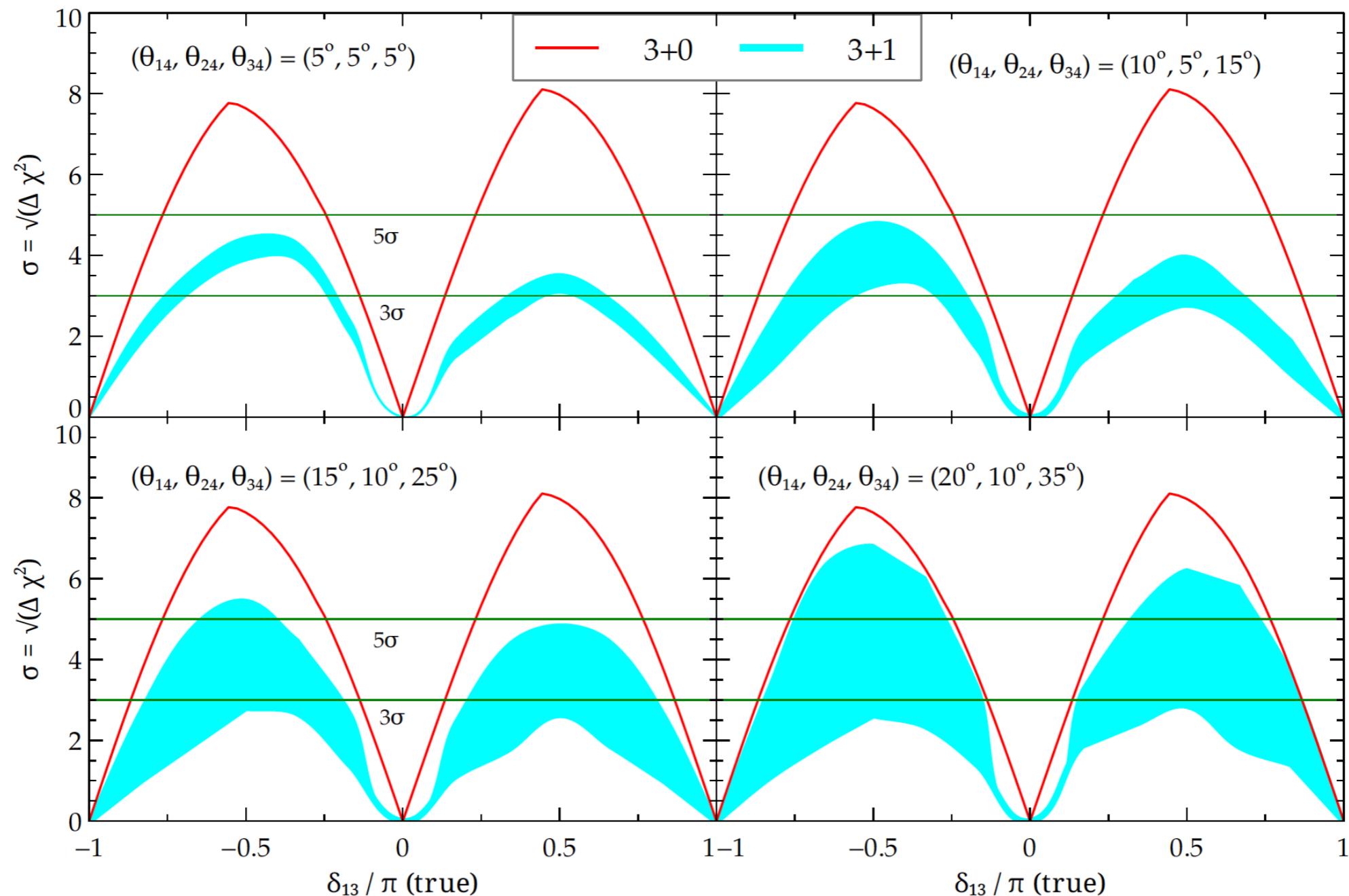
# Q4: New Physics and ν CPV

## ► Suprabh Prakash - Effect of sterile neutrinos on CPV measurements

Cyan band - variation of true  $\delta_{24}$  and  $\delta_{34}$

DUNE, 1300 km, 35 kt, 5+5

1% signal norm. error & 5% back norm. error



$$\Theta_{12} = 33.48^\circ, \Theta_{13} = 8.5^\circ, \Theta_{23} = 45^\circ$$

$$\Delta m^2_{31} = +2.457 \text{e-}3 \text{ eV}^2, \Delta m^2_{21} = 7.5 \text{e-}5 \text{ eV}^2$$

# Answer to Q4

**Question from NuFACT '14:** Do the current bounds on new physics in the neutrino sector (NSI, non-unitarity, steriles...) allow for effects large enough to interfere with CPV searches? Which experimental setups can improve these bounds?

- ▶ Answer is yes
- ▶ However, more work is required to understand how other parameter measurements may be affected, and how effects will change with more sterile neutrinos and active-sterile mixing driven by different mass scales
- ▶ We did not discuss specific experimental setups that could potentially disentangle these effects, but should be kept in mind for future NuFacts.

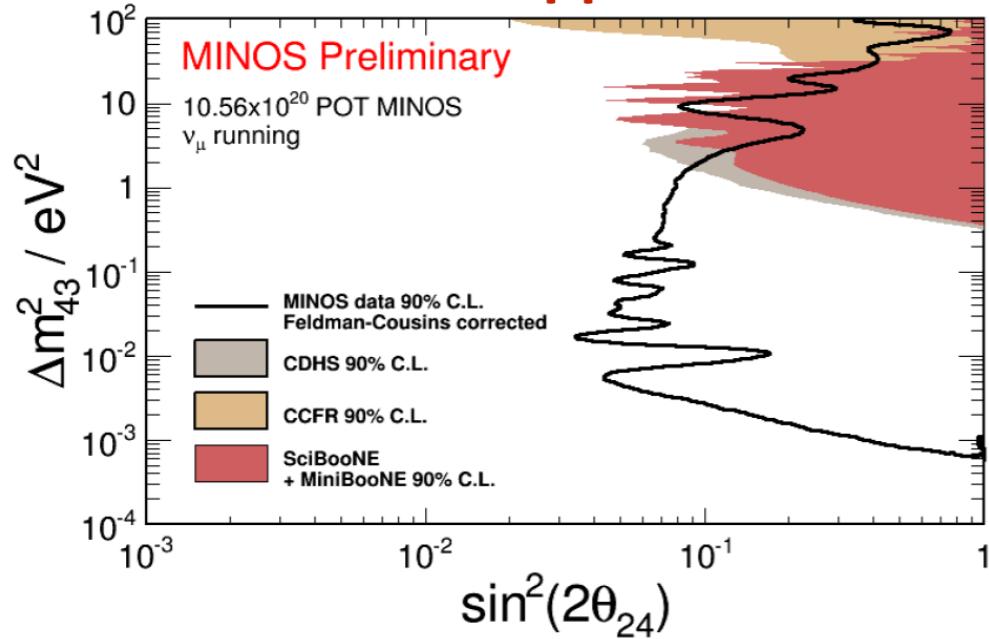
# Q5: Sterile Neutrinos

**Question from NuFact '14:** What is the best strategy to fully probe the LSND anomaly? And the reactor/gallium anomaly?

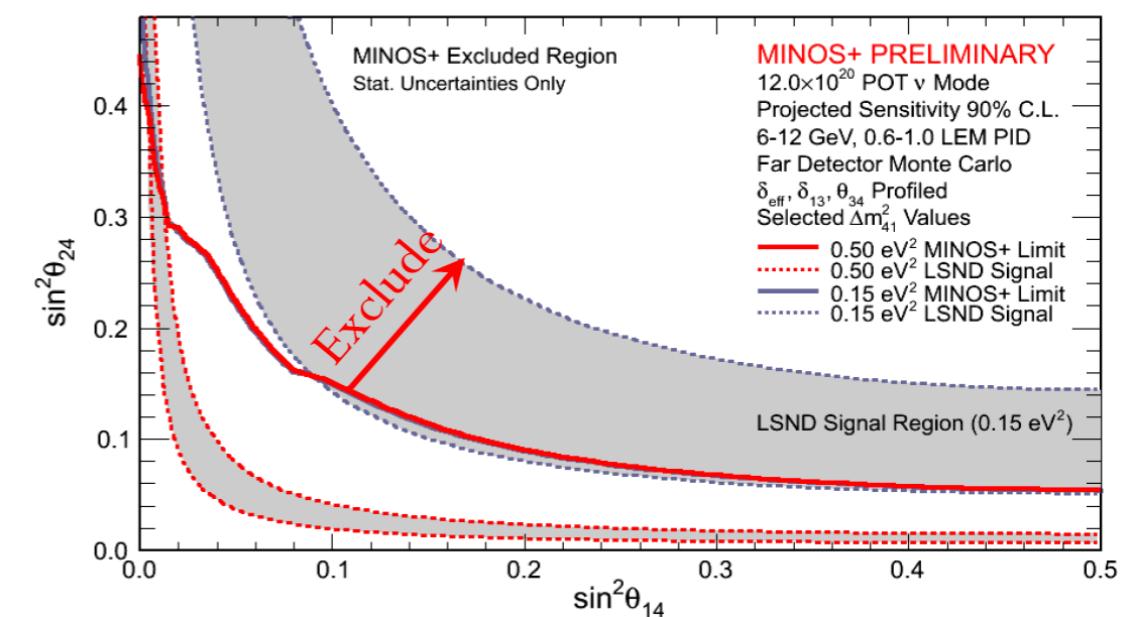
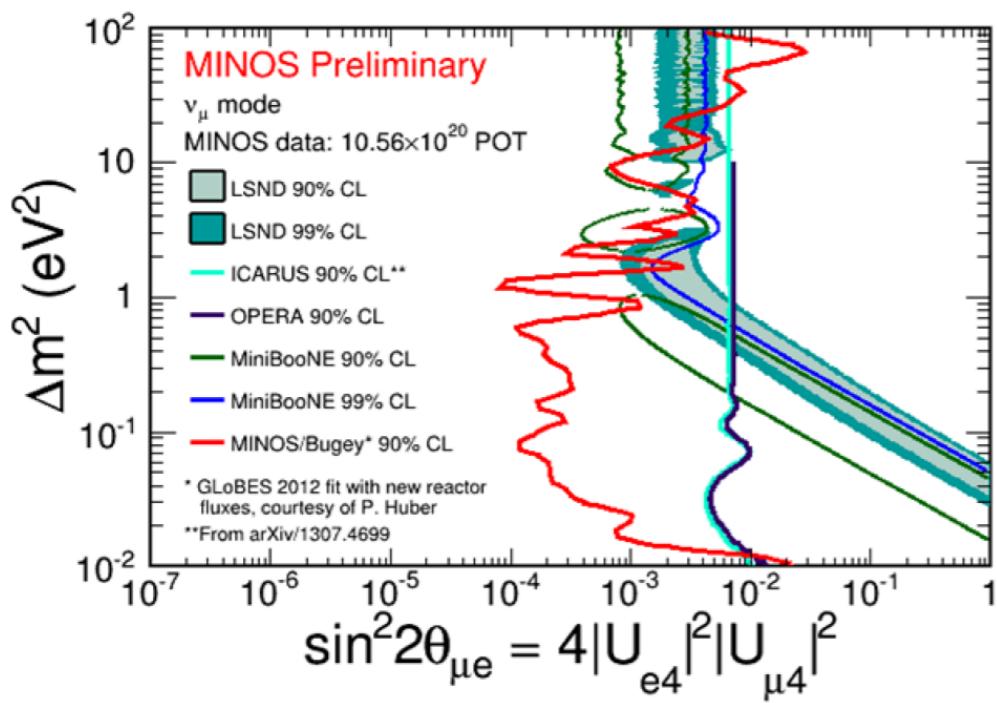
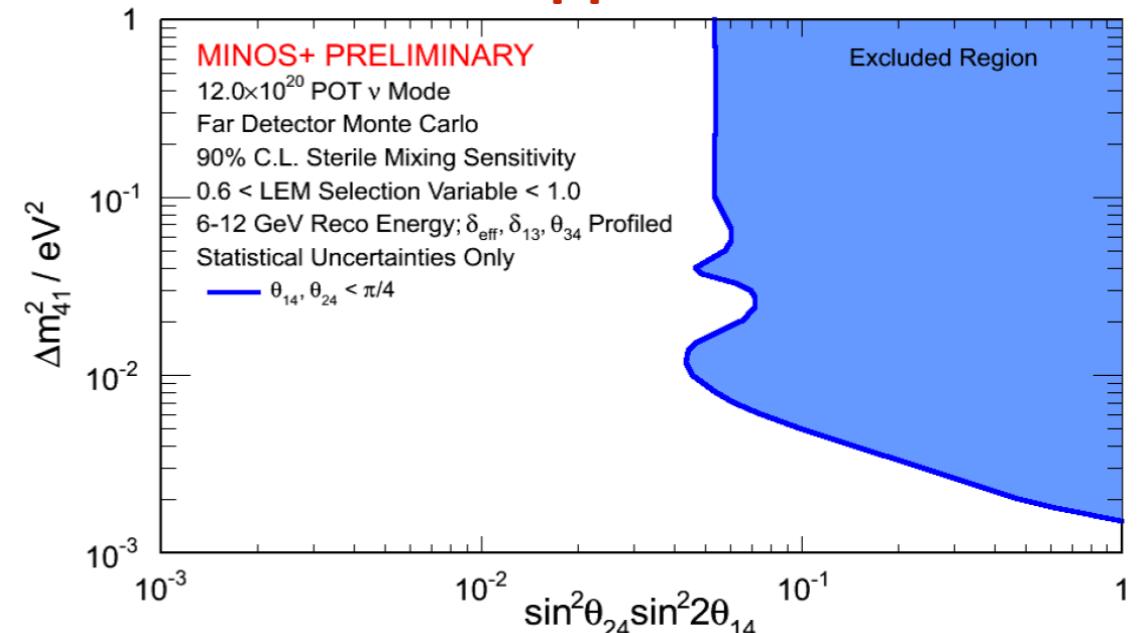
# Sterile Neutrinos

## ► MINOS/MINOS+ - João Coelho

### NuMu Disappearance



### NuE Appearance



- Excluded large unexplored region at low values of  $\Delta m_{43}^2$

Combined MINOS & Bugey data exclude most of the region allowed by LSND & MiniBooNE

- Background dominated, but 2 $\sigma$  significance for appearance

- Difficult measurement due to interference of 3-flavor and 4-flavor oscillation terms

# Sterile Neutrinos

## ► SuperK - Jun Kameda

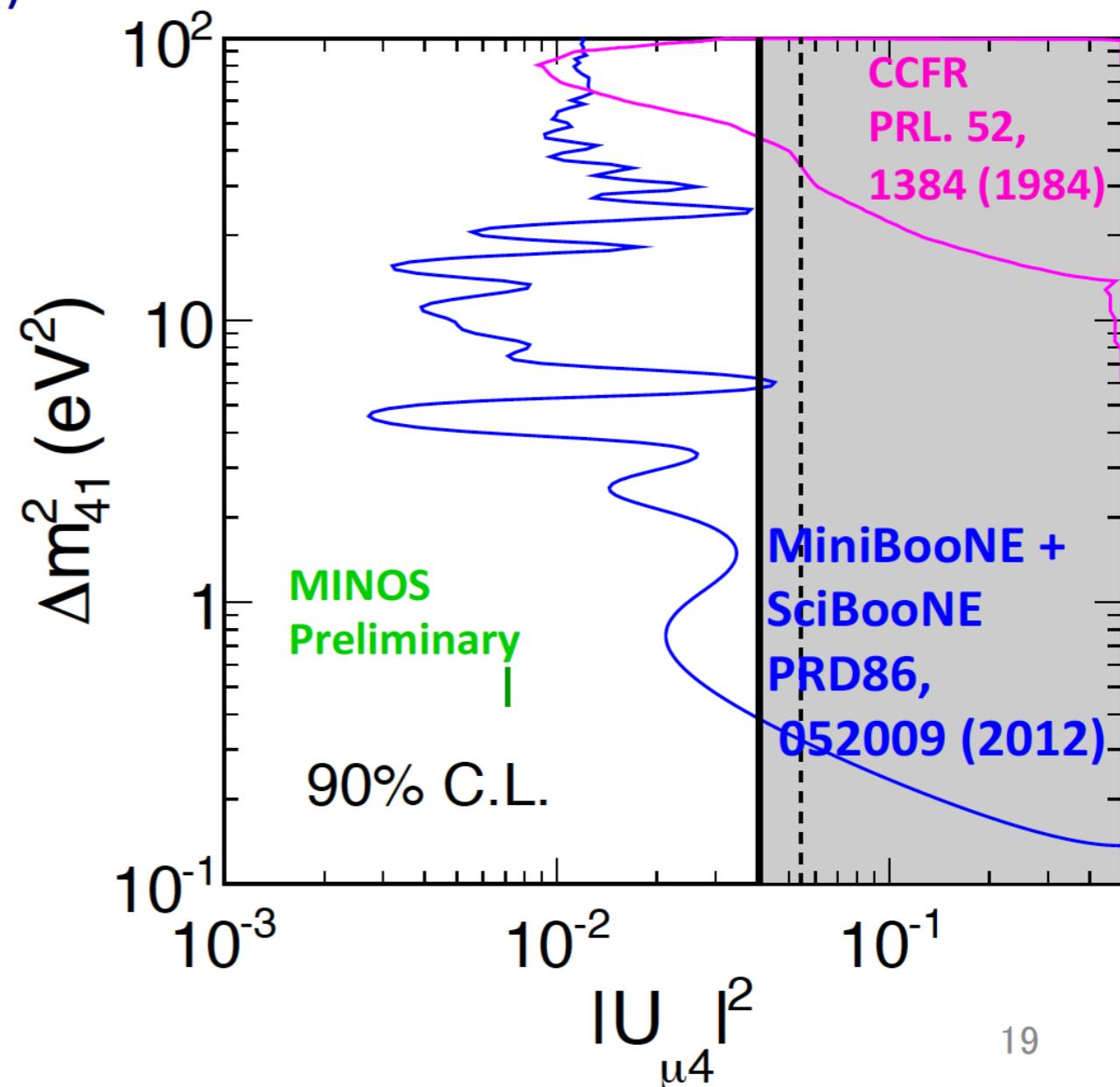
- (Standard neutrinos + sterile neutrinos) is tested for large  $\Delta m_{41}^2$  region ( $>0.01\text{eV}^2$ ) .

$$U = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} & U_{e4} & \dots \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} & U_{\mu 4} & \dots \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} & U_{\tau 4} & \dots \\ U_{s1} & U_{s2} & U_{s3} & U_{s4} & \dots \\ \vdots & \vdots & \vdots & \vdots & \ddots \end{pmatrix}$$

PMNS      Sterile

As with similar experiments, no strong sterile-driven  $\nu_\mu$  disappearance

$|U_{\mu 4}|^2 < 0.041$  at 90% C.L.



# Sterile Neutrinos

► OPERA - Chiara Sirignano

$$\nu_\mu \rightarrow \nu_\tau$$

$$P_{\nu_\mu \rightarrow \nu_\tau} = 4|U_{\mu 3}|^2 |U_{\tau 3}|^2 \sin^2 \frac{\Delta_{31}}{2} + 4|U_{\mu 4}|^2 |U_{\tau 4}|^2 \sin^2 \frac{\Delta_{41}}{2}$$

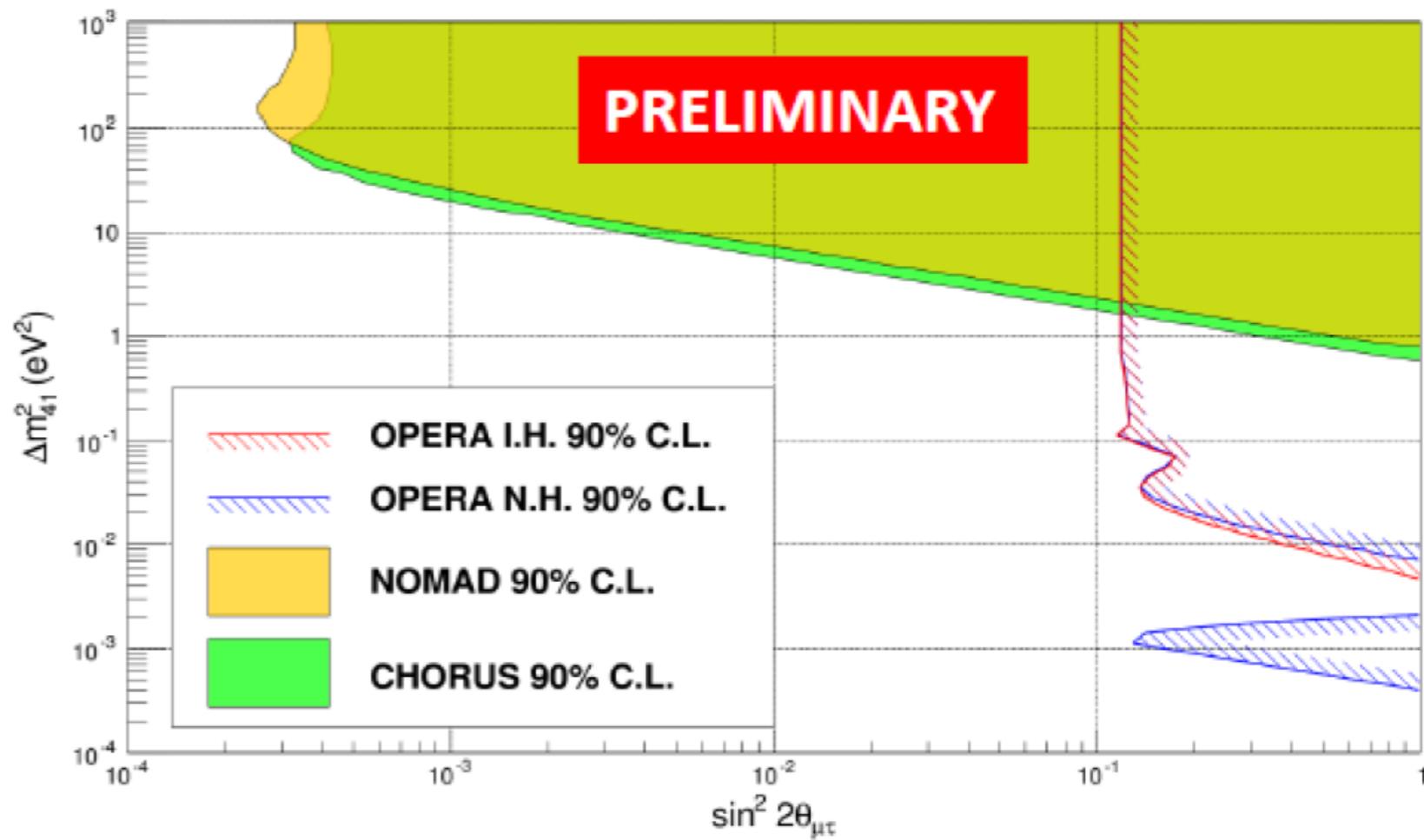
$$+ 2\Re[U_{\mu 4}^* U_{\tau 4} U_{\mu 3} U_{\tau 3}] \sin \Delta_{31} \sin \Delta_{41}$$

$$- 4\Im[U_{\mu 4}^* U_{\tau 4} U_{\mu 3} U_{\tau 3}] \sin^2 \frac{\Delta_{31}}{2} \sin \Delta_{41}$$

$$+ 8\Re[U_{\mu 4}^* U_{\tau 4} U_{\mu 3} U_{\tau 3}^*] \sin^2 \frac{\Delta_{31}}{2} \sin^2 \frac{\Delta_{41}}{2}$$

$$+ 4\Im[U_{\mu 4}^* U_{\tau 4} U_{\mu 3} U_{\tau 3}^*] \sin \Delta_{31} \sin^2 \frac{\Delta_{41}}{2}$$

interference terms



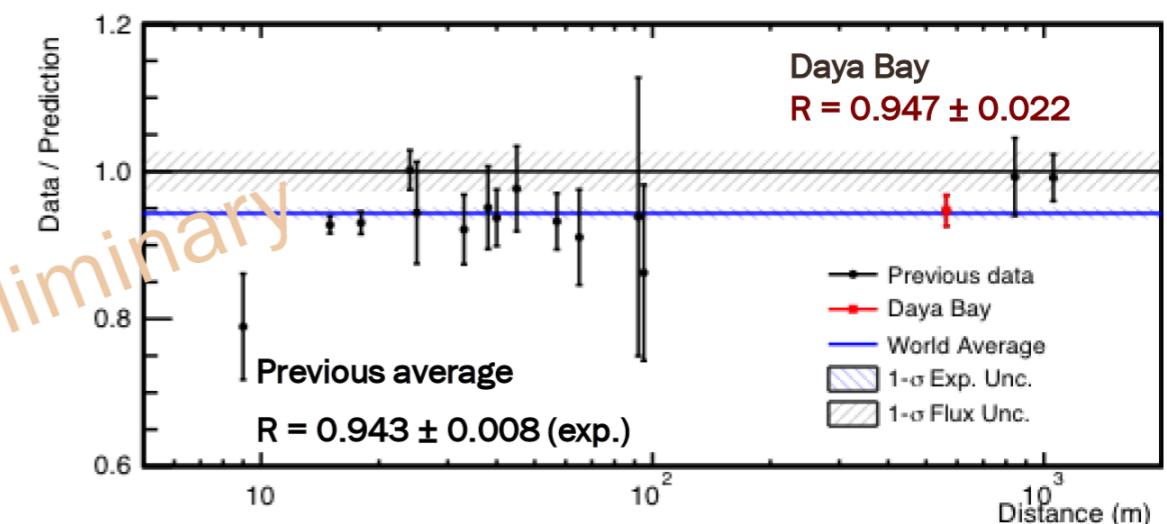
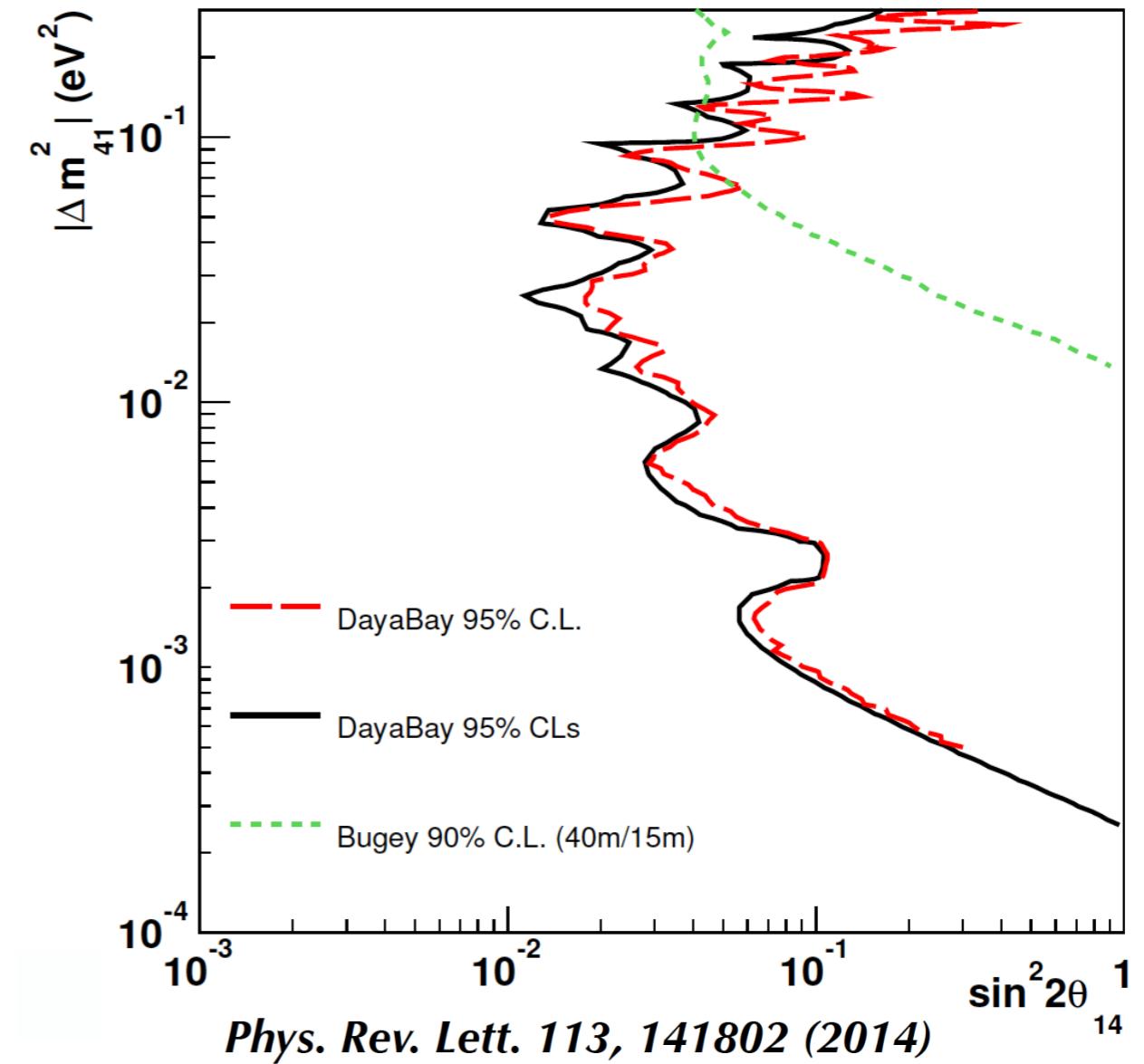
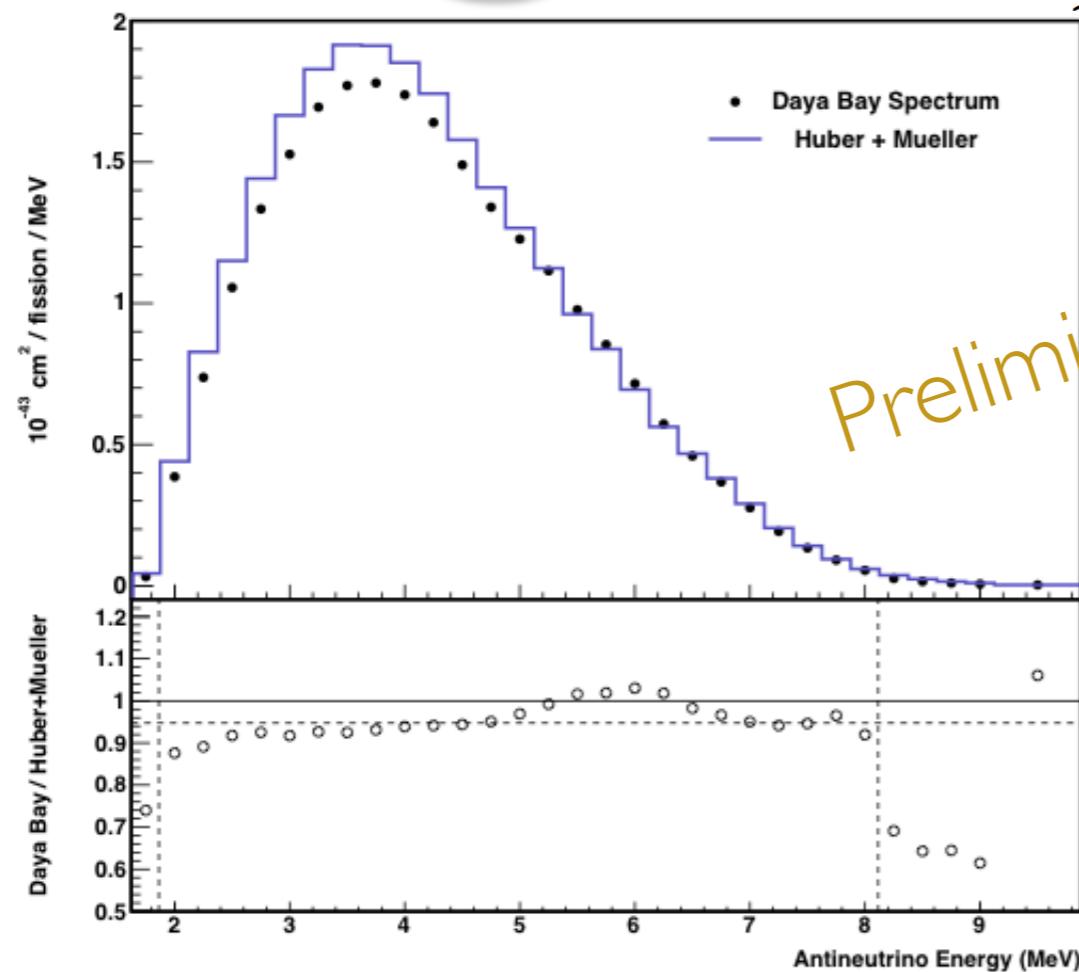
- Showing analysis updated to include 5  $\nu_\tau$  events
- 4 event results published in [\[JHEP 074 \(2015\) 0315\]](#)

First limits on  $|U_{\mu 4}|^2 |U_{\tau 4}|^2$  from direct measurement of  $\nu_\tau$

# Sterile Neutrinos

## ► Daya Bay - Wei Wang

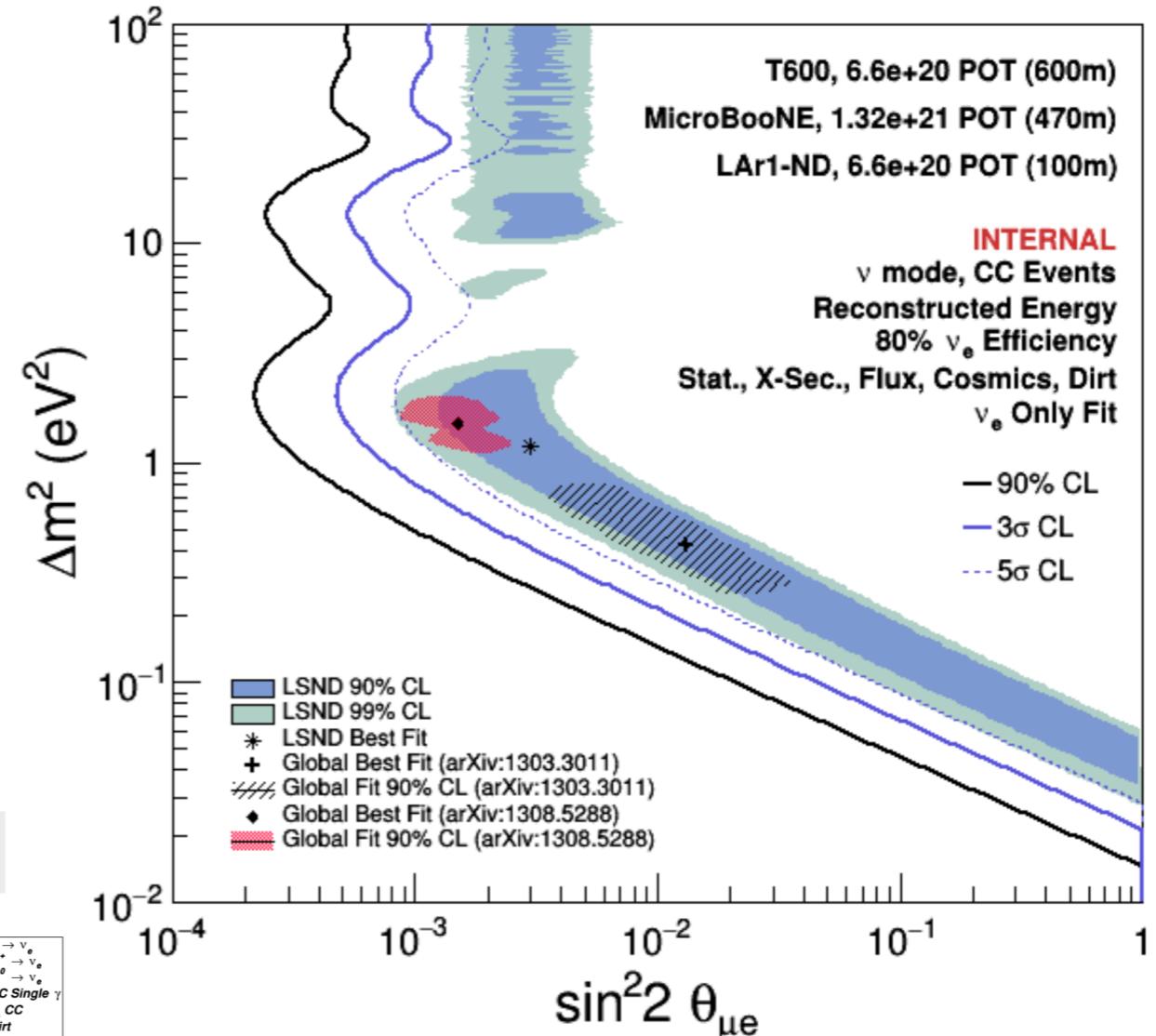
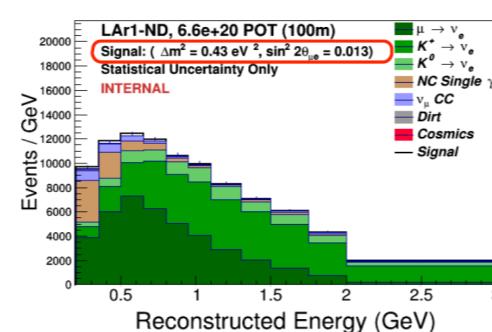
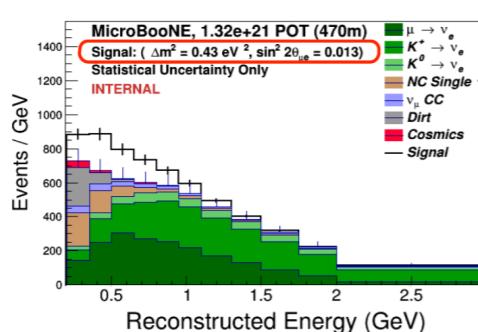
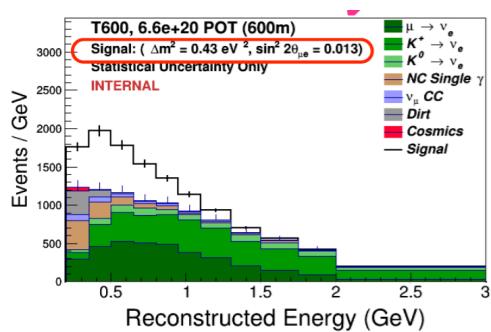
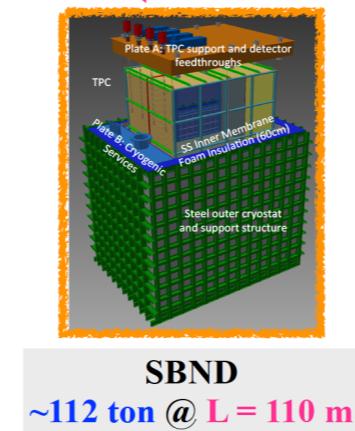
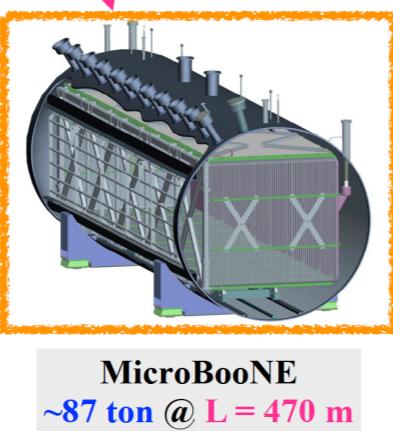
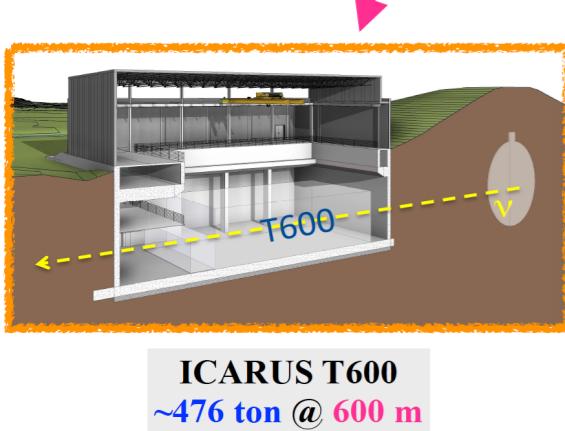
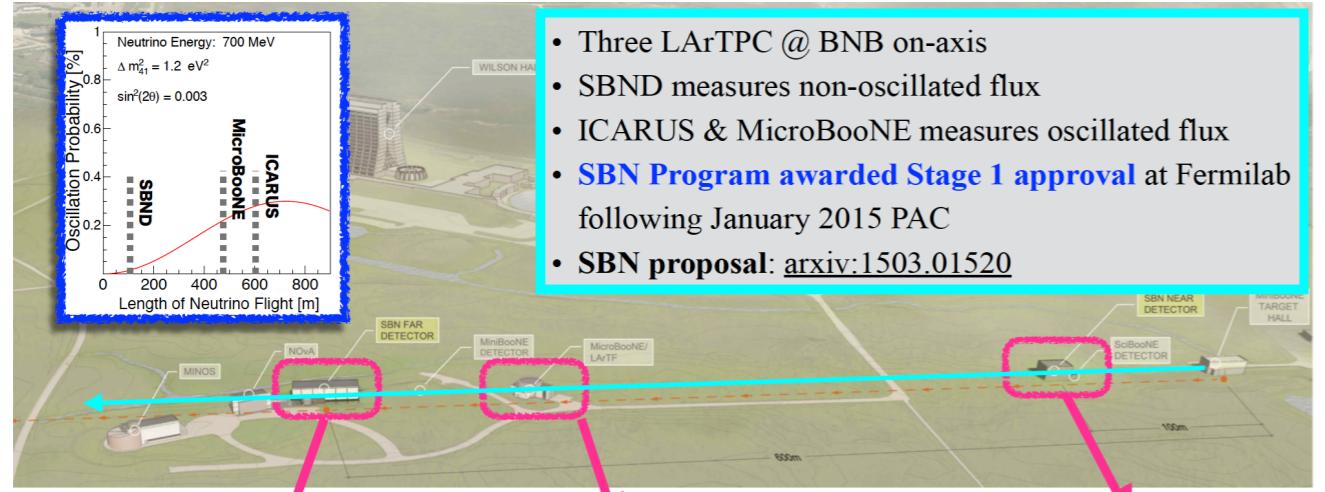
$$P(\bar{\nu}_e \rightarrow \bar{\nu}_e) \simeq 1 - \cos^4 \theta_{14} \sin^2 2\theta_{13} \sin^2 \left( \frac{\Delta m_{ee}^2 L}{4E_\nu} \right) - \sin^2 2\theta_{14} \sin^2 \left( \frac{\Delta m_{41}^2 L}{4E_\nu} \right)$$



- Daya Bay agrees with reactor flux average from previous experiments
- Have produced unfolded flux spectrum to be used for flux and spectrum predictions

# Sterile Neutrinos - Future

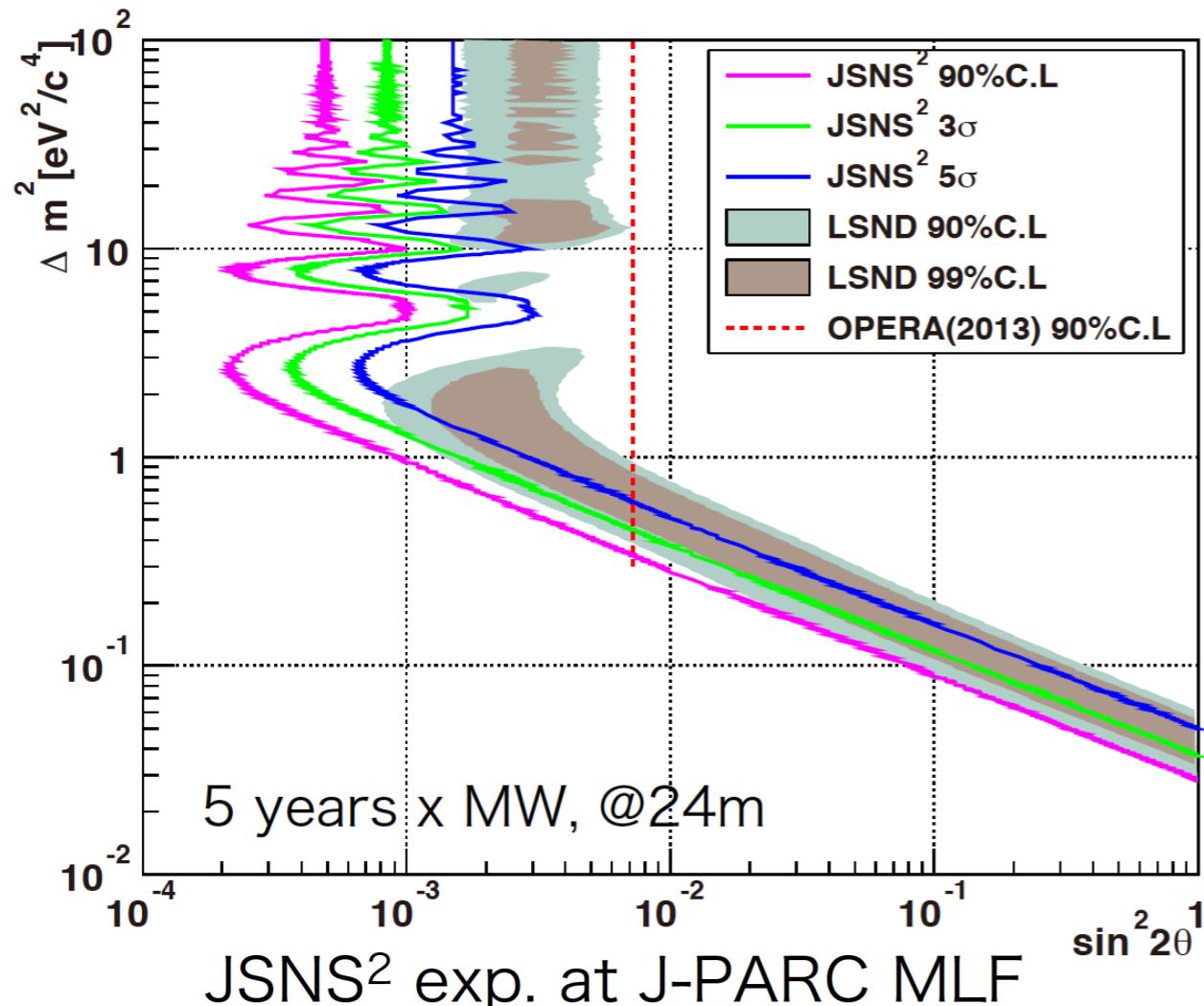
## ► Short-Baseline Program @ Fermilab - Kazuhiro Terao



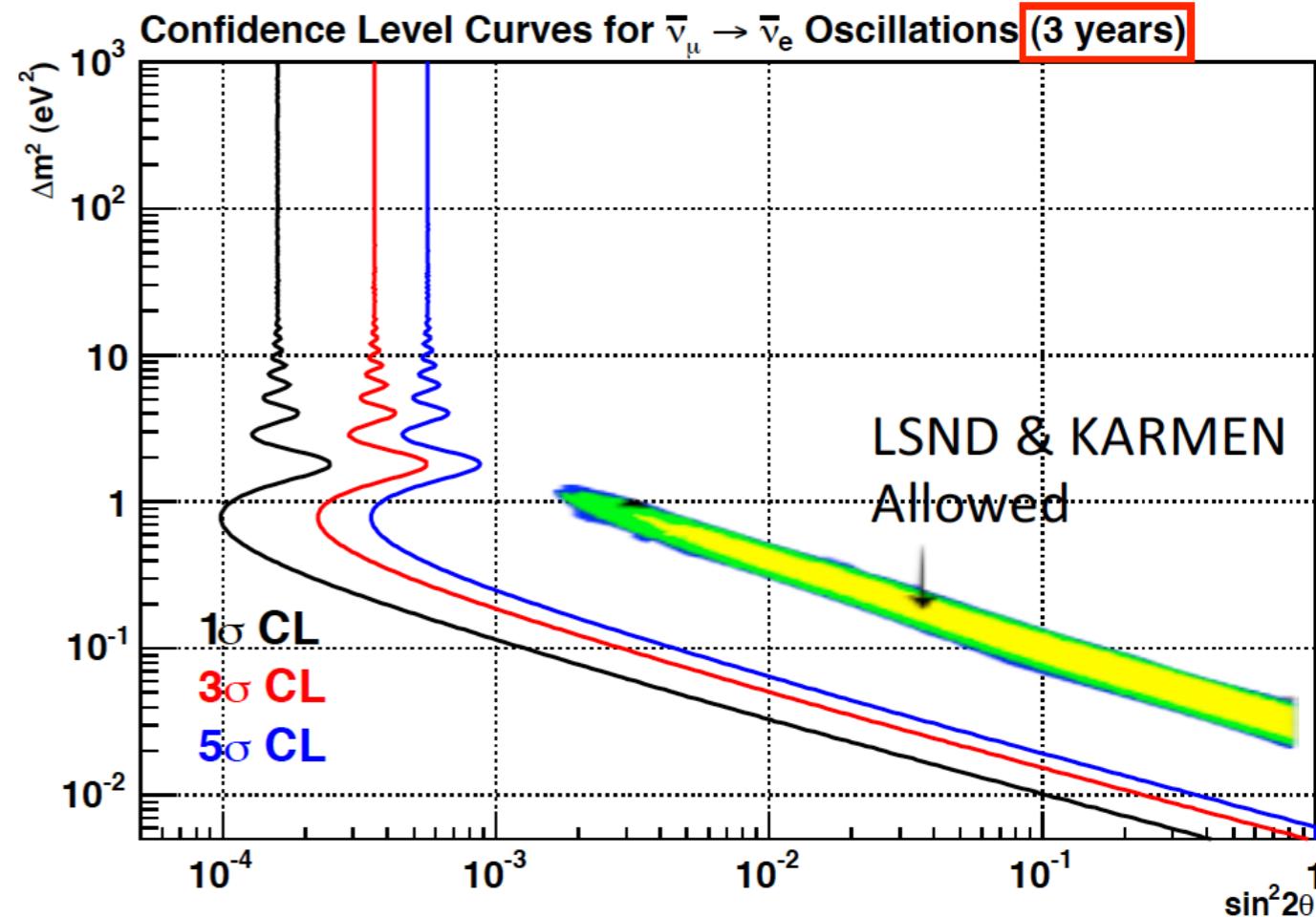
- MicroBooNE taking cosmic data!
- ICARUS undergoing refurbishment at CERN. Deployment at Fermilab in 2017
- Ongoing R&D program for SBND

# Sterile Neutrinos - Future

## ► Decay-at-rest Experiments - Eito Iwai



- beam energy: 3GeV
- beam power: 1MW designed, 500kW continuous
  - ▶ 24 m baseline
  - ▶ 2x50t LS detectors in phase 1 for \$4M
  - ▶ Phase 2 would use 1 kt detector

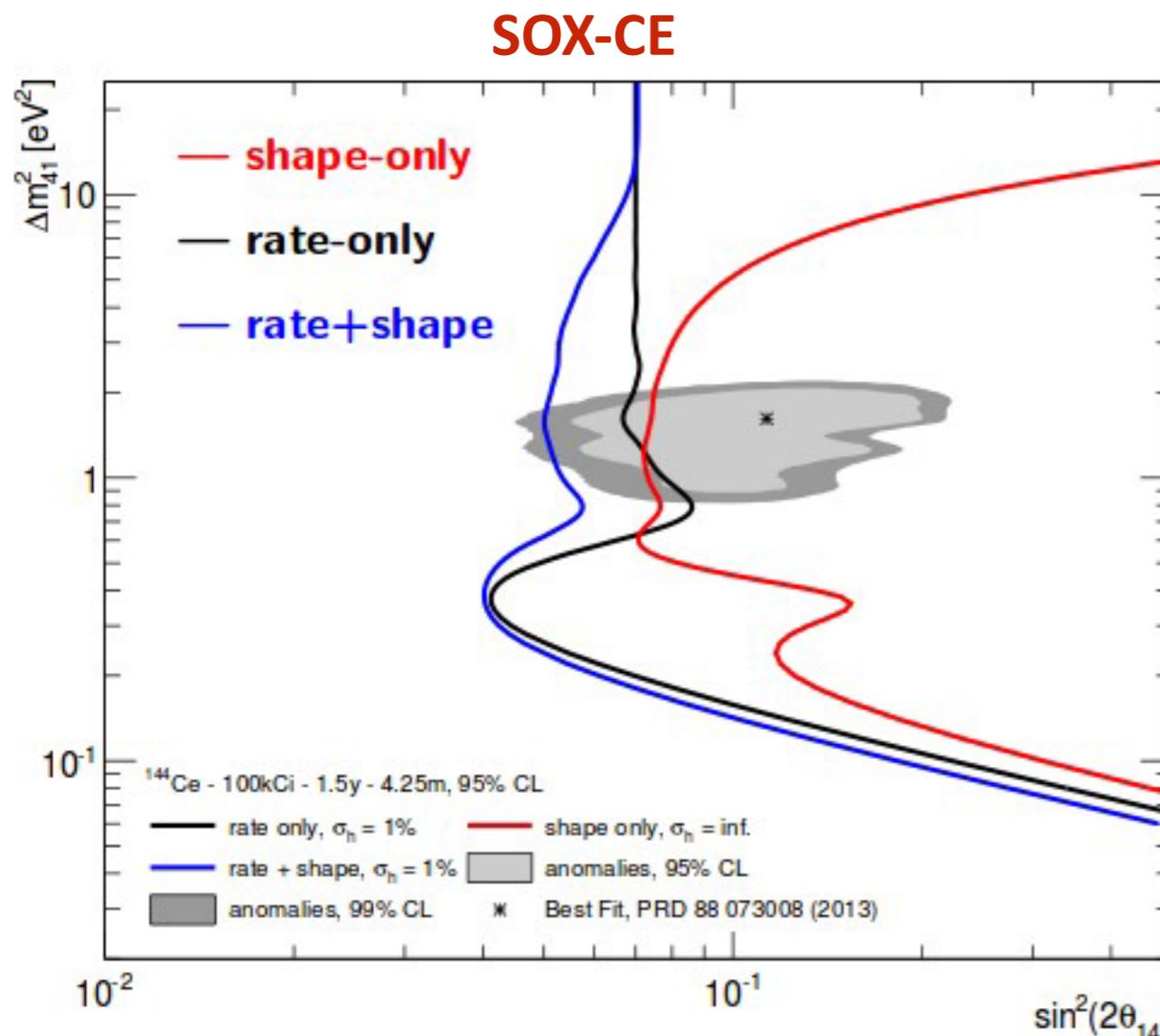


OscSNS at Oak Ridge National Laboratory

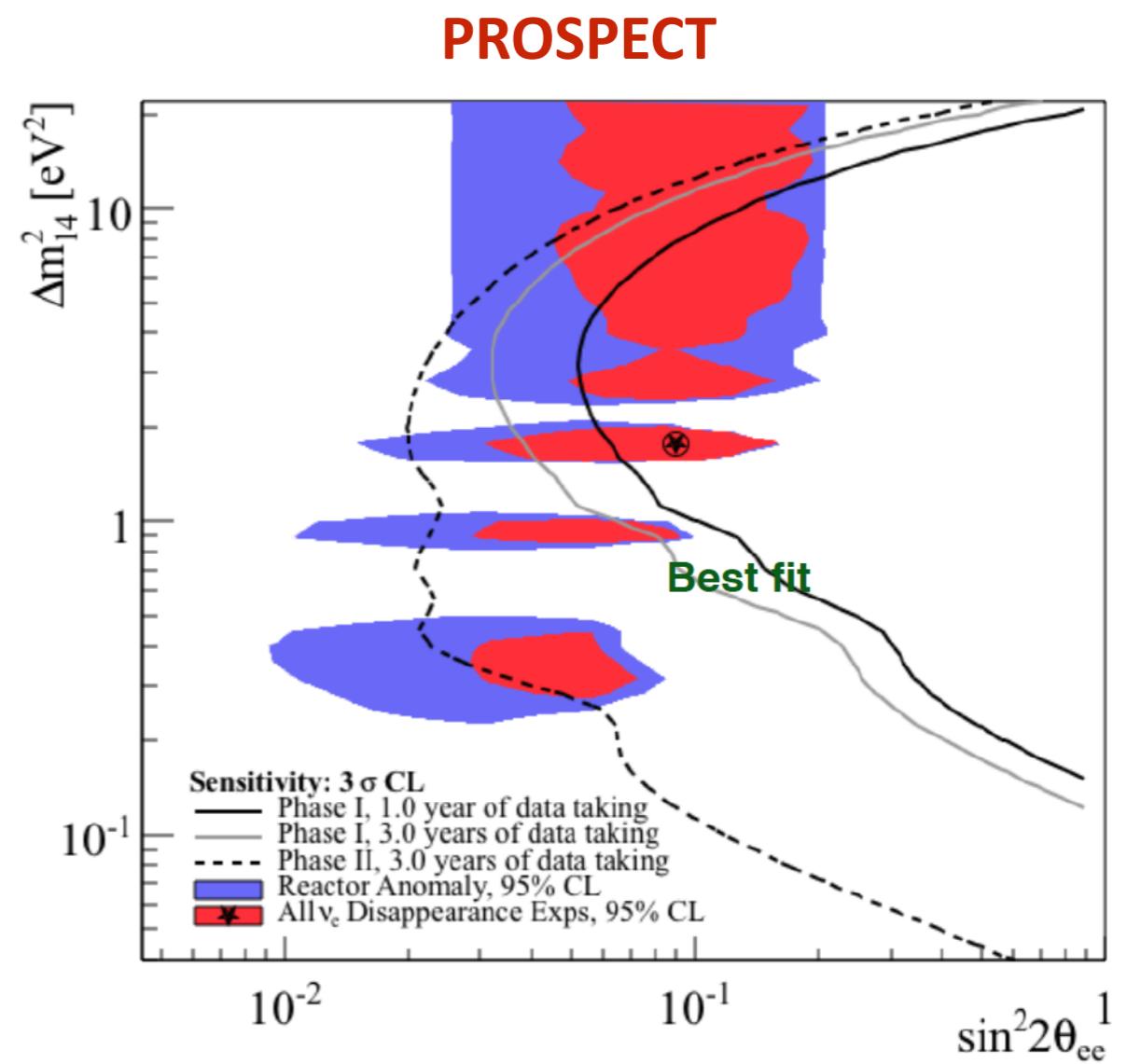
- proton beam: 1GeV, 1.4MW
  - ▶ 60 m baseline
  - ▶ 450t fiducial LS detector for \$22M
  - ▶ 3 years from ground breaking to operations

# Sterile Neutrinos - Future

## ► Source Experiments - Chiara Ghiano; SBL Reactors - David Martinez



- 3.7 Pbg  $^{144}\text{Ce}$  measured at Borexino
- Can exclude most of reactor/gallium anomalies at 95% CL from rate+oscillometry measurements
- Data taking starting in 2016, results in 2017



- Place 2.5 ton LS detector at Oak Ridge Research reactor
- Measure L-dependence of osc. prob.
- Ongoing R&D, seeking funding for Phase 1

# Answer to Q5

**Question from NuFact '14: What is the best strategy to fully probe the LSND anomaly? And the reactor/gallium anomaly?**

- ▶ Muon decay-at-rest experiments provide the most direct test of the LSND anomaly, but pion decay-in-flight SBN program at Fermilab has best near-term sensitivity to probe LSND region
- ▶ Long-baseline experiments provide complementary probes of disappearance and appearance due to sterile neutrino mixing
- ▶ Source experiments and SBL reactor experiments will both provide powerful tests of the reactor and gallium anomalies.

radioactive

reactor

Experiment	$\nu$ Source	$\nu$ Type	Channel	Host	Cost Category <sup>1</sup>
CeLAND [259]	$^{144}\text{Ce}-^{144}\text{Pr}$	$\bar{\nu}_e$	disapp.	Kamioka, Japan	small <sup>2</sup>
Daya Bay Source [260]	$^{144}\text{Ce}-^{144}\text{Pr}$	$\bar{\nu}_e$	disapp.	China	small
SOX [261]	$^{51}\text{Cr}$	$\nu_e$	disapp.	LNGS, Italy	small <sup>2</sup>
	$^{144}\text{Ce}-^{144}\text{Pr}$	$\bar{\nu}_e$	disapp.		
BEST [64]	$^{51}\text{Cr}$	$\nu_e$	disapp.	Russia	small
PROSPECT [262]	Reactor	$\bar{\nu}_e$	disapp.	US <sup>3</sup>	small
STEREO	Reactor	$\bar{\nu}_e$	disapp.	ILL, France	N/A <sup>4</sup>
DANSS [263]	Reactor	$\bar{\nu}_e$	disapp.	Russia	N/A <sup>4</sup>

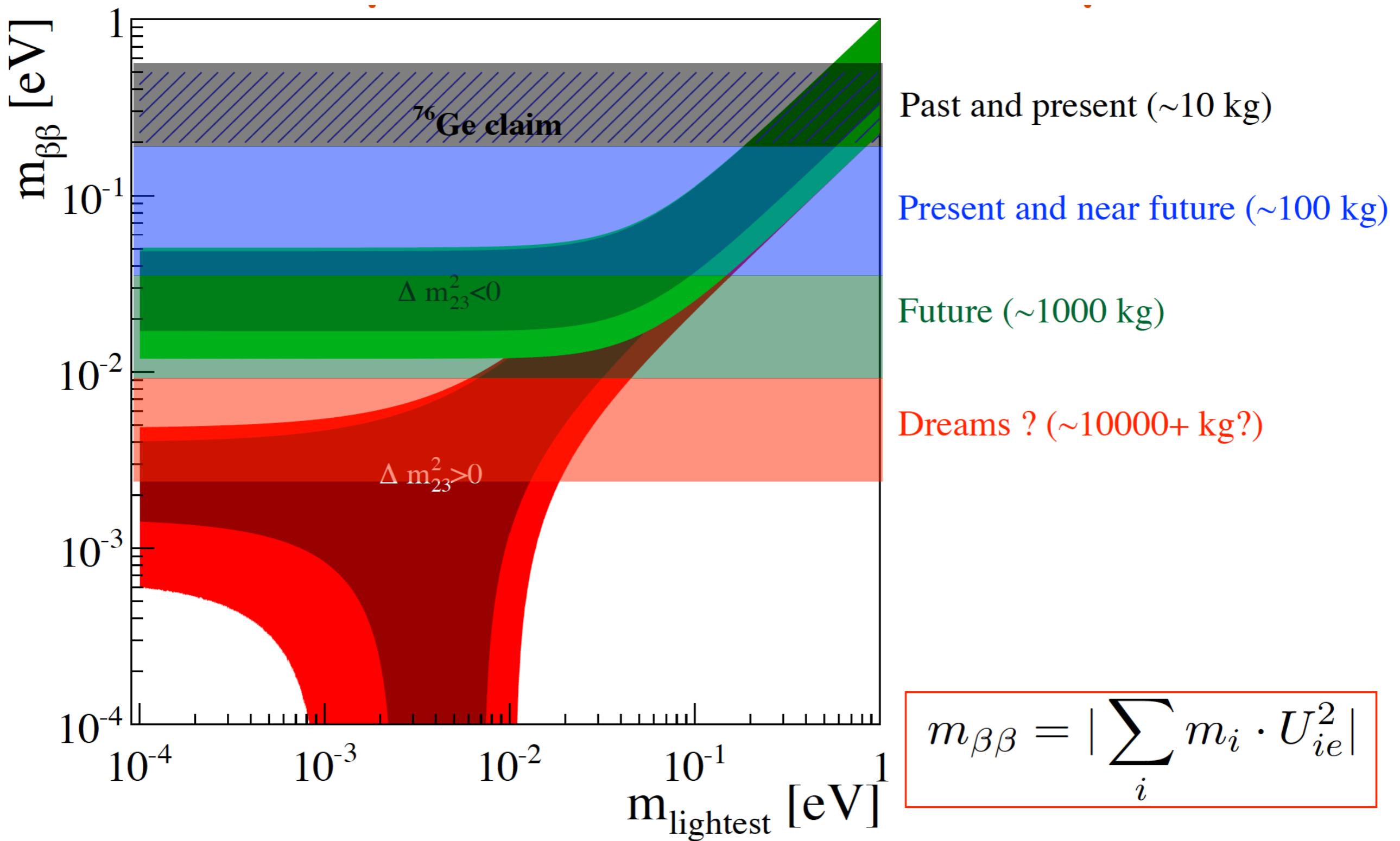
CSS 2013 (Snowmass) Neutrino Report, arXiv:1310.4340 (2013)

# Q6: Majorana mass scale

**Question from NuFact '14:** What can we say about the new Majorana mass scale implied by neutrino masses? What are the current bounds and how much will they improve over the next decade?

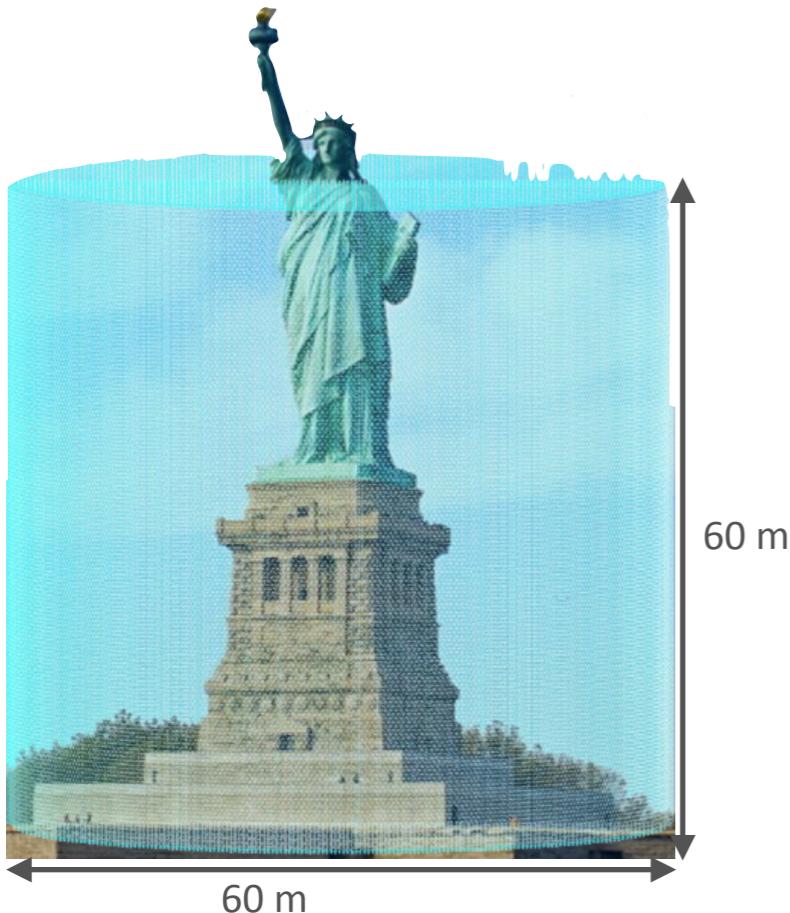
# Q6: Majorana mass scale

► Yury Kolomensky - Review of neutrinoless double beta decay



# Majorana mass scale

## ► THEIA - Gabriel Orebi Gann

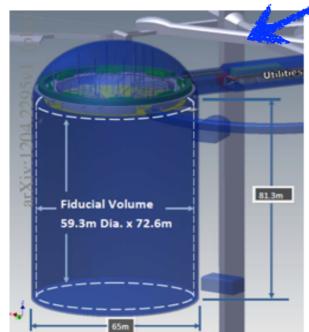
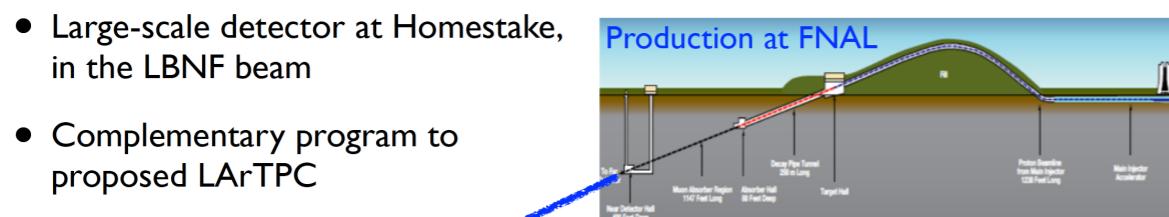


- Large-scale detector (50-100 kton)
- WbLS target
- Fast, high-efficiency photon detection with high coverage
- Deep u/ground (Pyhäsalmi, Homestake)
- Isotope loading (Gd, Te, Li...)
- Flexible! Target, loading, configuration

**→ Broad physics program!**

1. Neutrinoless double beta decay
2. Solar neutrinos (solar metallicity, luminosity)
3. Geo-neutrinos
4. Supernova burst neutrinos & DSNB
5. Source-based sterile searches
6. Nucleon decay
7. Long-baseline physics (mass hierarchy, CP violation)

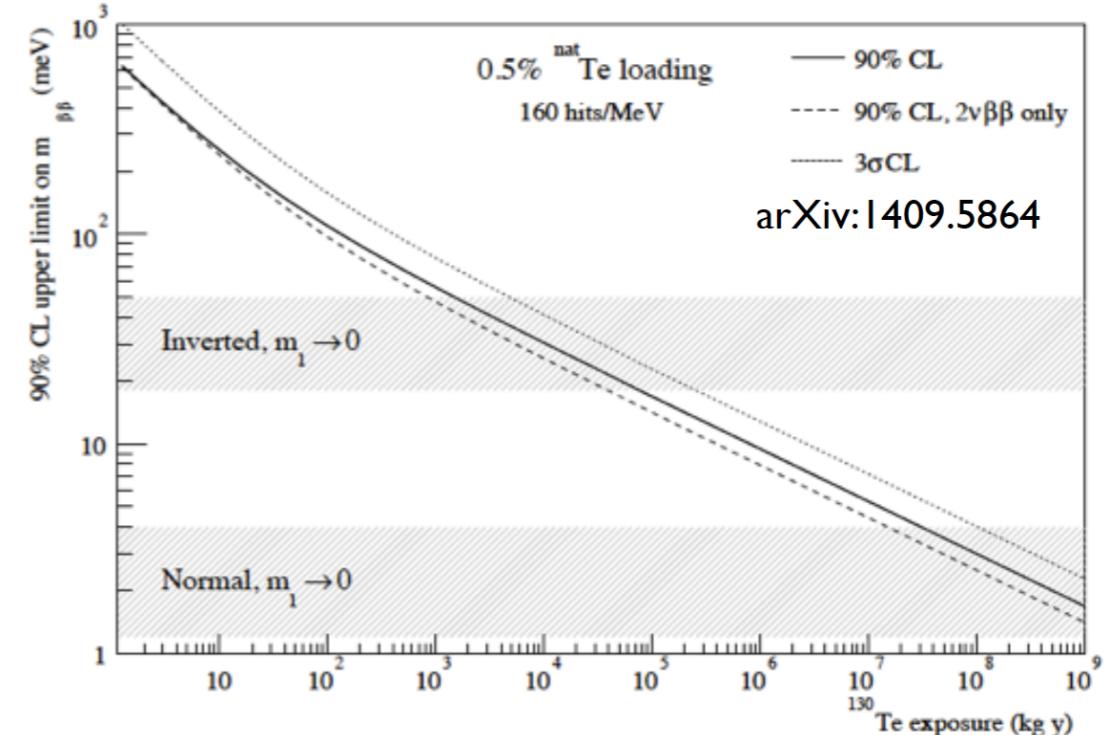
- Large-scale detector at Homestake, in the LBNF beam
- Complementary program to proposed LArTPC
- Build on WCD studies (arXiv:1204.2295)



Images from arXiv:1204.2295

- Ring-imaging of a water Cherenkov detector
- Particle ID from Cher/scint separation
- n and low-E hadron detection (low threshold)
  - ▶ reduce wrong-sign component (nu vs anti-nu)
  - ▶ reduce NC background by detecting  $\pi^0 \rightarrow \gamma\gamma$
- Large size → sensitivity to 2nd oscn max

MH sensitivity for 34kT WbLS alone  $> 4.8 \sigma$



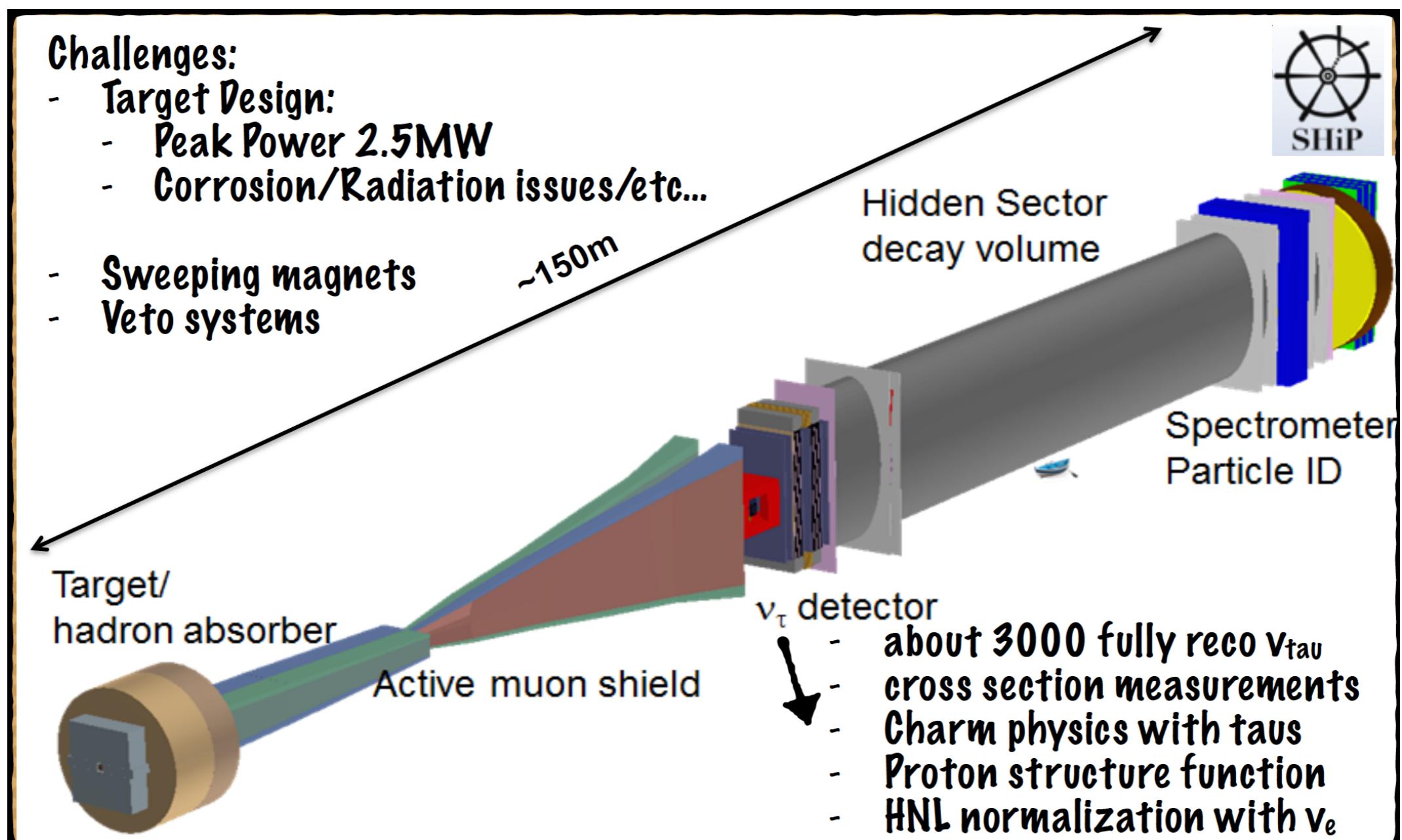
⇒ **3σ discovery for  $m_{\beta\beta} = 15$  meV in 10 yrs**

- Ongoing R&D for modeling of light production in WbLS, isotope loading, photon detectors (LAPPD), etc.

# Q7: Non-Oscillation Constraints

Question from NuFact '14: How can we explore synergies between neutrino oscillations and other experiments (absolute mass searches, cosmological constraints, CLFV) to constrain new physics?

## ► Heavy Neutrinos - Nicola Serra (ShiP Experiment@CERN)



# New Questions For Next NuFact

**Question:** What are the new developments and predictions from flavor models on neutrino oscillation parameters? What precision do we need to achieve to probe them? Which parameters (or combinations of them) are more powerful to test them?

**Question:** Are atmospheric neutrino and medium-baseline reactor measurements competitive with next generation long-baseline facilities in the determination of the mass hierarchy? And the octant of  $\theta_{23}$ ? How much complementarity is there between them?

**Question:** What is the best statistical metric to report sensitivity of experiments to mass hierarchy and CP violation?

**Question:** How important is it to account for correlations between systematic uncertainties in different experiments when producing global fits? How do we increase involvement of experiments in production of joint results?

# New Questions For Next NuFact

**Question:** What do we need to do to reach the required systematic error budget level? How much improvement in constraining flux and cross section uncertainties can we expect from dedicated hadron production, cross-section measurements, and near detector efforts? How can we improve knowledge of nue cross sections and improve energy reconstruction?

**Question:** Do the current bounds on new physics in the neutrino sector (NSI, non-unitarity, steriles...) allow for effects large enough to interfere with CPV searches? And with other parameters like mass hierarchy ? Which experimental setups can improve these bounds?

**Question:** Is the current landscape of proposed experiments able to conclusively probe the LSND anomaly? And the reactor/gallium anomaly?

# New Questions For Next NuFact

**Question:** *What can we say about the Majorana mass scale implied by neutrino masses? How much will the current bounds improve over the next decade? If normal hierarchy is confirmed, what new strategies are available to probe the Majorana/Dirac nature of the neutrino.*

**Question:** *How can we explore synergies between neutrino oscillations and other experiments (absolute mass searches, cosmological constraints, CLFV) to constrain new physics?*

**Question:** *What could an R&D program leading to a neutrino factory look like in the present landscape of LBL experiments? What physics could we probe with it and how complementary would it be to the ongoing LBL efforts?*



Many Thanks to all WG1 Speakers  
and a very lively Audience!